

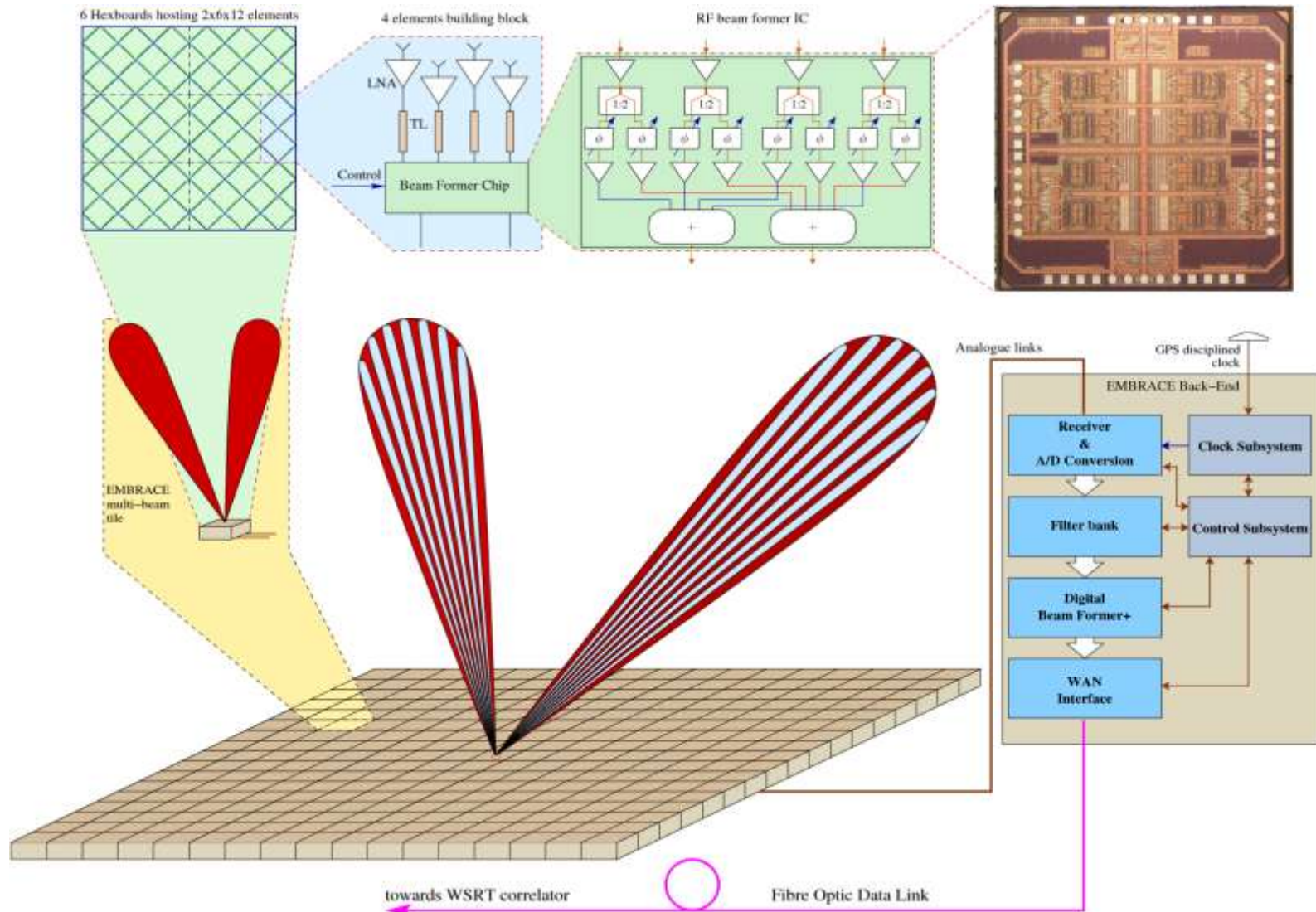


EMBRACE Status and Results

Dion Kant
on behalf of the EMBRACE team
ASTRON, OPAR, INAF, MPIfR



EMBRACE Architecture overview





EMBRACE Specifications

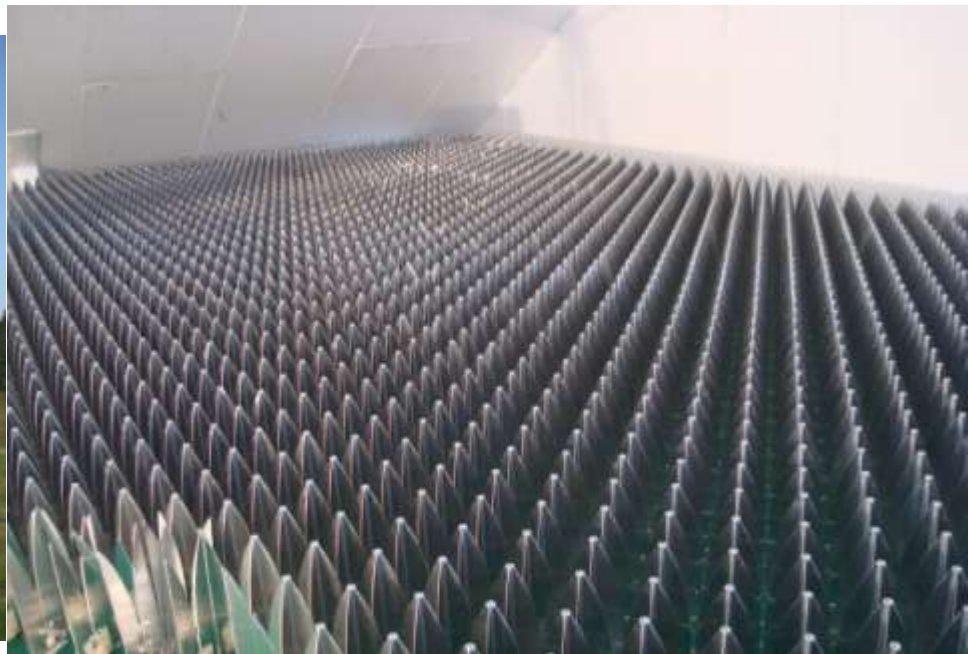
Frequency range	500 MHz - 1500 MHz.
Polarisation	Single polarisation*
Physical Collecting area	~160 m ² WSRT / 90 m ² Nançay
Aperture Efficiency	> 80%
Electronic Scan Range	>+-45 deg
T	100K @ 1GHz
Element phase control accuracy	3 bit (also time delays)
Instantaneous bandwidth	40 MHz Output bandwidth
Dynamic range A/D Converter	60dB (effective # of bits ~10)
Number of independent FoV (RF beams)	2
No of digital beams	8+

* Antenna design is dual polarized



EMBRACE

Station at Westerbork





Station at Nançay



image by SATorchinsky



Status Nançay Station



- ~ 62 m² area 56 Tiles
- 4032 Vivaldi elements (single linear polarisation)
- Currently 48 tiles cabled up end-to-end
- Grouped into tilesets of 4 tiles
- 64 tiles to be installed by end 2010
- Ultimately, 80 tiles in early 2011





Status Westerbork Station

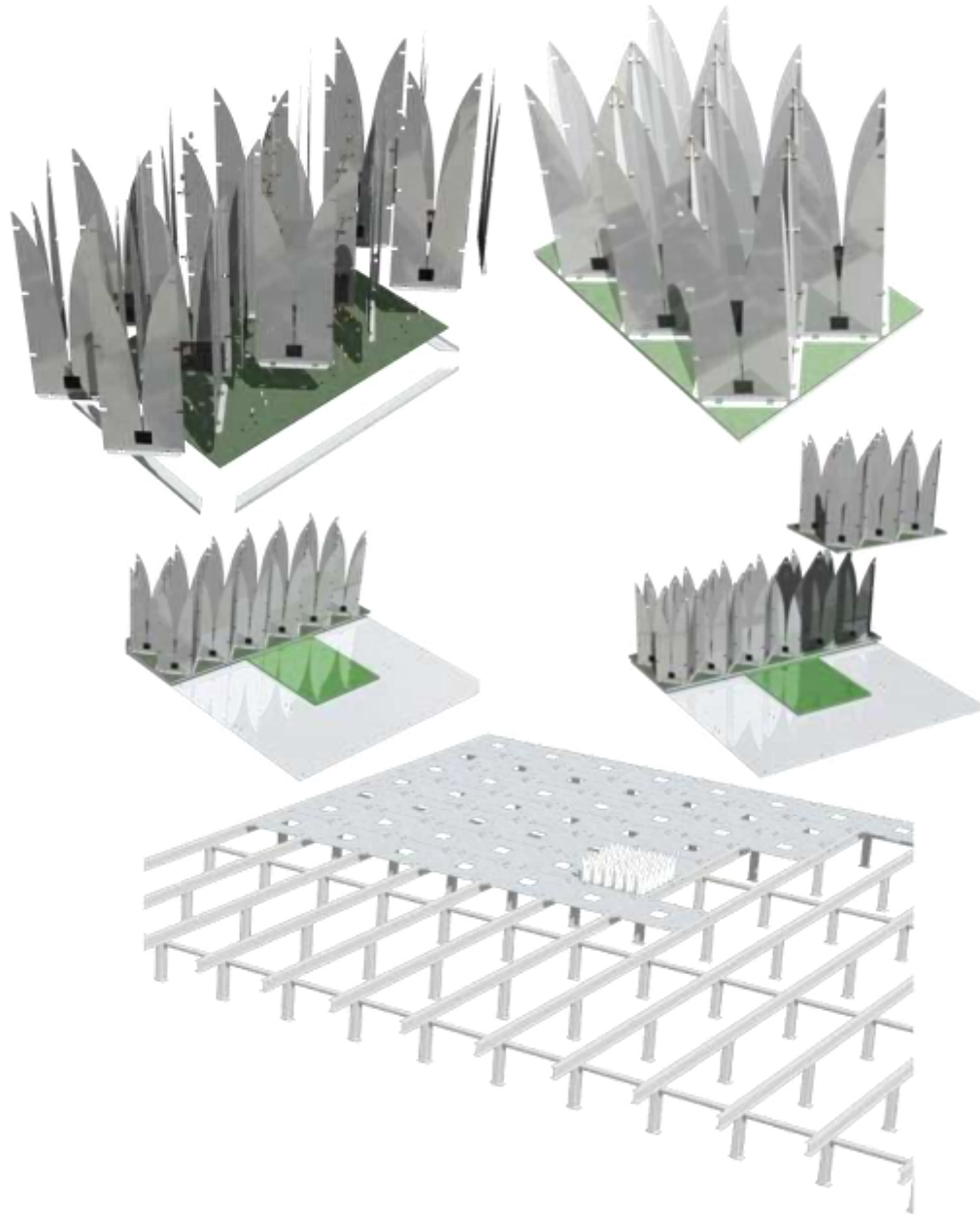


- 64 Tiles, ($\sim 70 \text{ m}^2$)
- 25 Tiles online in a 3x3 and 4x4 array configuration
- Merging control software with latest LOFAR devel
- Ultimately 144 Tiles in 2011

Current measurement objective:

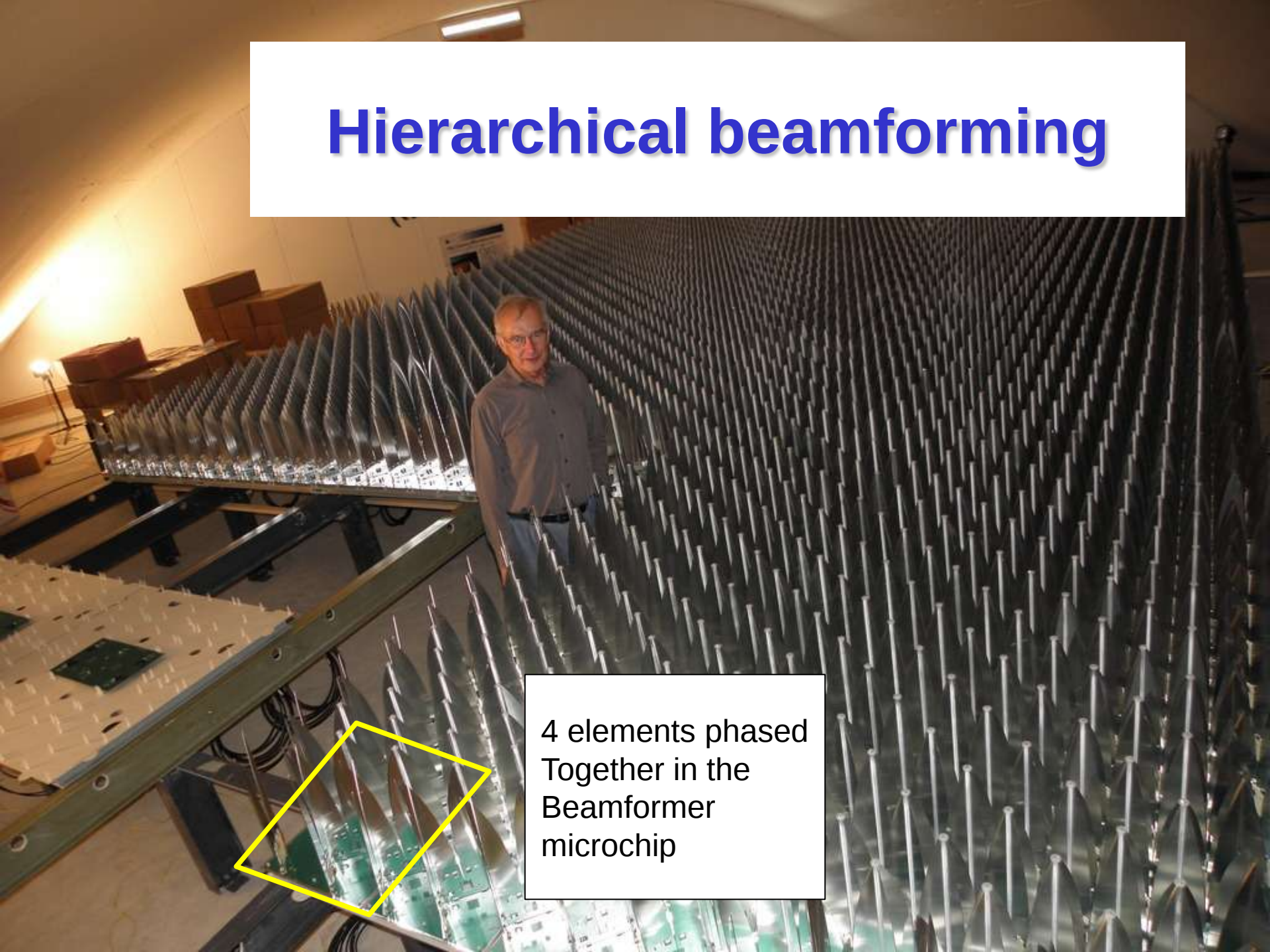
- Pulsar detection
- Simultaneous detection of pulsar and HI





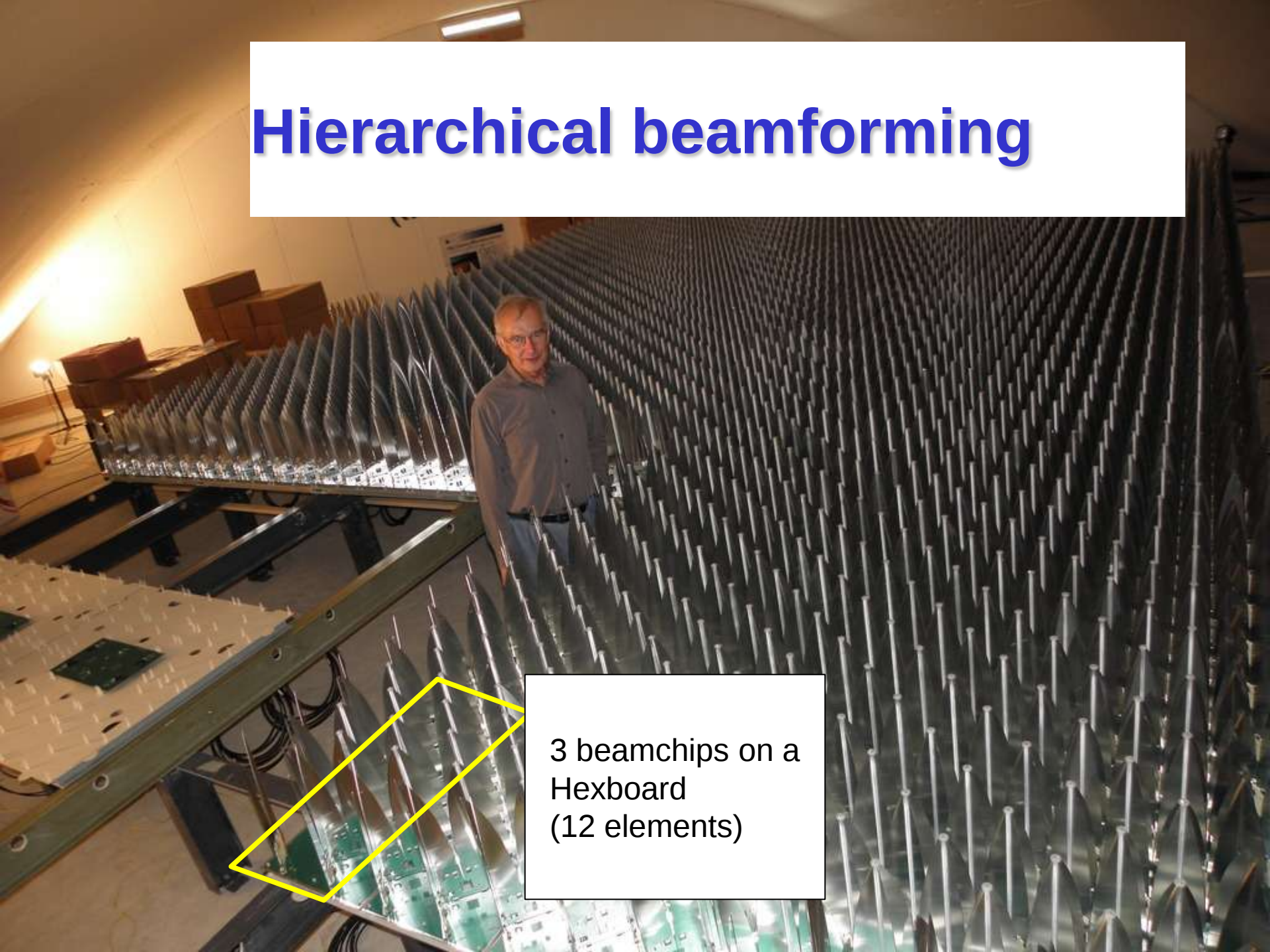
Mechanical break down of
EMBRACE array

Hierarchical beamforming



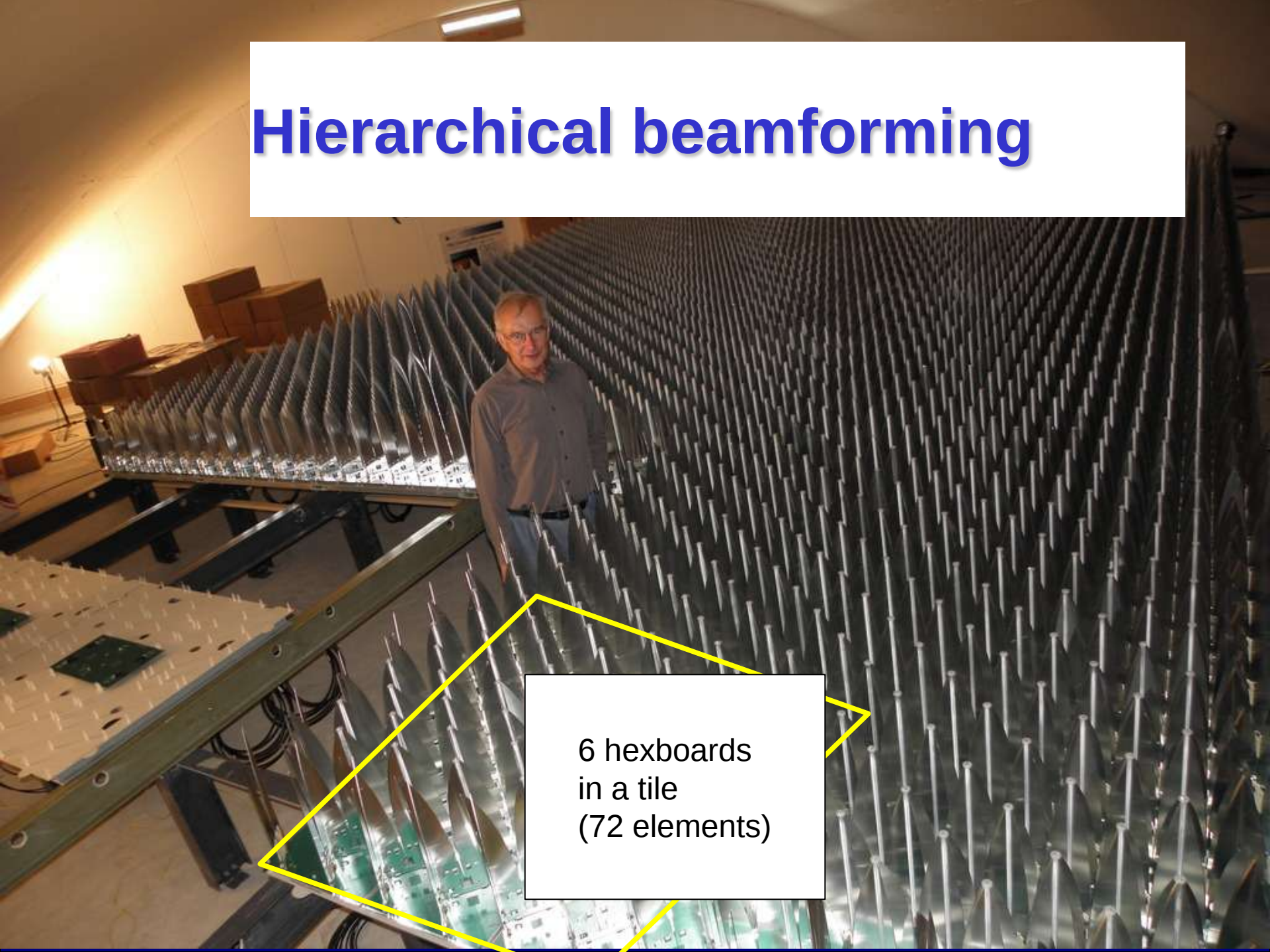
4 elements phased
Together in the
Beamformer
microchip

Hierarchical beamforming



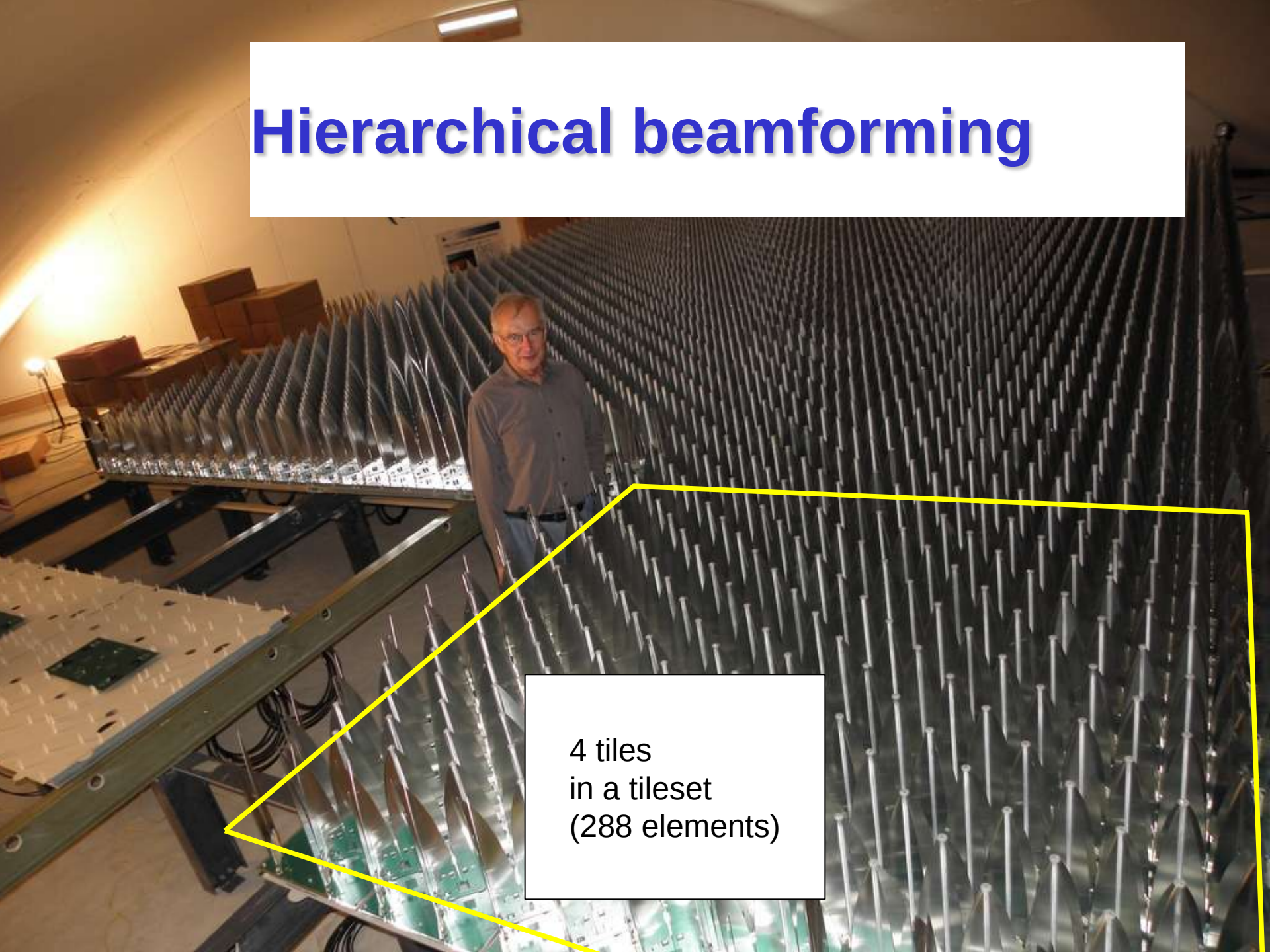
3 beamchips on a
Hexboard
(12 elements)

Hierarchical beamforming



6 hexboards
in a tile
(72 elements)

Hierarchical beamforming

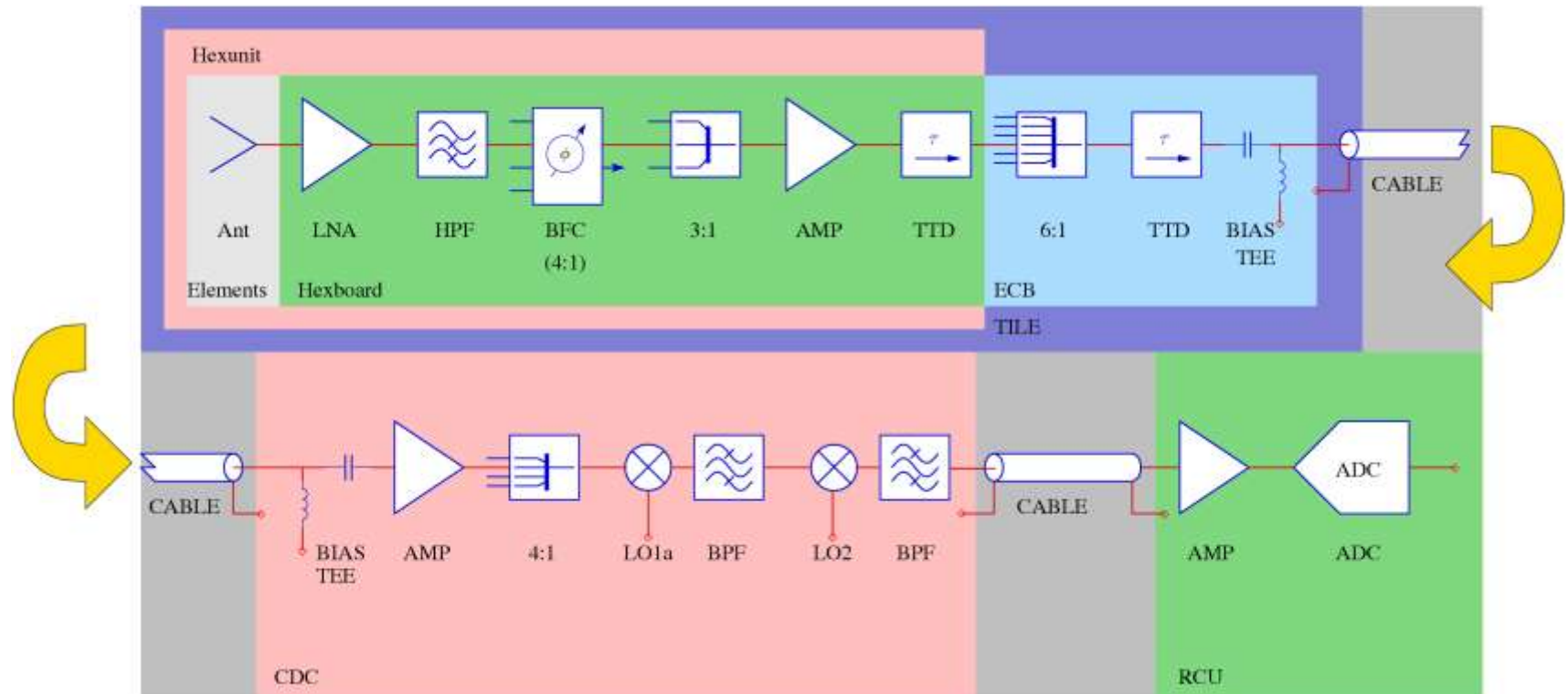


4 tiles
in a tileset
(288 elements)



EMBRACE

Signal flow block diagram



Control & Down Conversion

- CDC cards group 4 tiles together (tileset)
- RF and control/command multiplexed
- Currently debugging



Image by SA Torchinsky



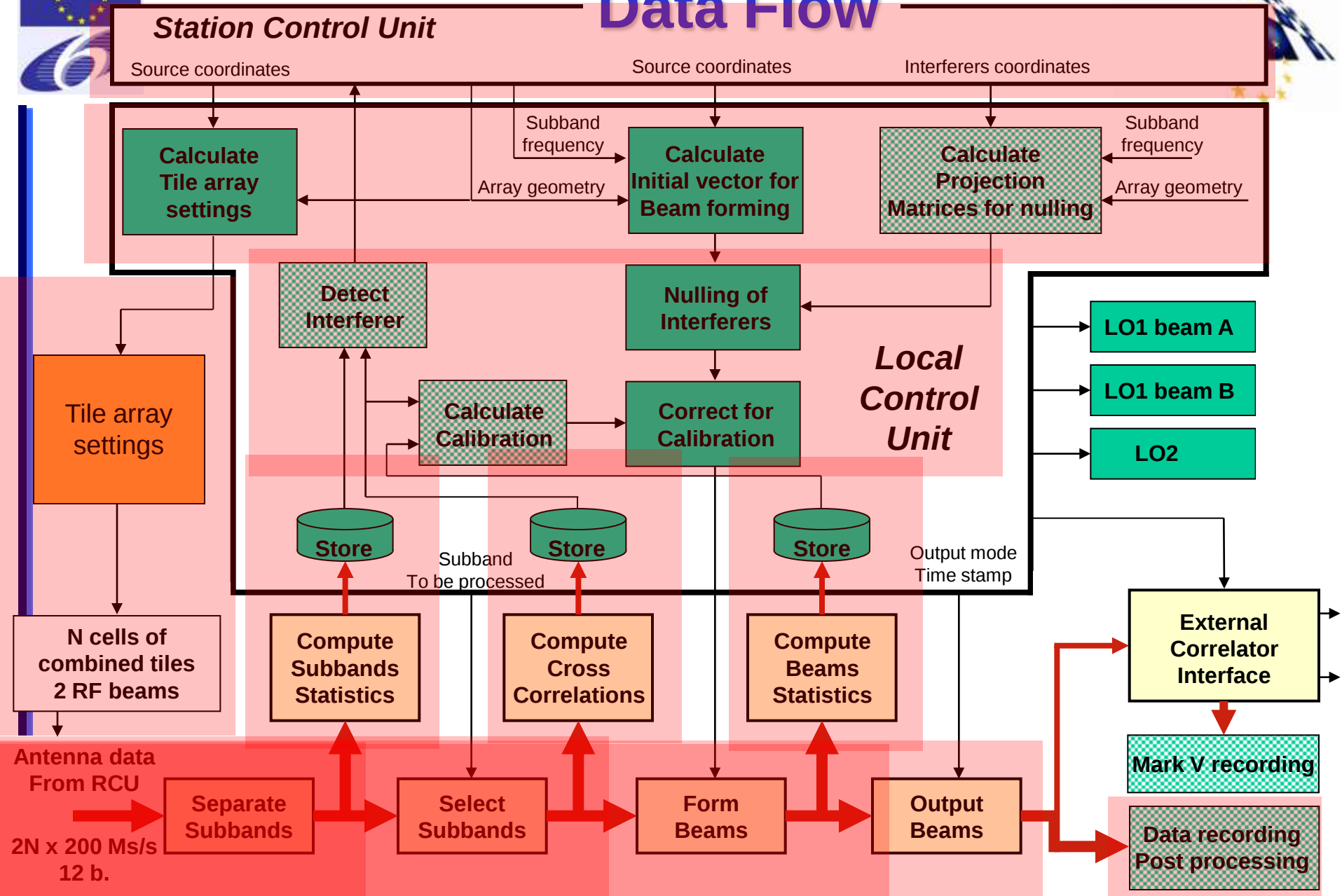


- Control and Down Conversion (CDC) Rack
- Back end Rack Layout (right)

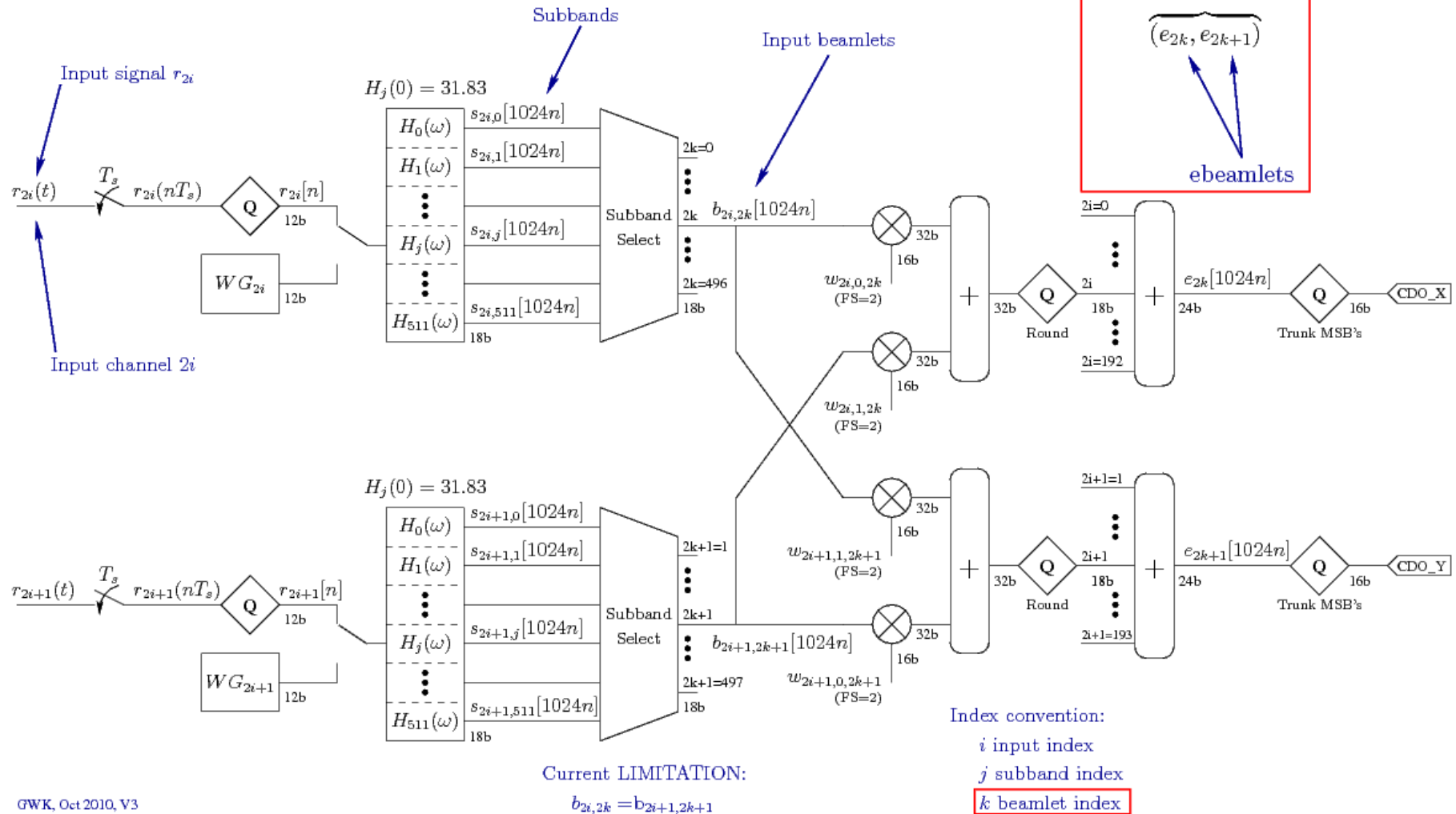




Data Flow

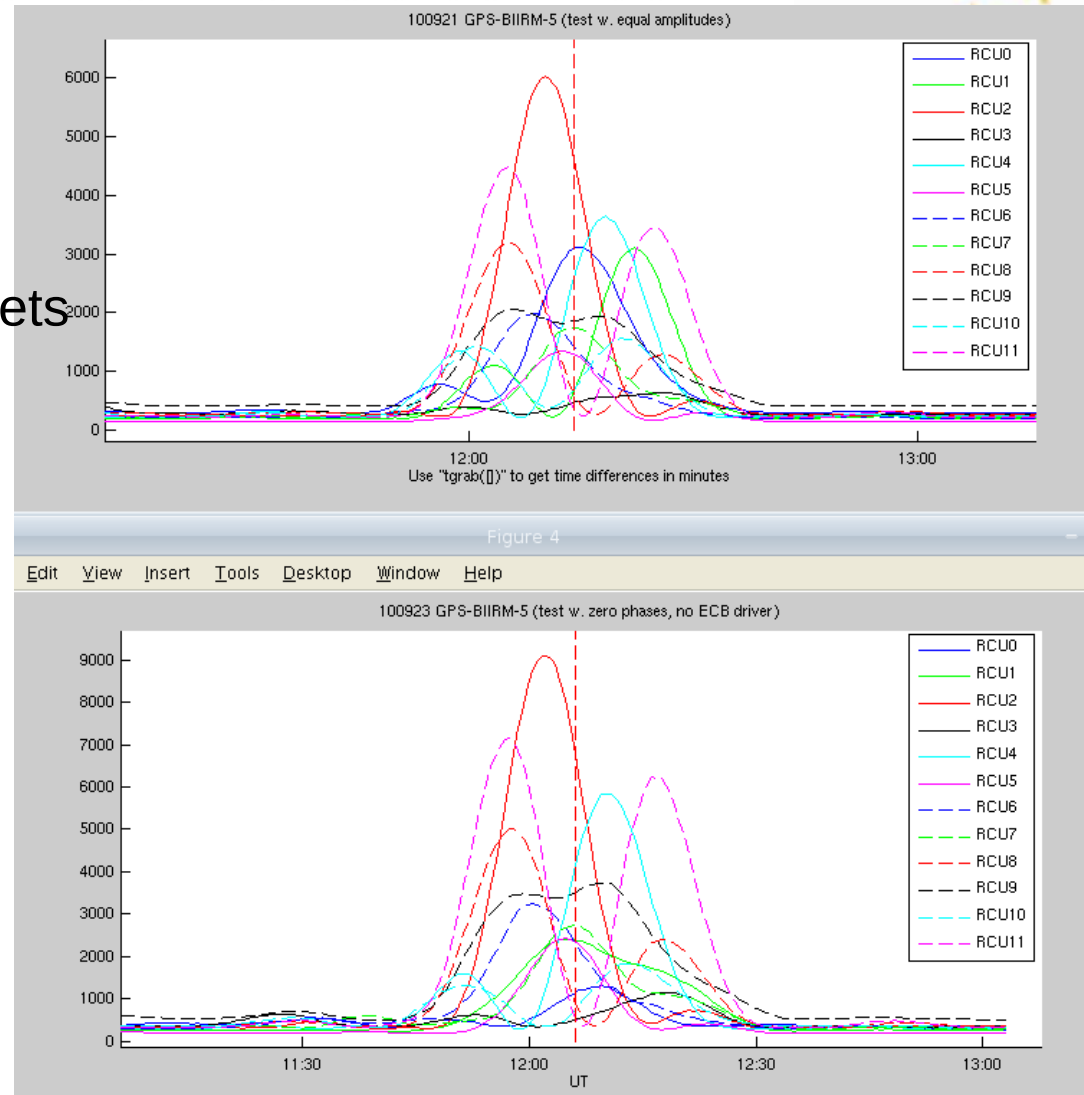


Name conventions



Drift scan of GPS satellite

- Strong carrier at 1227MHz
- Subband statistics
- Pointing offset between tilesets
 - problem with commanding phase settings
 - CDC + centre board driver





Early highlights: solar fringes

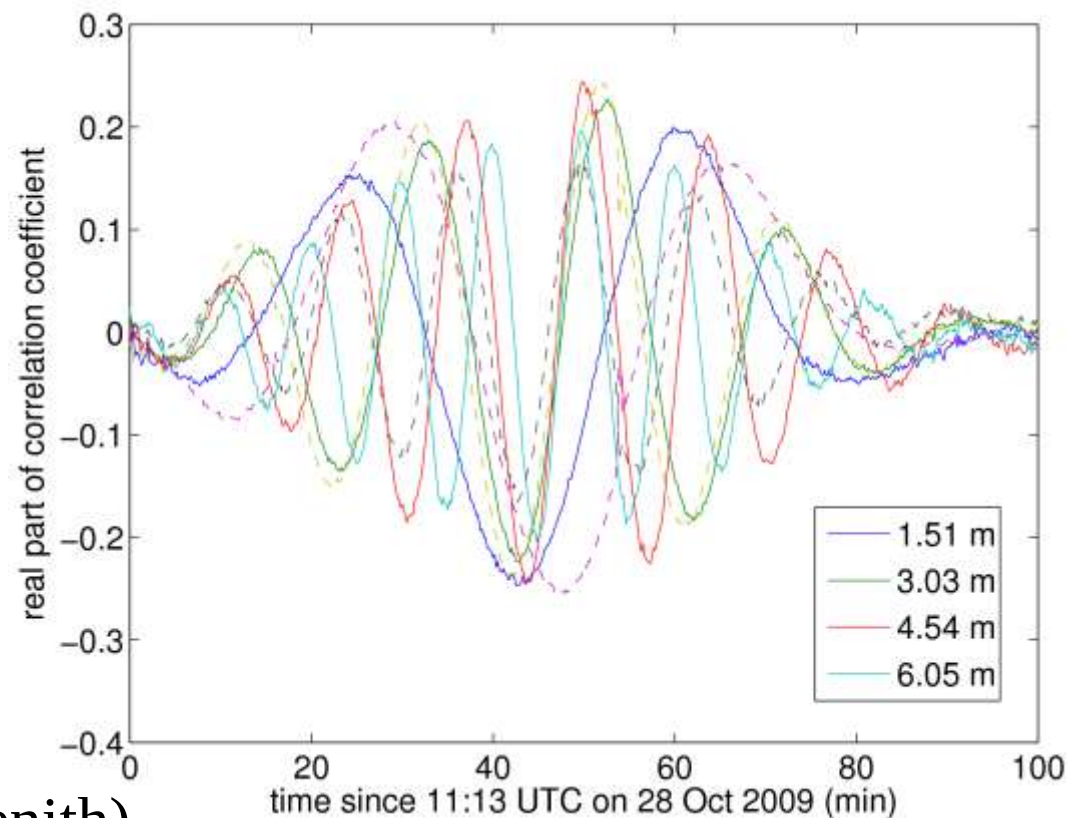


Interferometric measurement with 10 tiles

- two E-W ULAs
- frequency: 1179 MHz
- integration: 10 s
- bandwidth: 195 kHz

Initial performance est.

- $A/T = 4.7 \cdot 10^{-3}$ (to sun)
- $T_{\text{sys}} = 103 - 117$ K (to zenith)

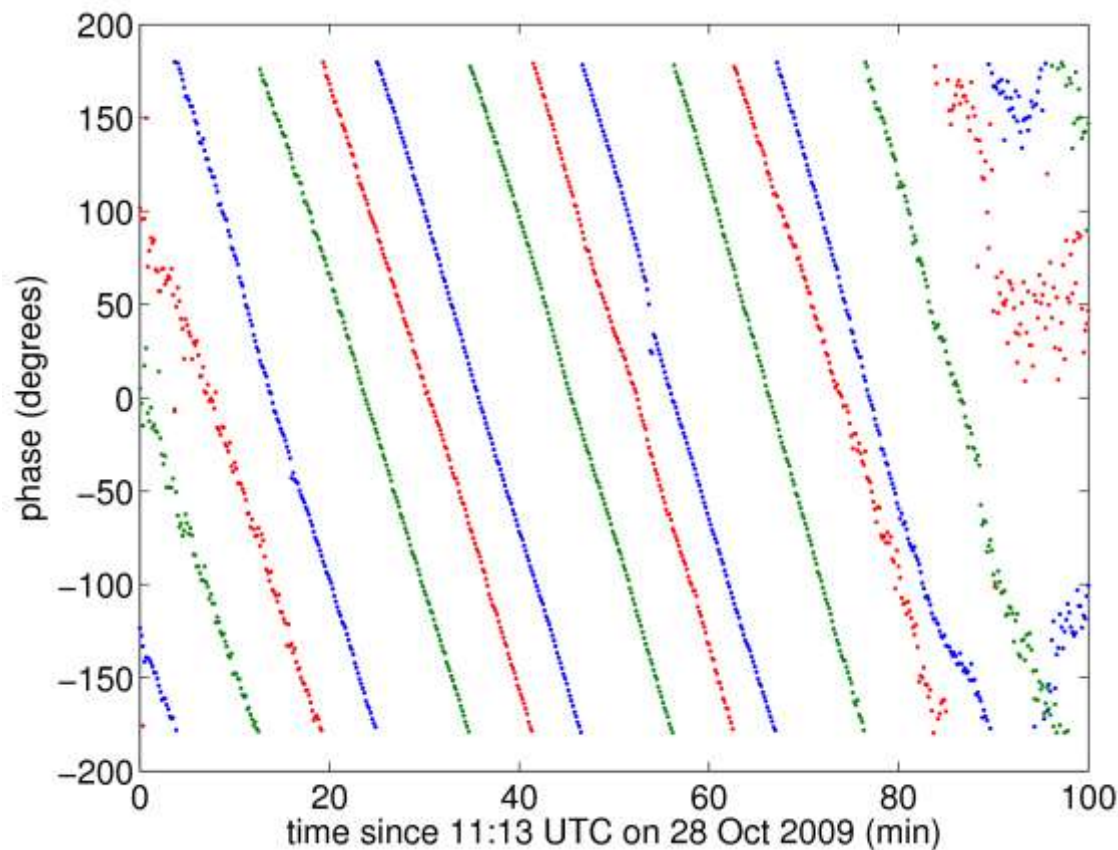




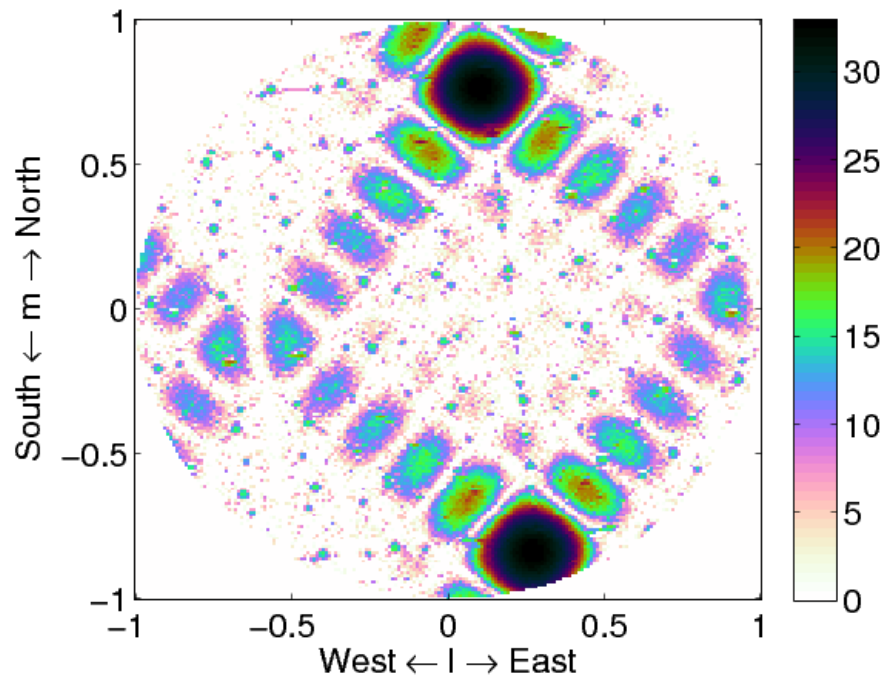
Early highlights: instrument quality



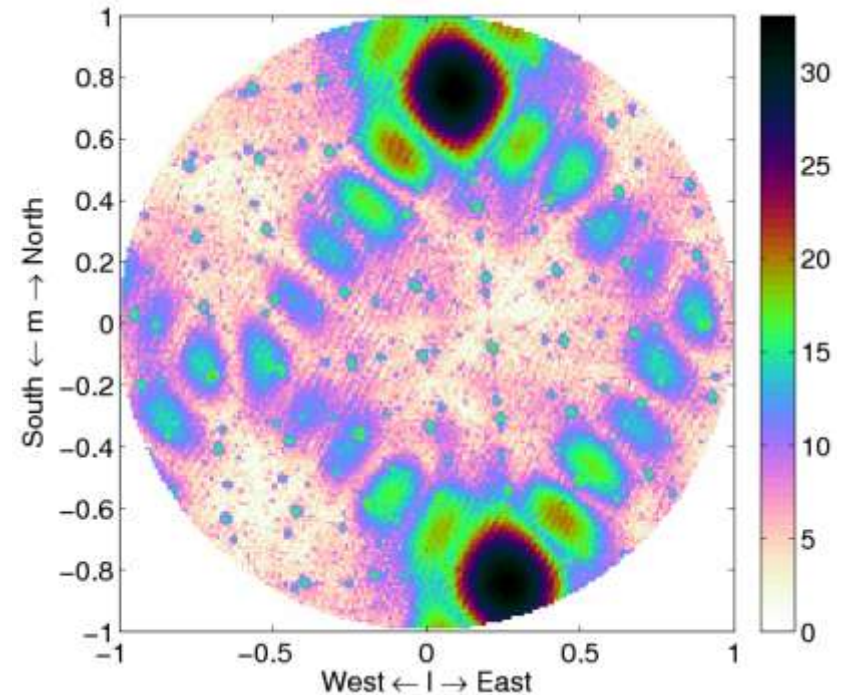
- Redundant baselines produce redundant visibilities



Scanned Tile Beam pattern measured on Afristar satellite



Model



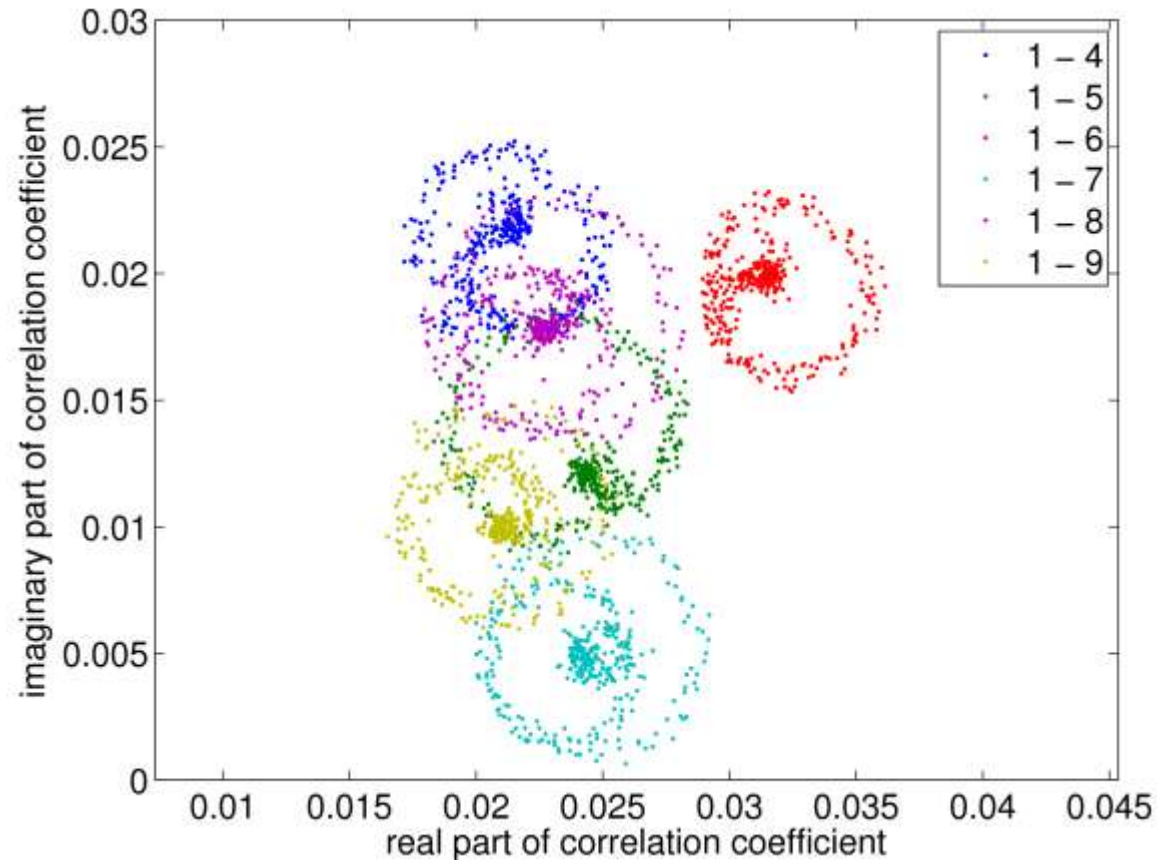
Measurement



Fringes of Cas A!

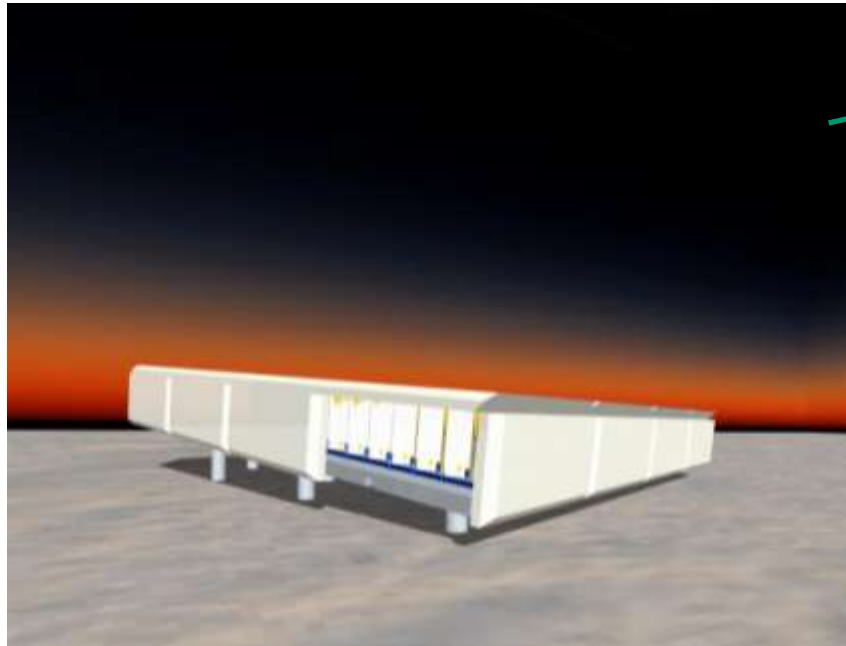


- **Experimental settings**
- 3x3 array
- 1254 MHz
- 30 s integration
- 195 kHz bandwidth
- **Initial conclusions**
- Confirms A/T
- Correlation offsets



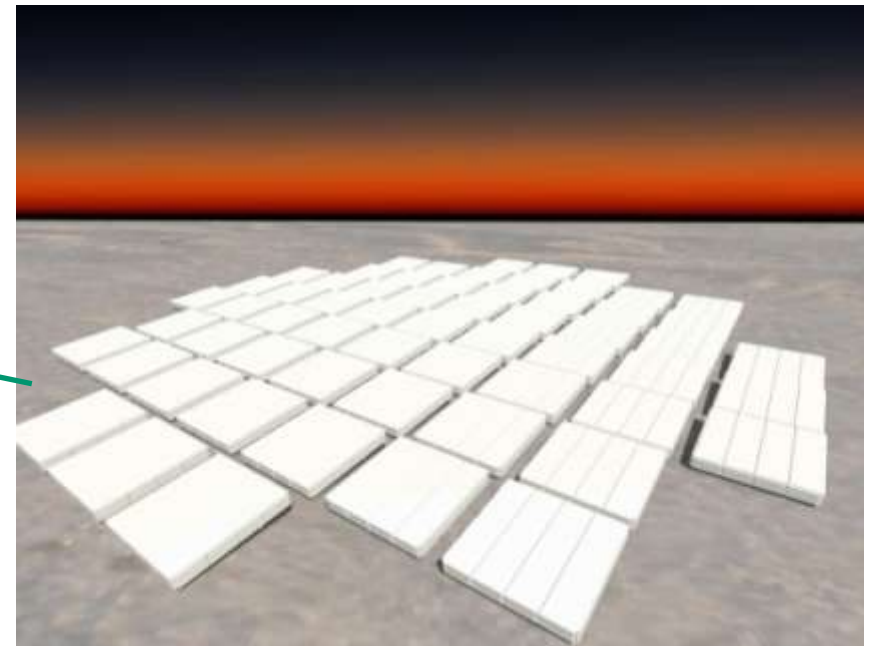


AAVS Design



- XX “tiles” in a substation.
- Every substation slightly tilted.
- Substation closed by EPS wall on top.
- Passive cooling by ground connection through poles.

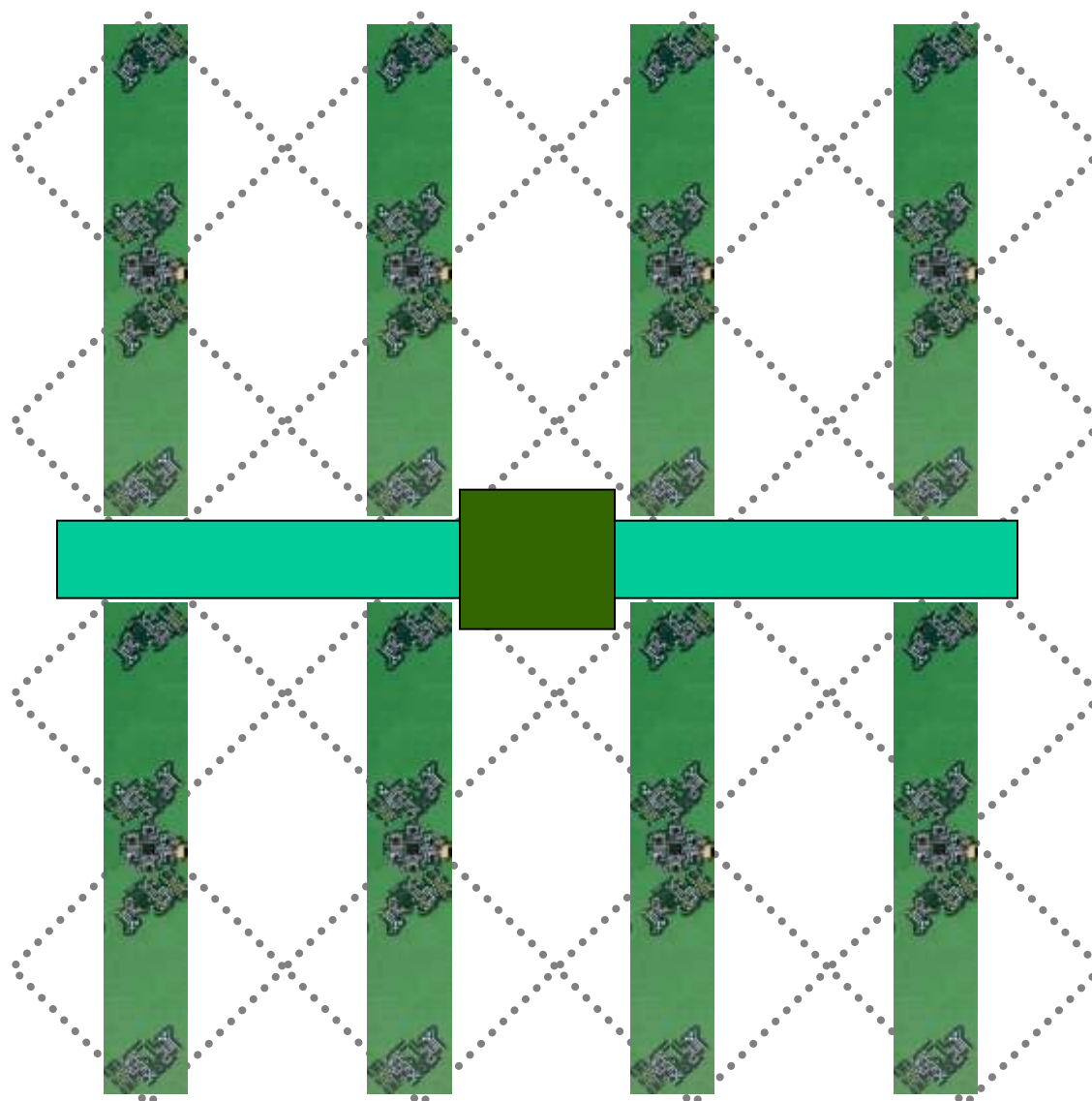
- SKA station formed by multiple substations.
- Walkable spaces between substations.
- 56m diameter



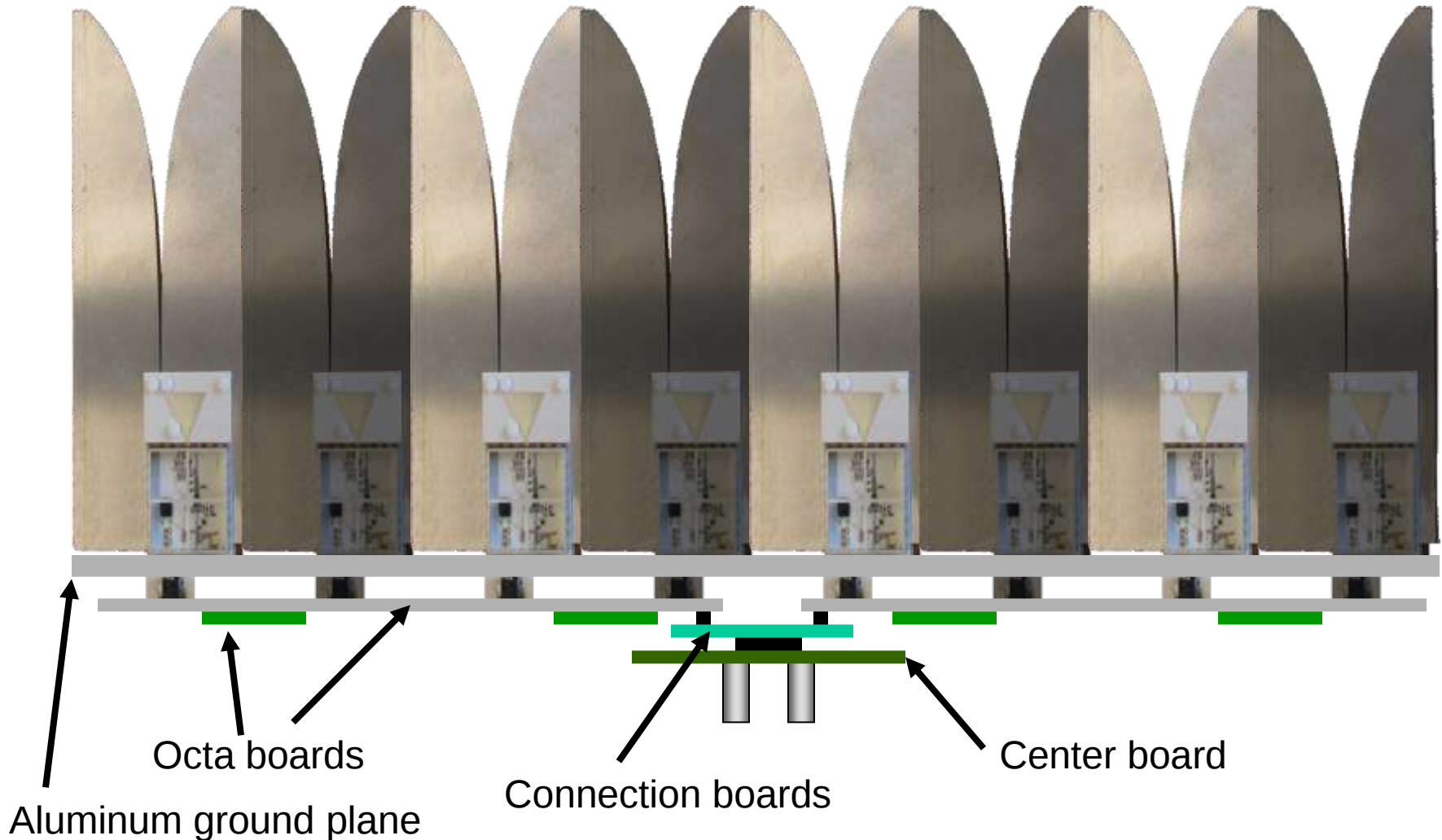
288.000 antenna elements



AAVS Octa concept; bottom view

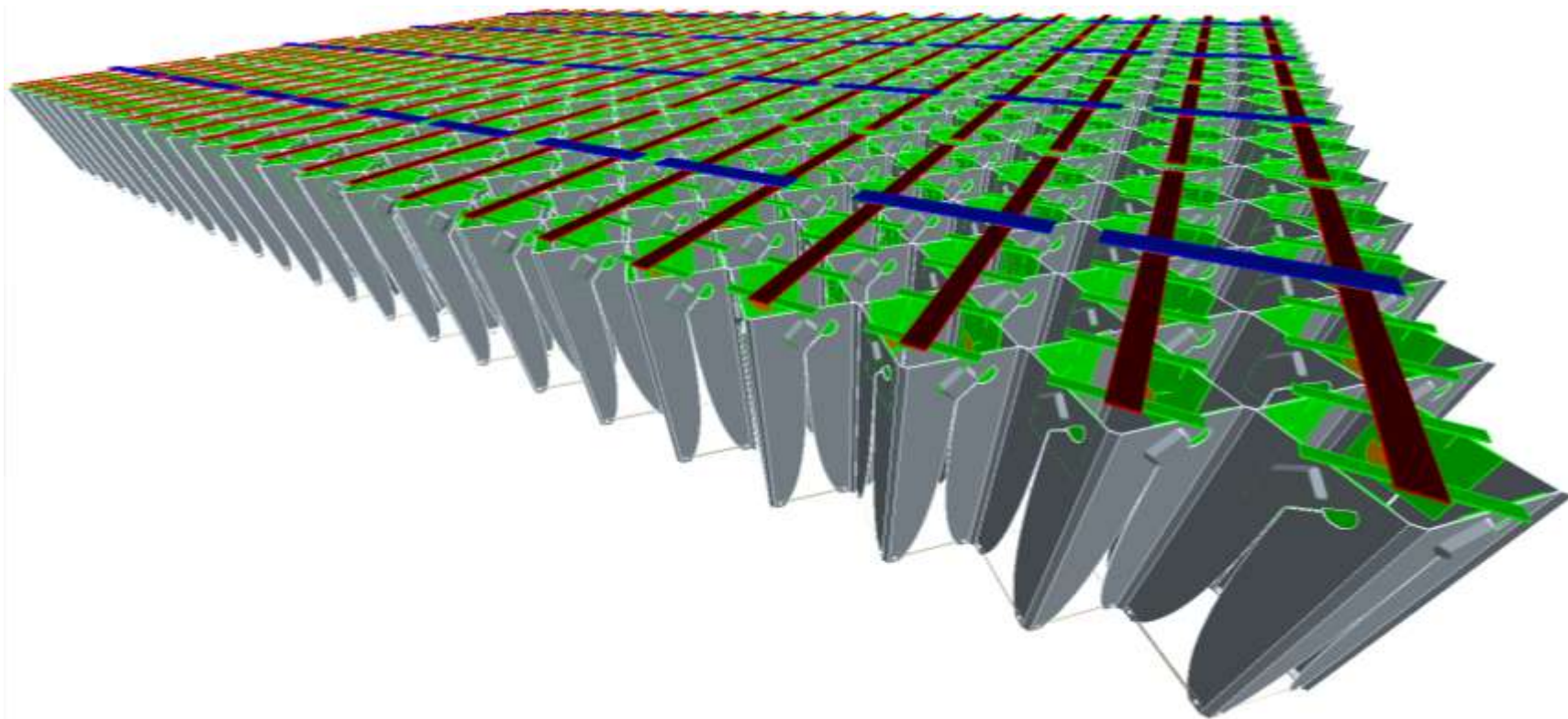


AAVS Octa concept; side view





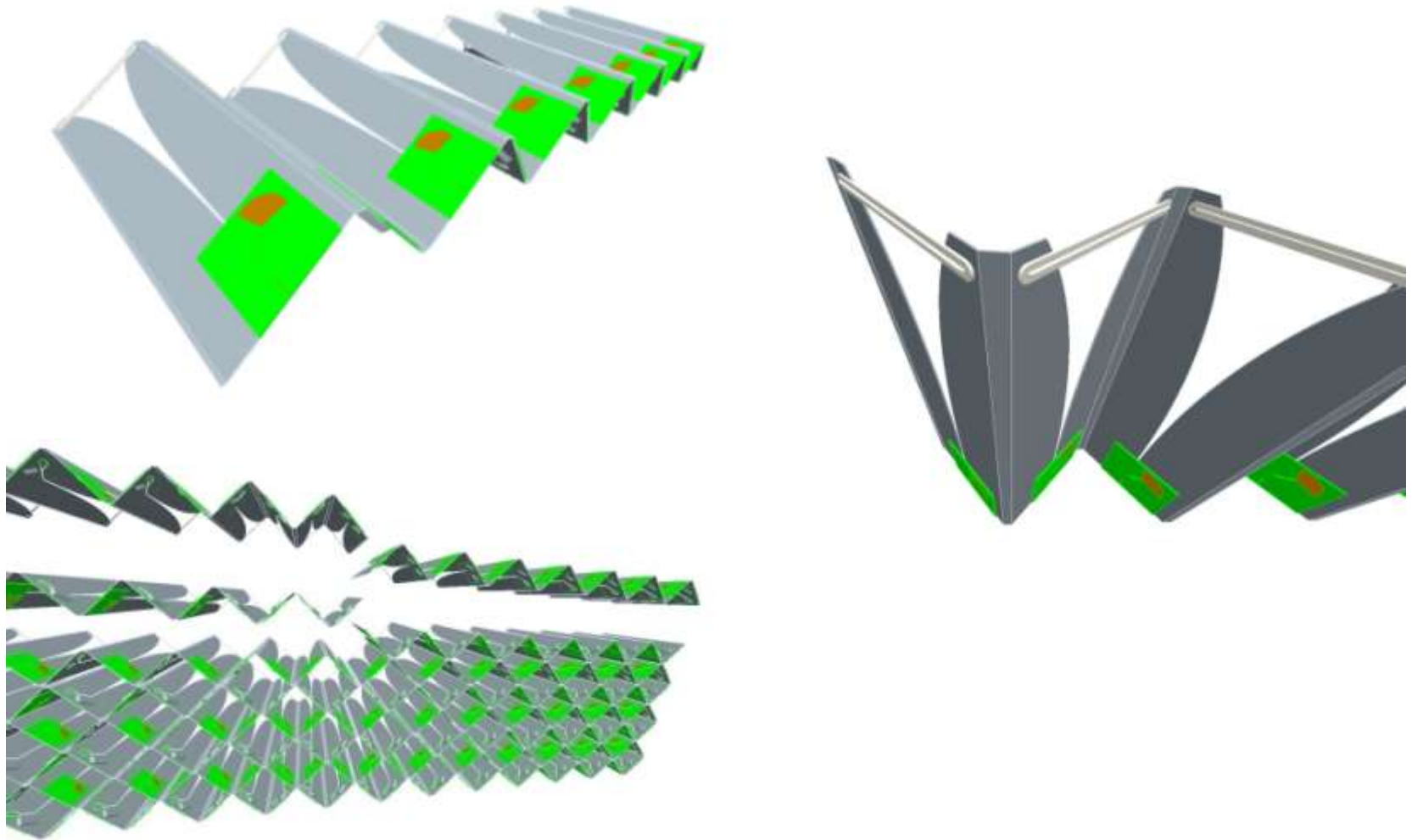
AAVS Octa Concept; 3D bottom view





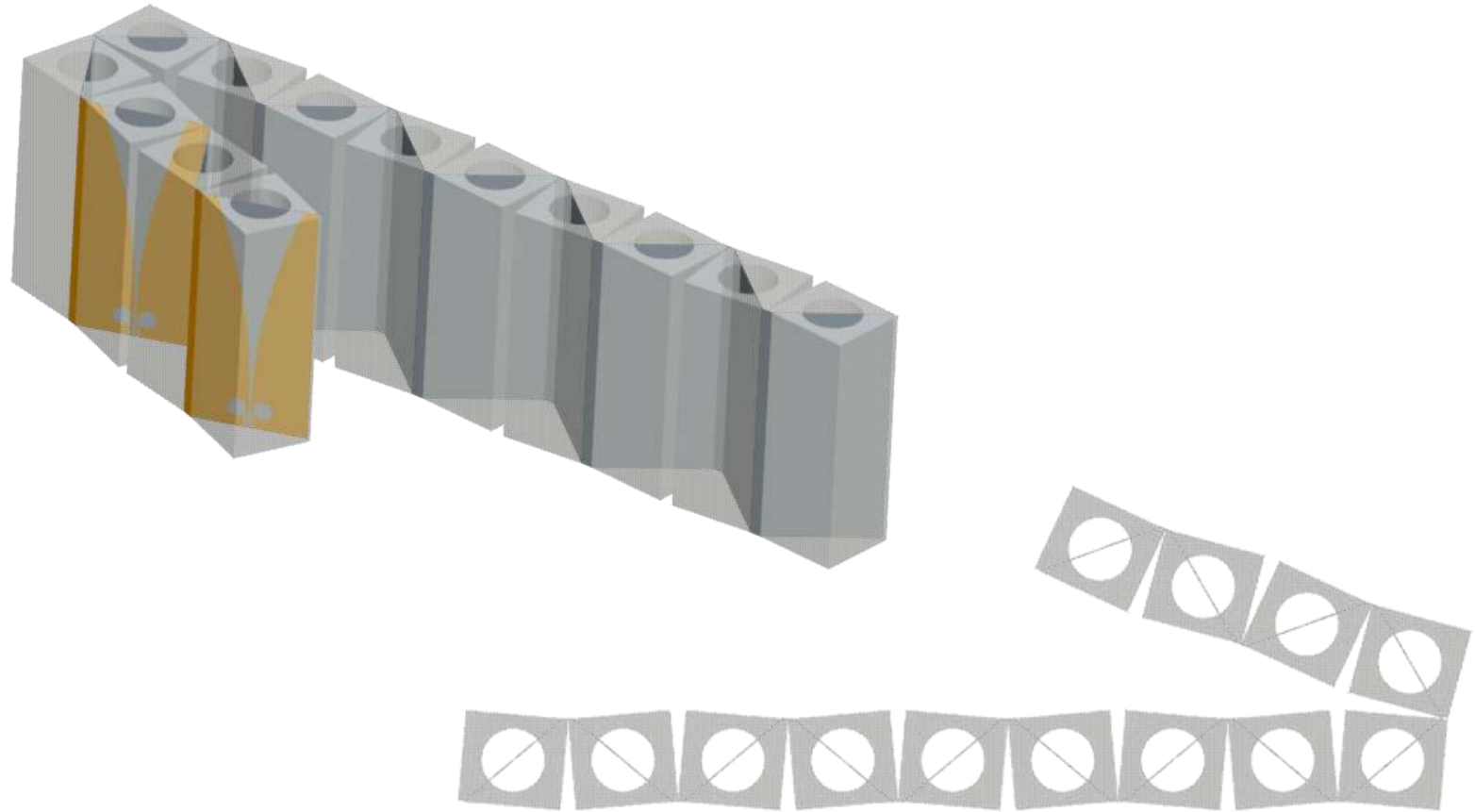
AAVS antenna concepts

Aluminium plate based





AAVS antenna concepts Foil based





Conclusion



- Both EMBRACE Stations are finalised
- First results are available and much more to come
- EMBRACE evaluation is continued in the framework of AAVP
- EMBRACE will be used as a basis for AAVS development



Questions?

ASTRON



Unité Scientifique de Nançay

