



UNIVERSITÉ
DE GENÈVE

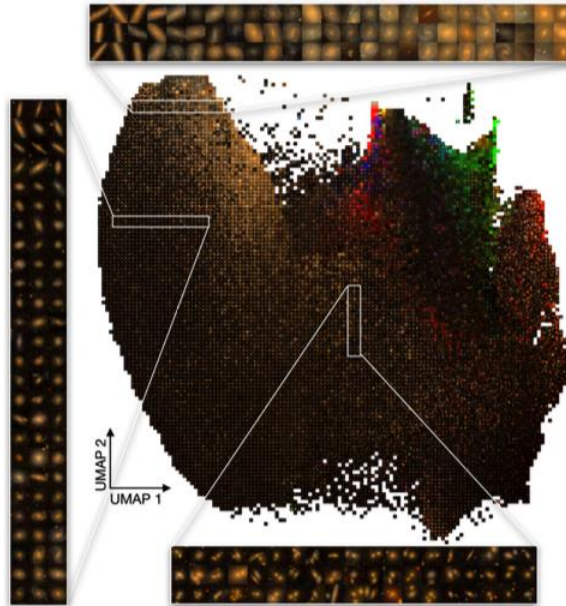
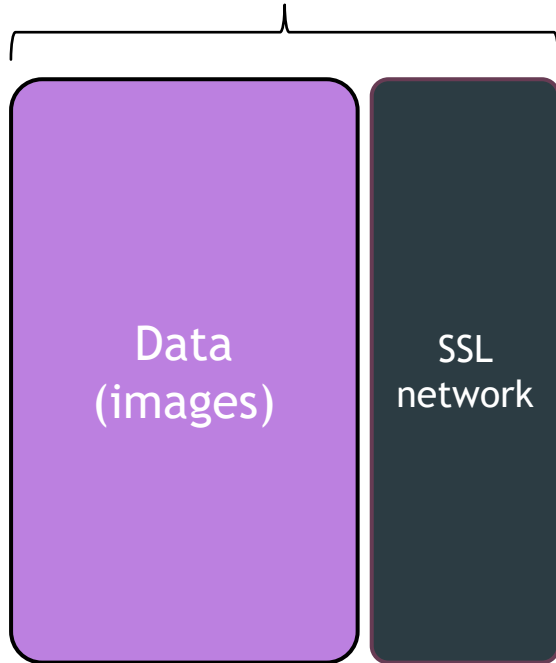
Self-Supervised Learning for MeerKAT Images

E. Lastufka

M. Audard, O. Bait, M. Dessauges-Zavadsky, M. Drozdova, T. Holotyak,
V. Kinakh, D. Piras, O. Taran, D. Schaerer, S. Voloshynovskiy

Self-supervised learning is crucial for training foundation models

Train



[Hayat et al 2021](#)

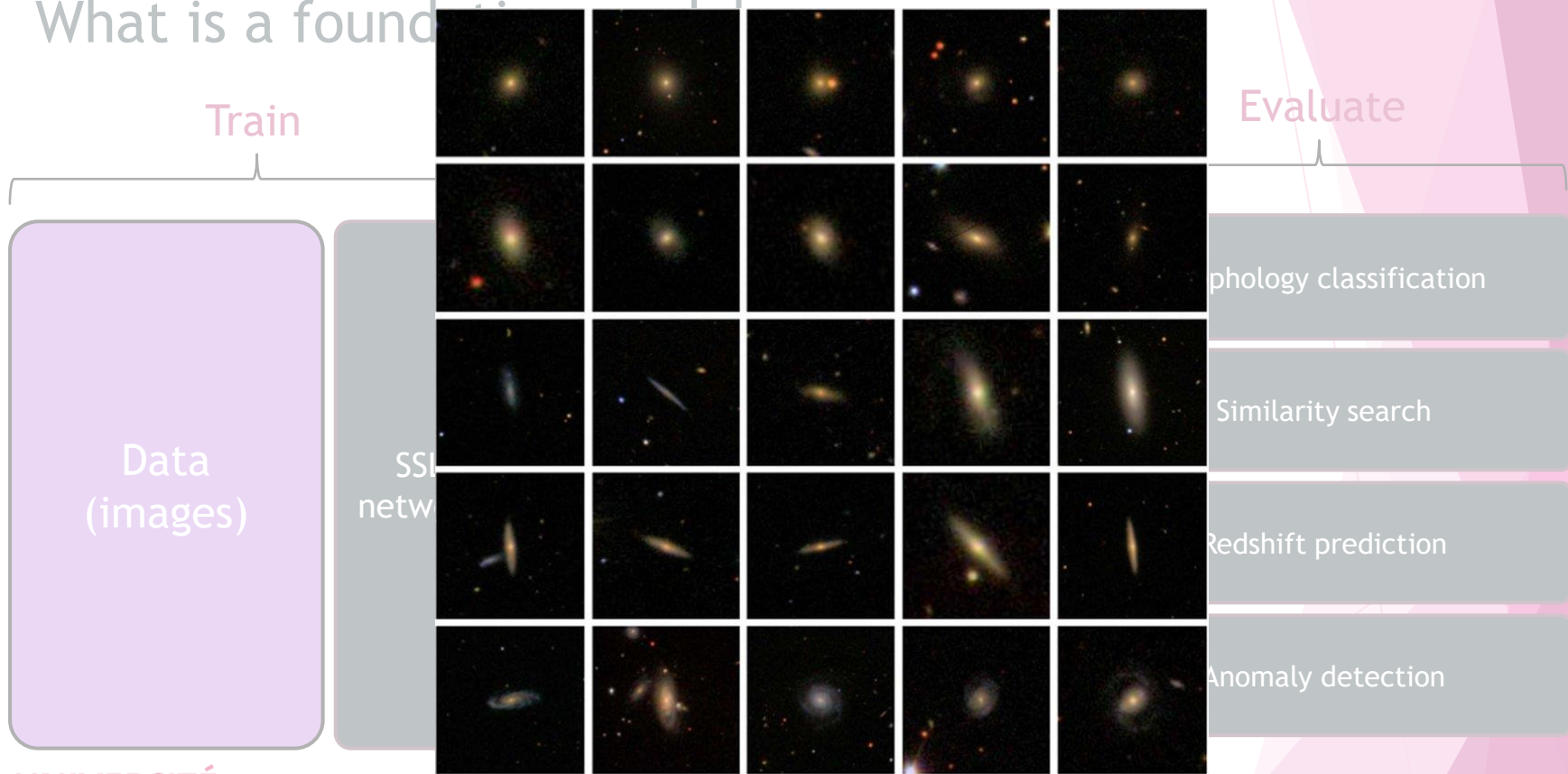
Evaluate

-
- Morphology classification
 - Similarity search
 - Redshift prediction
 - Anomaly detection
- A list of four evaluation tasks, each in a dark blue rounded rectangle, grouped under the "Evaluate" header. A bracket above the list spans the width of the "Evaluate" section.



What is a foundation model?

Galaxy Zoo SDSS (scales vary)

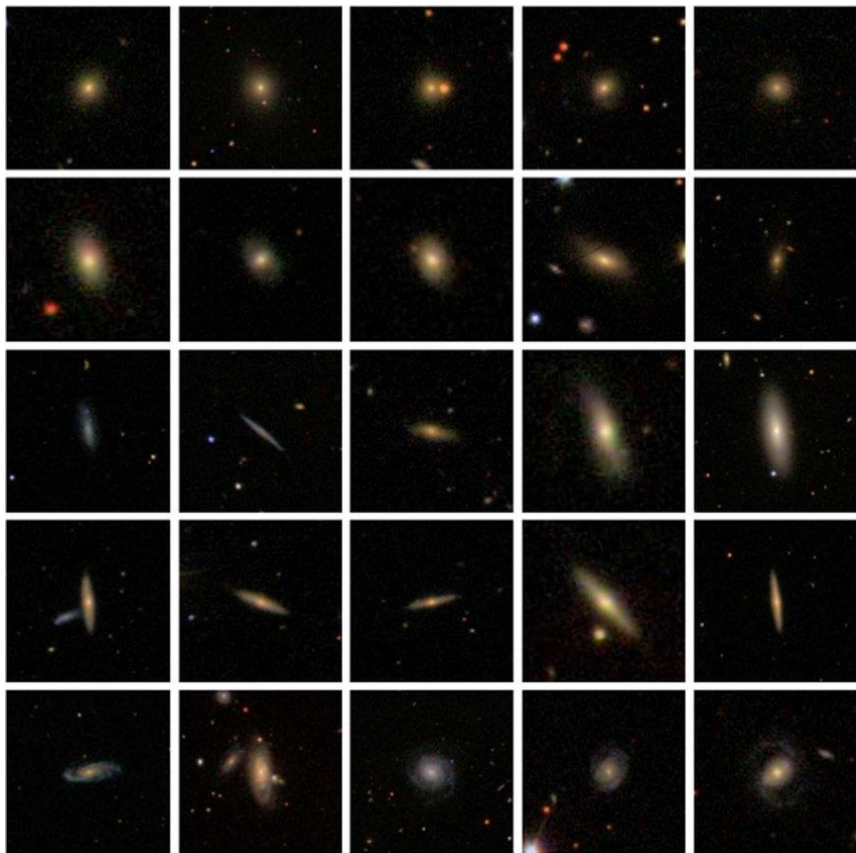


Current foundation models in astronomy

Regime	Work	Dataset	Network
Optical	Hayat et al 2021	Galaxy Zoo 2 (SDSS)	MoCo v2, Resnet50
Optical	Mohale & Lochner 2023	Galaxy Zoo DECaLS	BYOL, Resnet50
Radio	Slijepcevic et al 2023	Radio Galaxy Zoo	BYOL, Resnet18
Radio	Andrianomena & Tang 2023	Radio Galaxy Zoo	VDVAE, SimCLR BYOL, SimSiam Resnet34

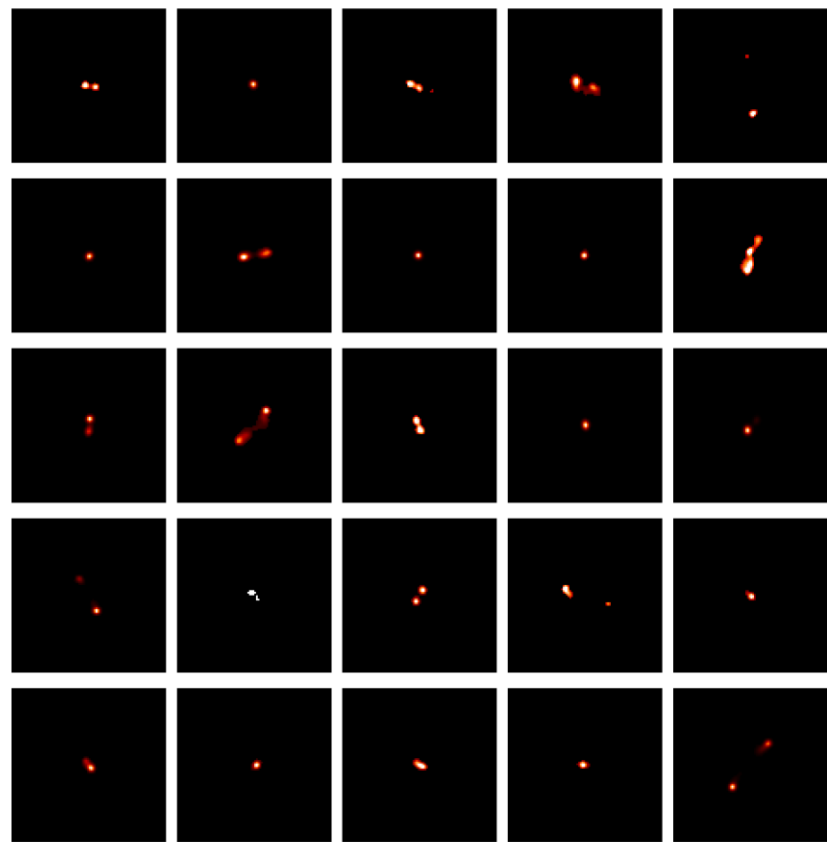


Galaxy Zoo SDSS (scales vary)

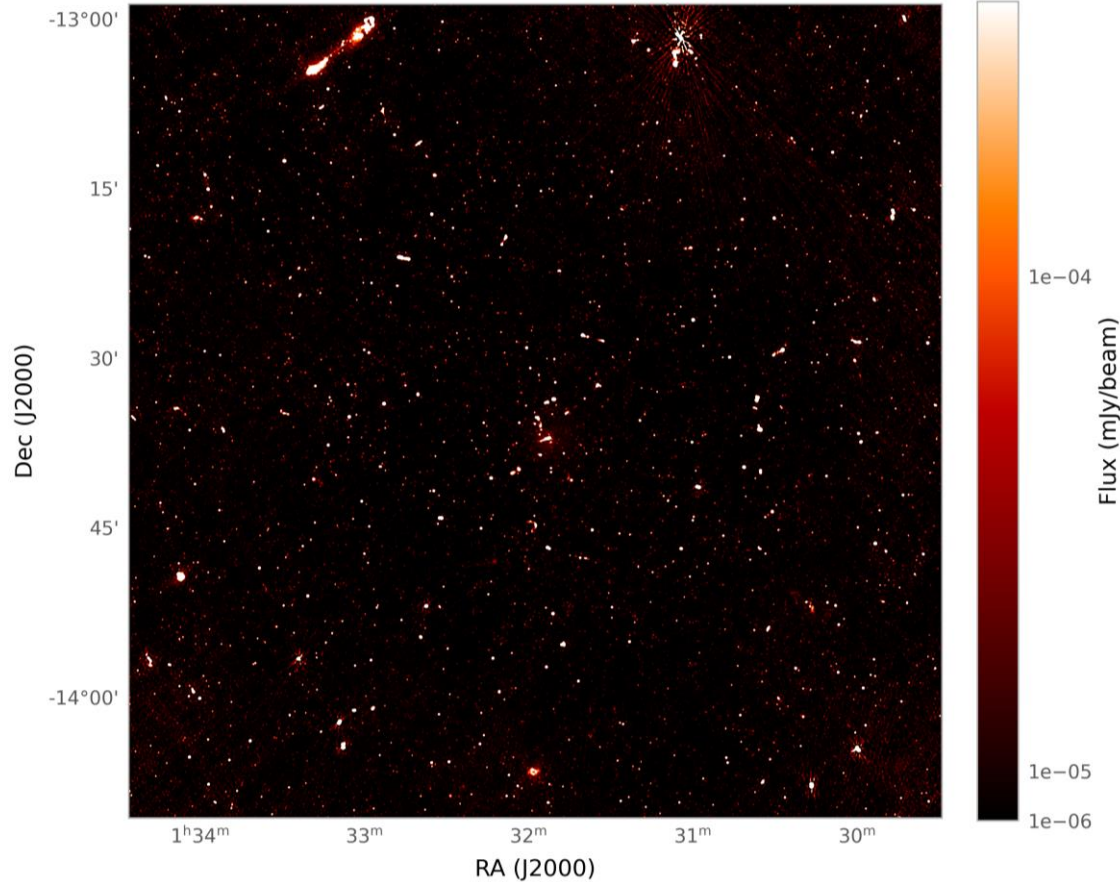


180''

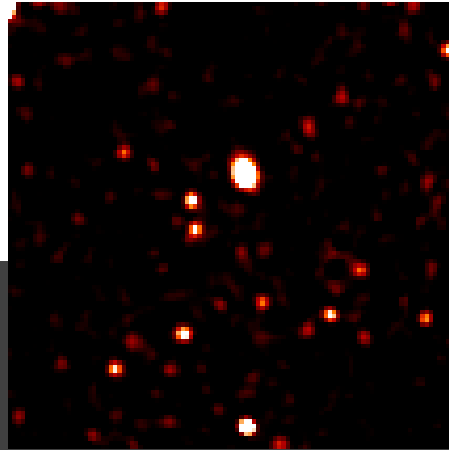
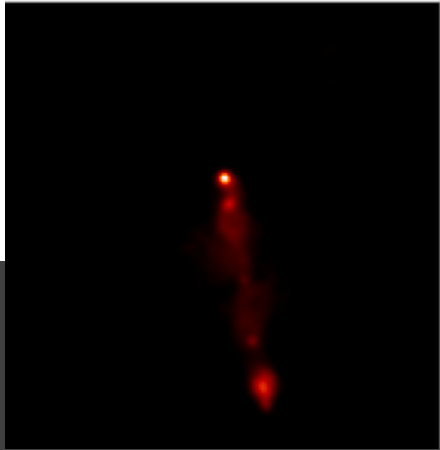
Radio Galaxy Zoo (VLA)



MeerKAT L-band Abell 209



Can we train multi-purpose foundation models with pipeline data?



MGCLS

MeerKAT Galaxy Cluster
Legacy Survey

- ▶ MeerKAT observations of 115 galactic clusters
- ▶ Wide-field 1.2 degree images
 - ▶ ~20,000 crops of pixel size 256 x 256
- ▶ Continuum CLEAN images

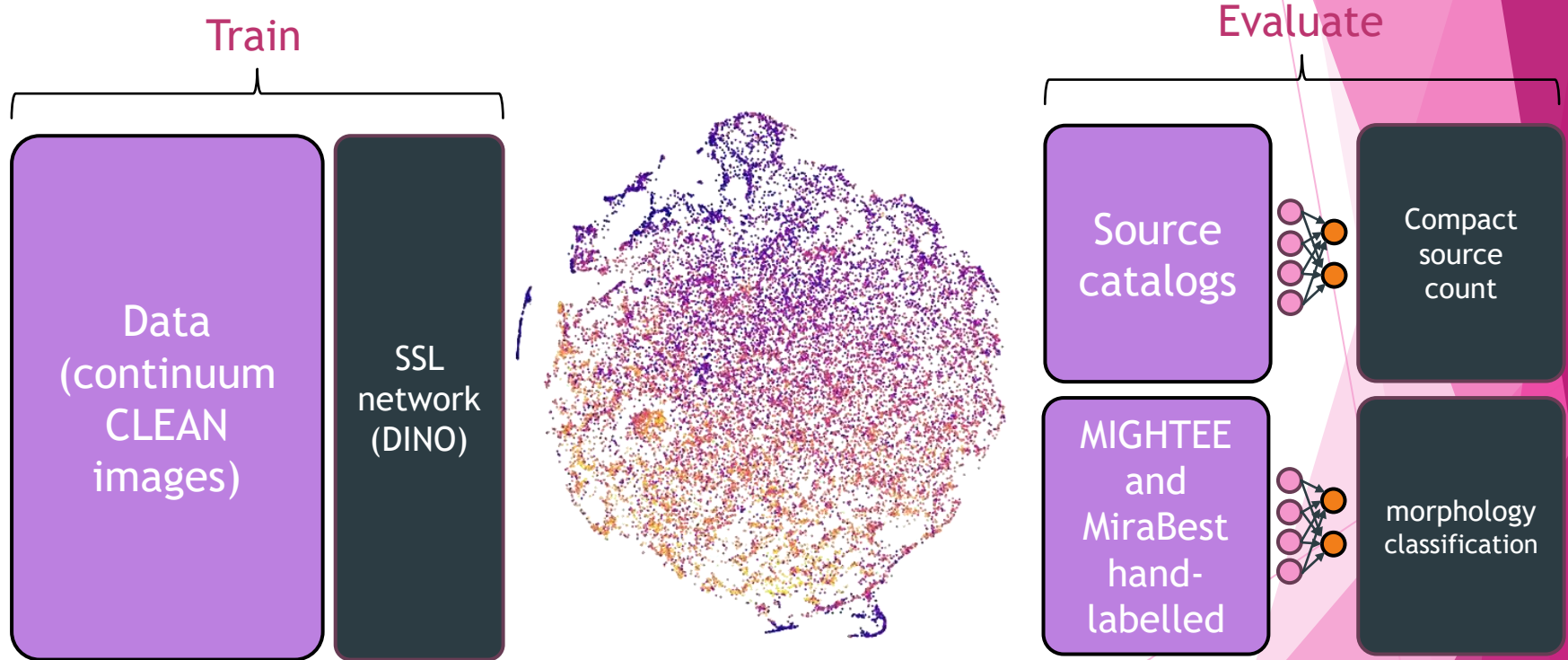


Knowledge distillation with no labels:

DINO

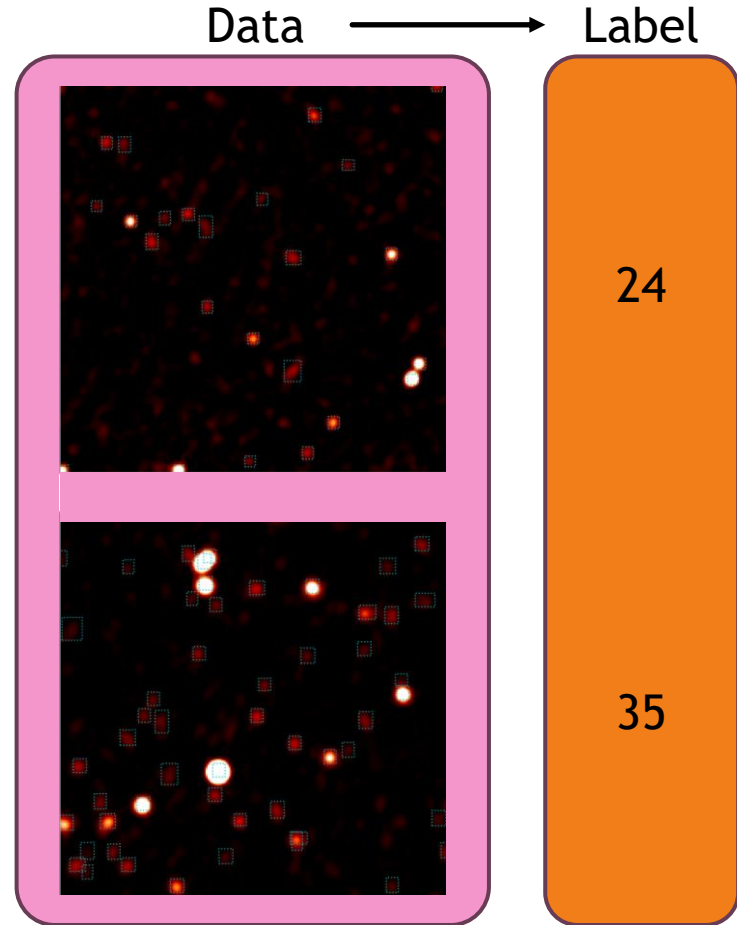


Can SSL create representations that characterize source-rich MeerKAT data?

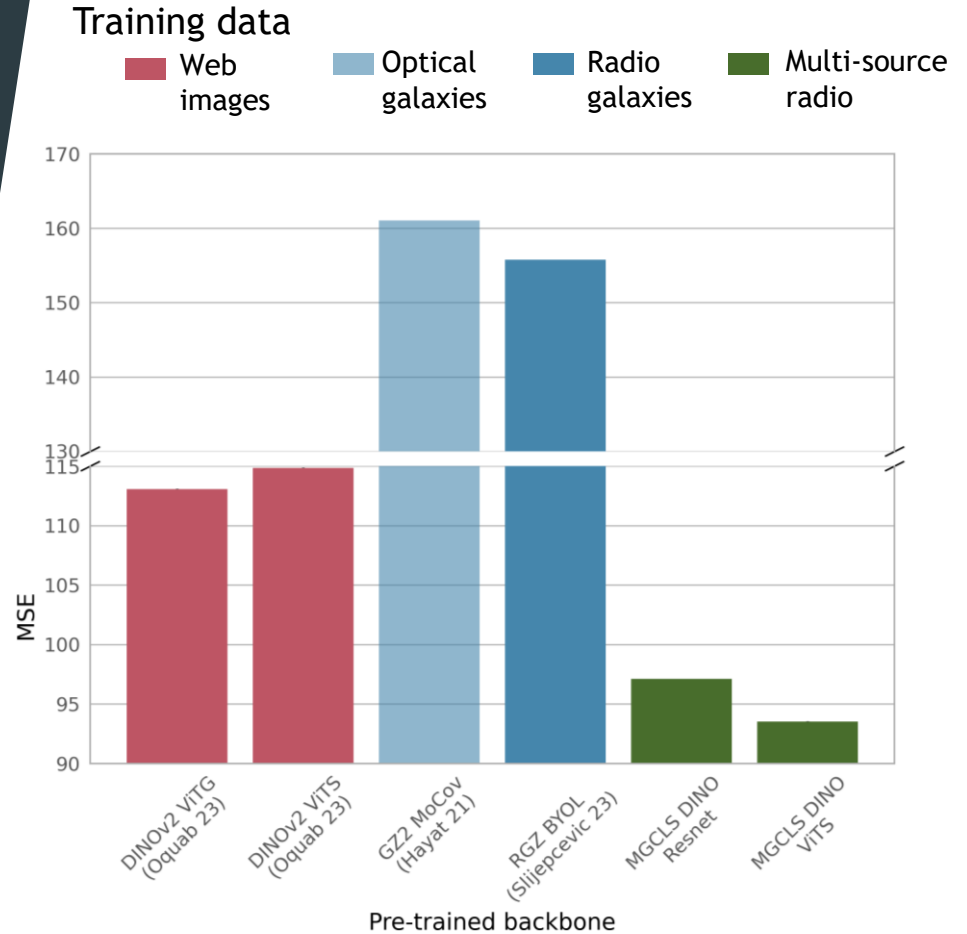


Evaluation: compact source count prediction

- ▶ Labels are from compact source catalog (pyBDSF)
- ▶ Task: predict source count with image as input
- ▶ Metric: mean squared error (MSE), the lower the better

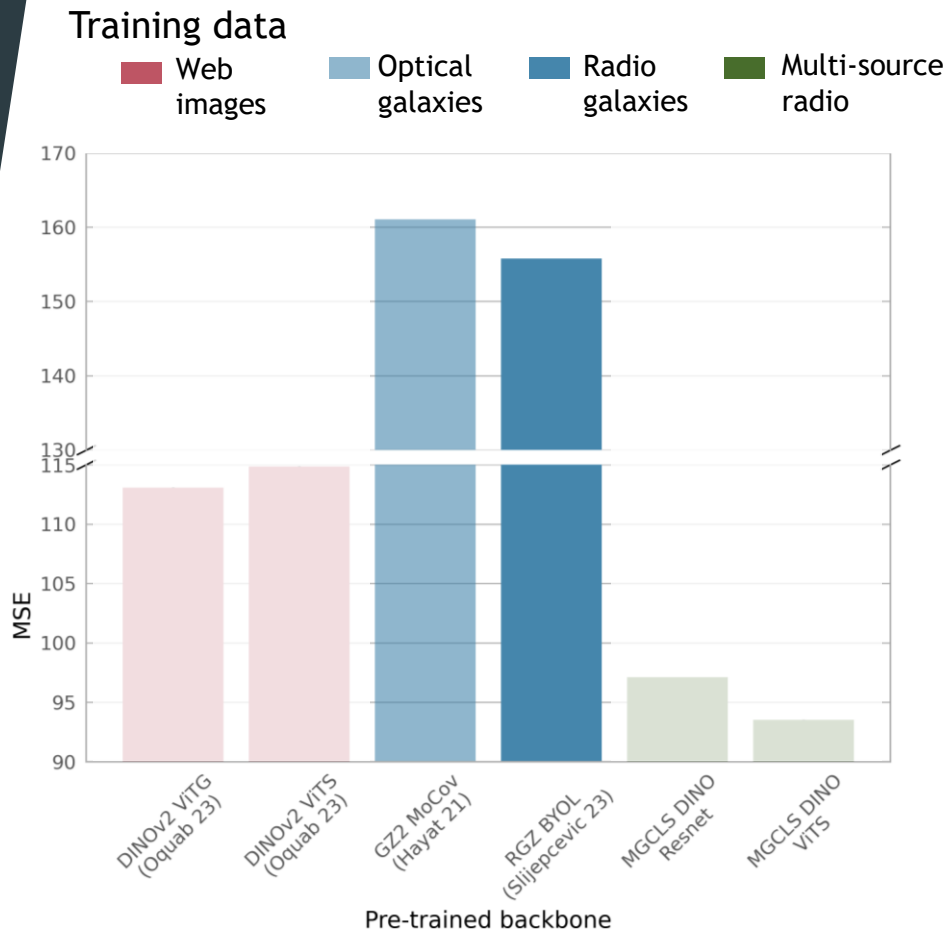


Evaluation: compact source count prediction



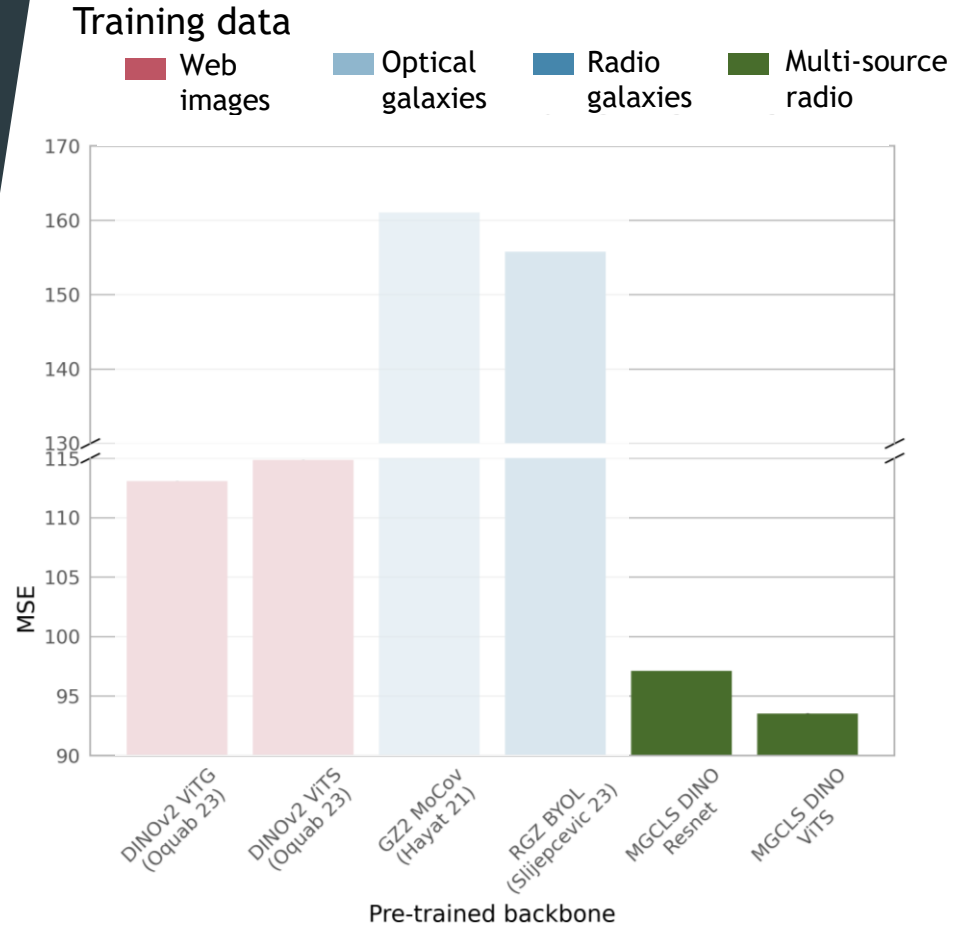
Evaluation: compact source count prediction

- ▶ Backbones trained on curated, single-galaxy images cannot perform this task!



Evaluation: compact source count prediction

- ▶ Domain-specific training data is best for this task

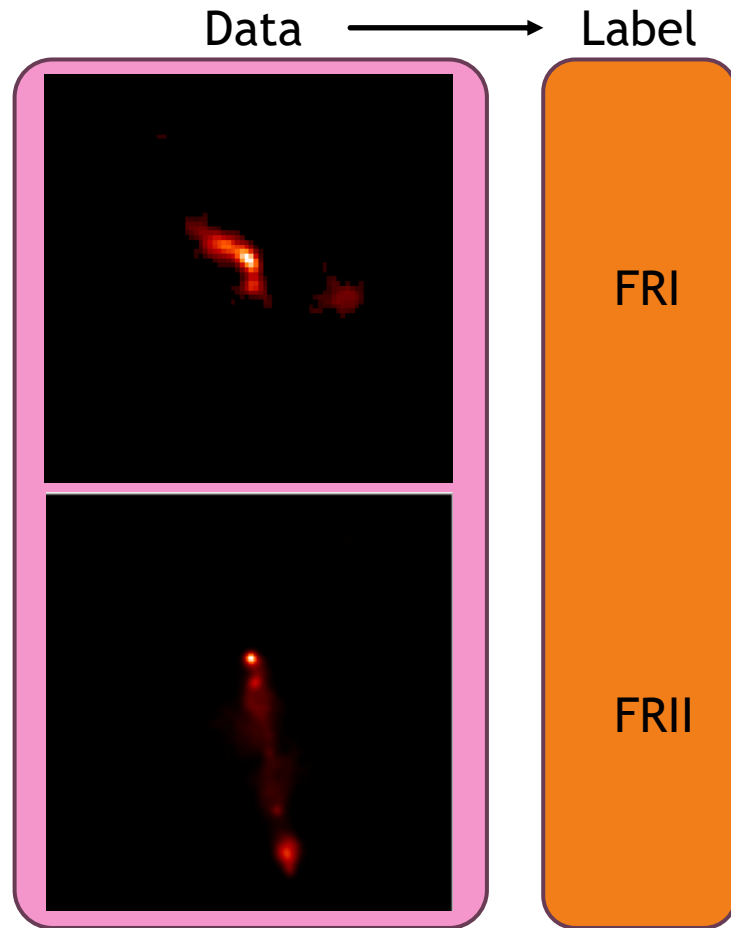


Can the network do more than one thing?



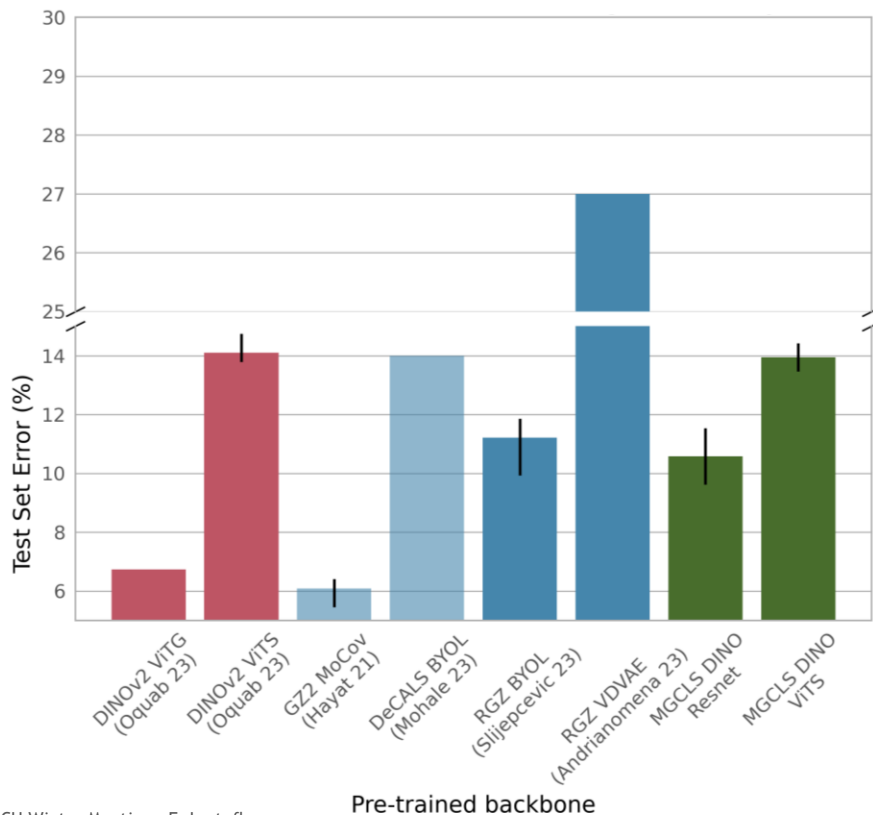
Evaluation: FRI/FRII galaxy morphology classification

- ▶ Public [MiraBest](#) dataset from VLA images
- ▶ Metric: Test set error (1-accuracy), lower is better



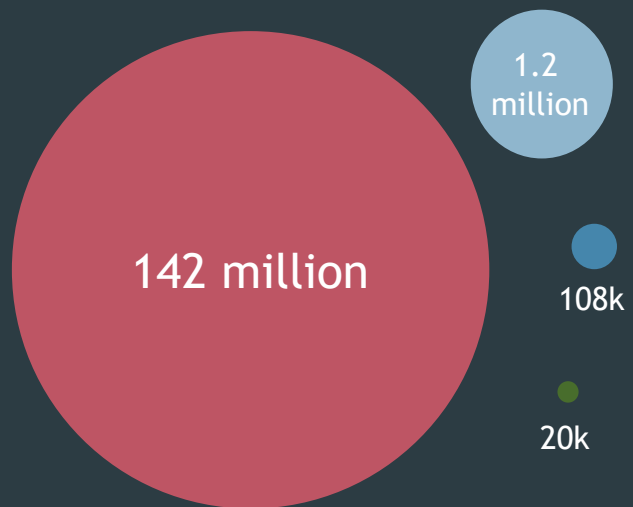
Evaluation: FRI/FRII galaxy morphology classification

Training data



Evaluation: FRI/FRII galaxy morphology classification

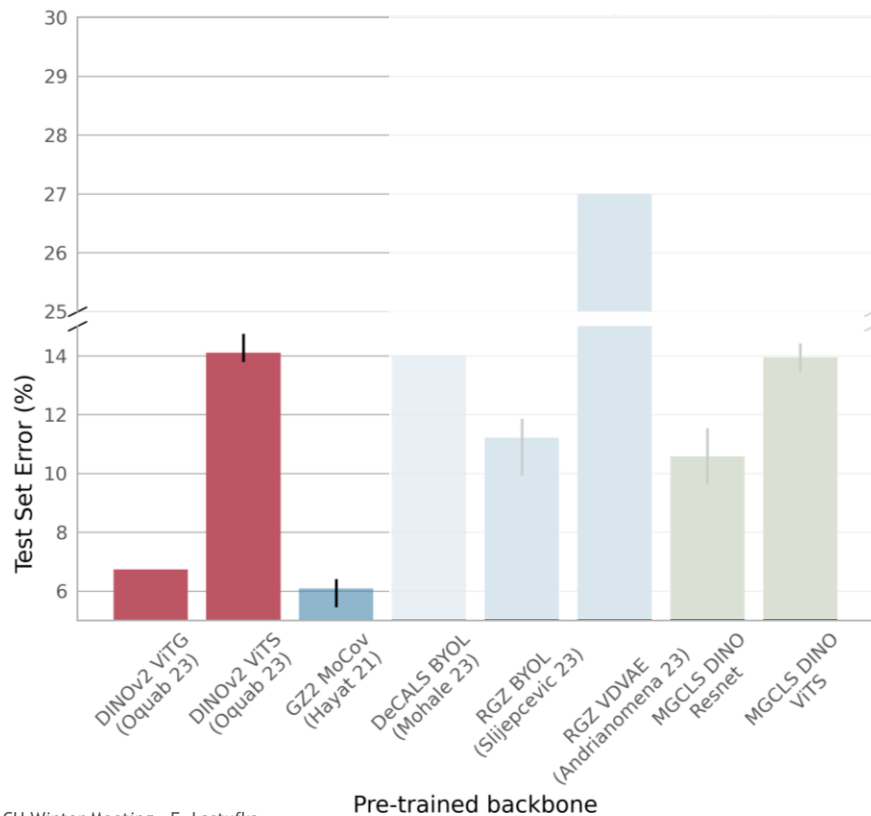
- ▶ Big training datasets gives good performance!



22-Jan 2024

Training data

- Web images
- Optical galaxies
- Radio galaxies
- Multi-source radio

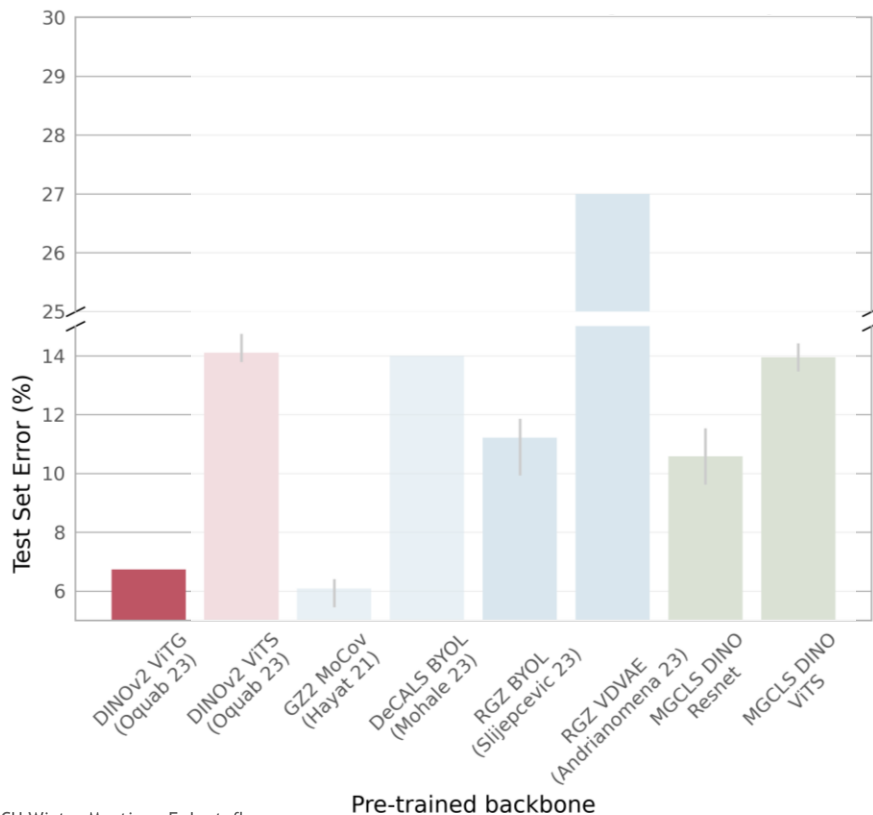


SKACH Winter Meeting - E. Lastufka

Evaluation: FRI/FRII galaxy morphology classification

- ▶ Big model makes up for lack of domain-specific training data?

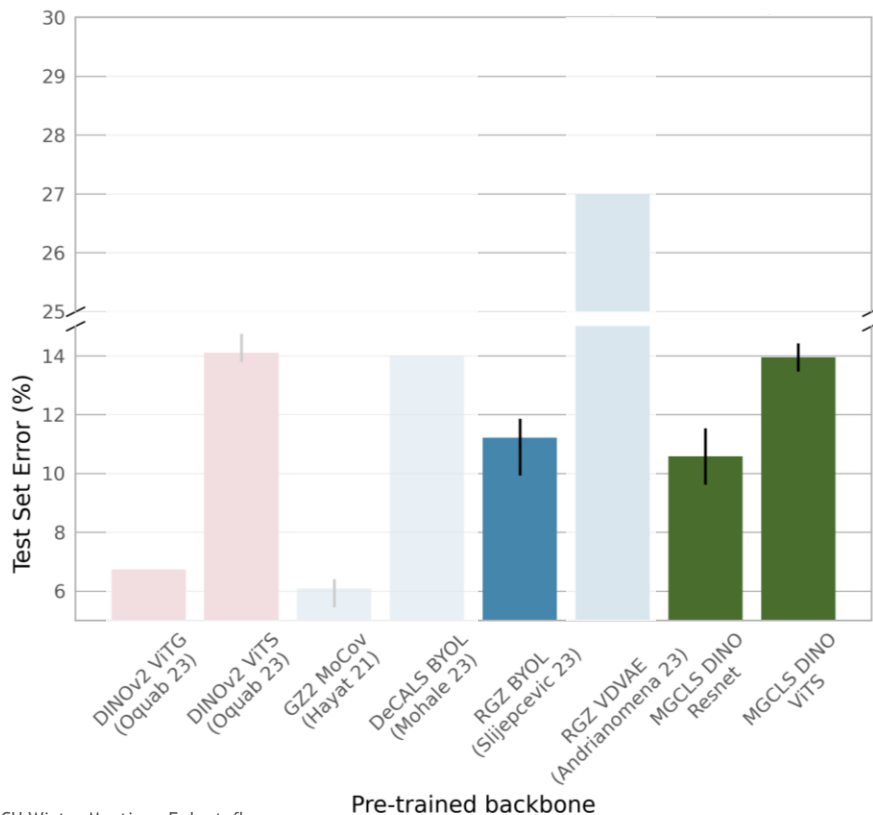
Training data



Evaluation: FRI/FRII galaxy morphology classification

- ▶ Our model trained on source-rich data can perform as well as academic state-of-the-art trained on highly curated data

Training data



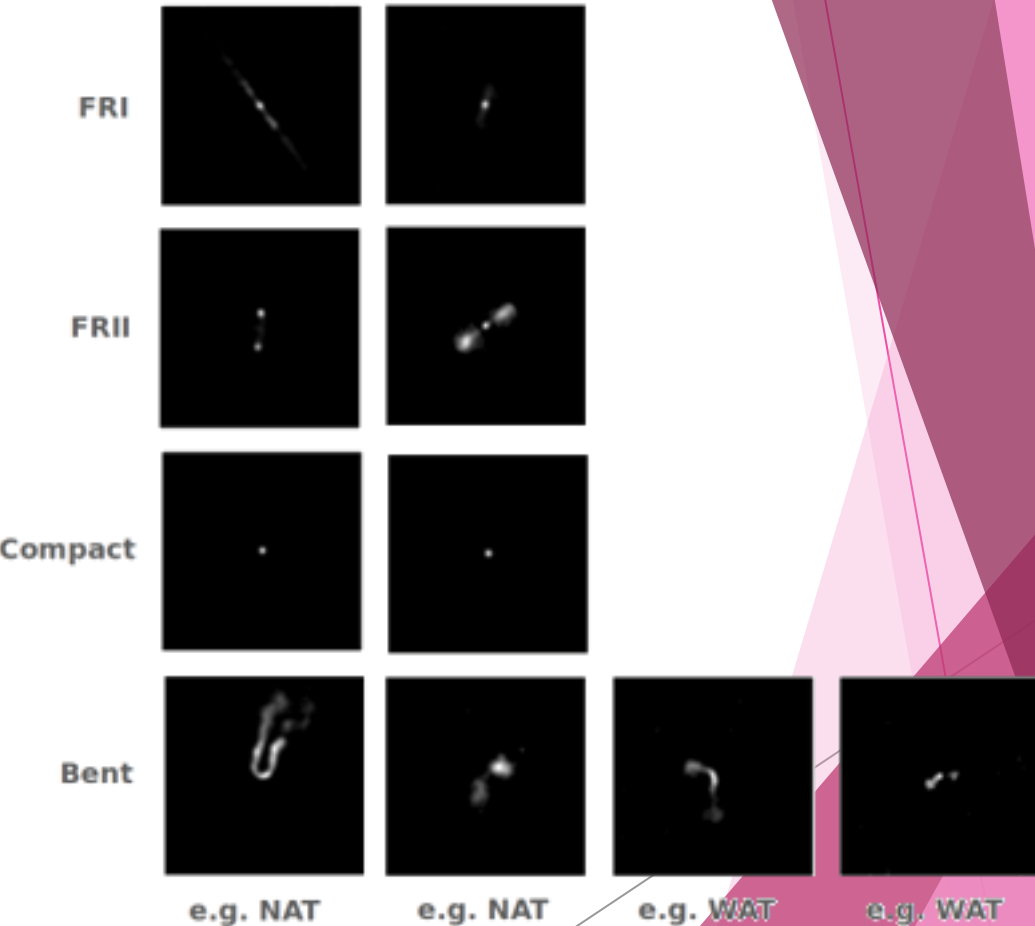
Can the network do complex things?

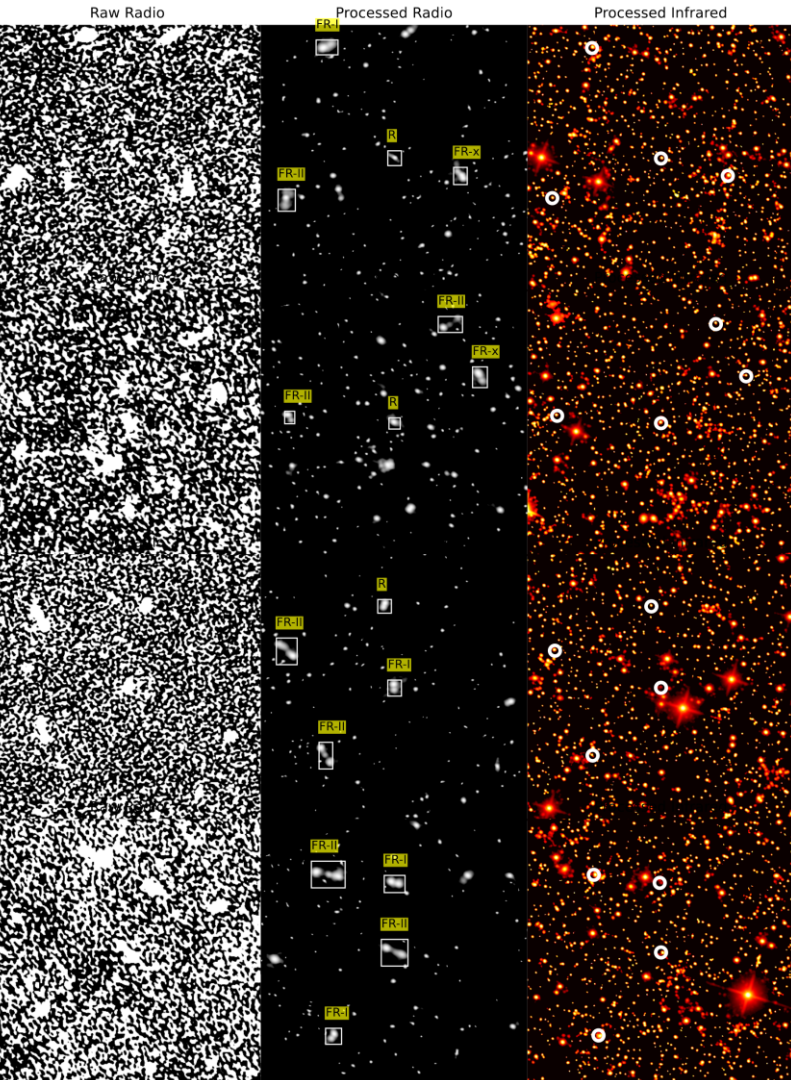


Multi-class morphology classification

- ▶ RadioGalaxyDataset

- ▶ VLA images of four morphology classes





Source detection and segmentation

- ▶ RadioGalaxyNET
 - ▶ ASKAP + WISE radio/infrared dataset for object detection

Future foundation models in astronomy

Training data can come straight from the data processing pipeline!

A mixture of source-rich and single-source?

Evaluated on multiple high-complexity tasks



Extra slides



Evaluation: multi-class galaxy morphology classification

