

Sustainability of Extreme-Scale Simulations with SPH-EXA

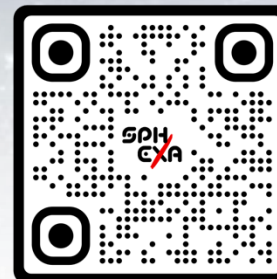
SKA Days 2024

September 02-04, 2024, Geneva

Osman Seckin Simsek



<https://github.com/unibas-dmi-hpc/SPH-EXA>



Florina Ciorba (PI)
Ruben Cabezon (Co-PI)
Osman Seckin Simsek
Yiqing Zhu
Lukas Schmidt
José Escartin



Lucio Mayer (Co-PI)
Noah Kubli
Darren Reed



Sebastian Keller
Jean-Guillaume Piccinali
Jean Favre
Jonathan Coles



Axel Sanz (UPC)
Joseph Touzet (Paris-Saclay)

TGSF: The role of Turbulence and Gravity in Star Formation



EuroHPC
Joint Undertaking

Extreme Scale Access

Allocation: 22,000,000 GPUh* on LUMI-G

Duration: 12 months, Nov.'23 – Oct.'24

*Largest allocation in Europe to date.



Objectives

Study the formation of pre-stellar cores and their initial mass function at unprecedented resolution

Scalability limitation for previous codes

Study turbulent transport and mixing

More natural with Lagrangian codes

Contribute to the general theory of turbulence (Lyapunov exponents)

Study the load imbalance, performance, and energy consumption at unprecedented scales

HPC research at extreme scale

Study large scale compression techniques for checkpointing, compression, and visualization

Publications | News

PASC project principle investigators discussing their new astrophysical simulation code, which helped them win a large

Cosmology & Astrophysics

Computer Science

TGSF: The role of Turbulence and Gravity in Star Formation

Definition

Sonic scale (l_s) : is the scale at which the transition from supersonic to subsonic turbulence occurs.

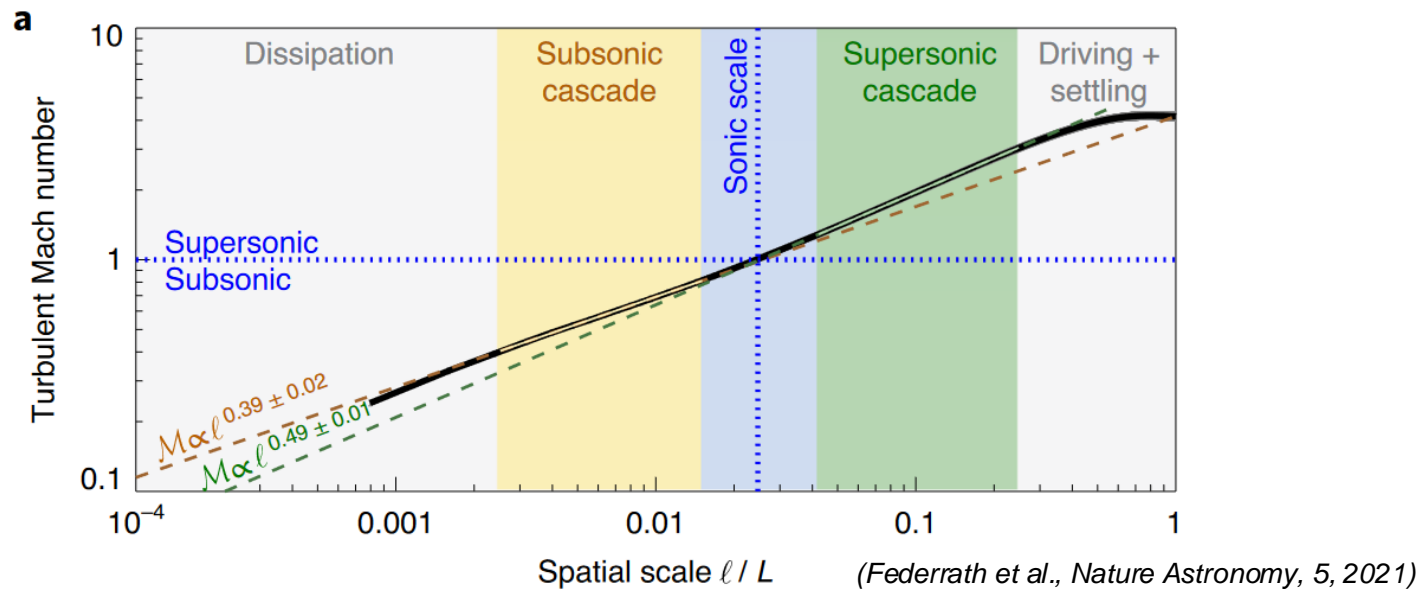
$$l_s = \phi_s L(\mathcal{M})^{-2}$$

ϕ_s encompasses our lack of knowledge about the exact position of the Sonic scale. Usually taken as $\phi_s = 1$

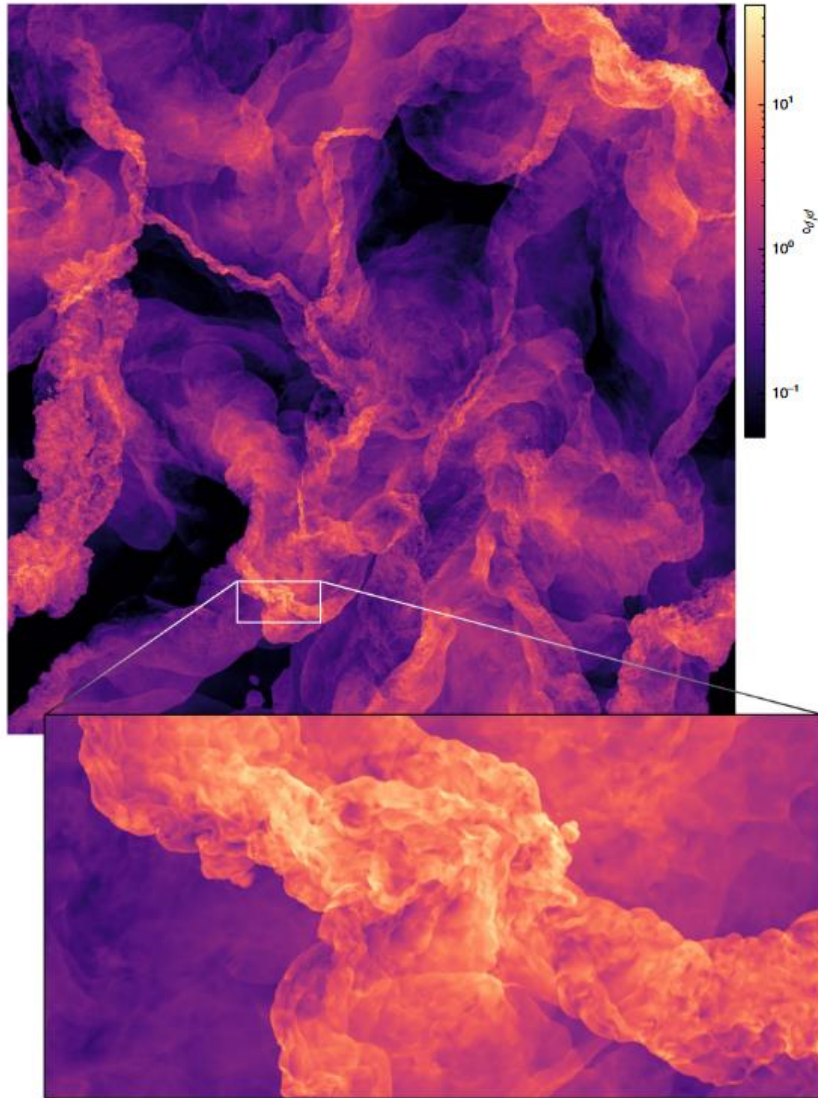
Nevertheless, a large-scale simulation (Federrath et al. 2021) has directly measured ϕ_s to be x2.4 smaller.

This pushes the collapse scale to smaller scales than previously considered and it has a critical relevance in the predictive power of star formation theories.

In order to test this, **self-gravity** must be included in such simulations.

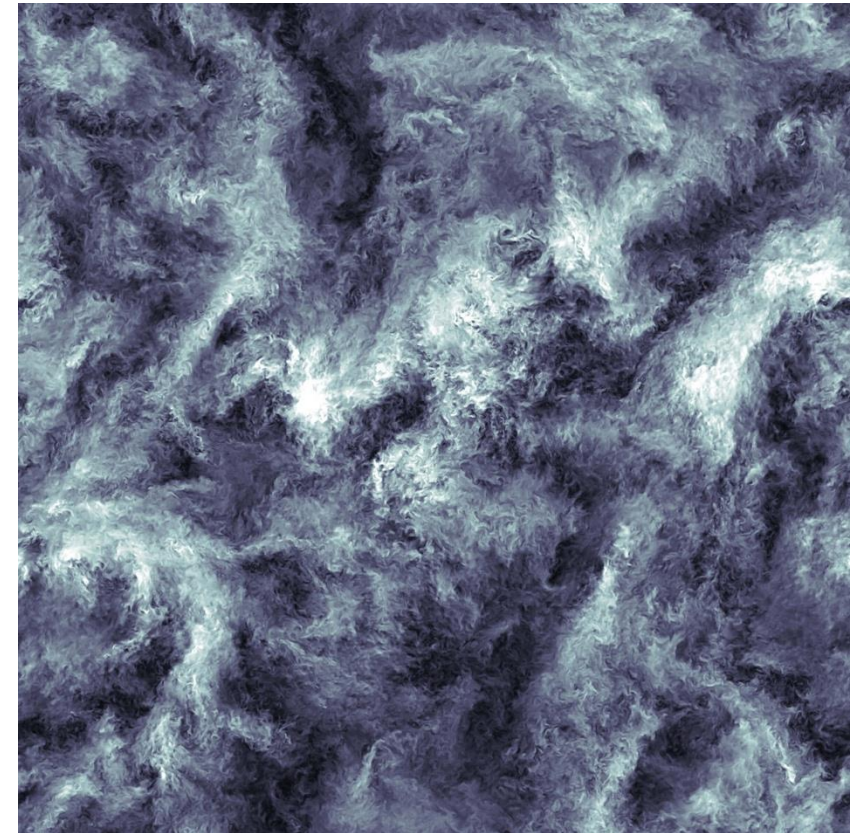


TGSF: The role of Turbulence and Gravity in Star Formation



FLASH
Eulerian code
10,048³ grid cells
Hydrodynamics only (no gravity)
CPU only (65,536 cores)

SPH-EXA
Lagrangian code
10,079³ SPH particles
Hydrodynamics + gravity
CPU + GPU (16,416 GPUs)



Gas density contrast distribution of ISM turbulence. (Federrath et al., Nature Astronomy, 5, 2021)

Velocity field distribution of subsonic turbulence. 3000³ particles (SPH-EXA team, 2023)

SPH-EXA: Smoothed Particle Hydrodynamics at Exascale



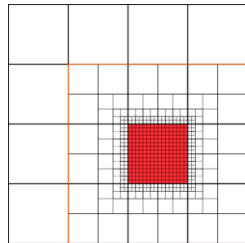
is a *scalable* smoothed particle hydrodynamics simulation framework **interdisciplinarily co-designed** by computational physicists and computer scientists to exploit **Exascale** supercomputers.

SPH-EXA: Framework Components



SPH-EXA application frontend

Parallel I/O and test case setup (ICs)



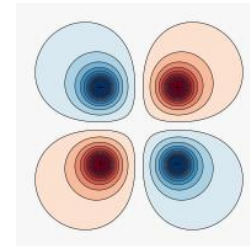
Cornerstone

Octree and domain decomposition framework

SPHYON

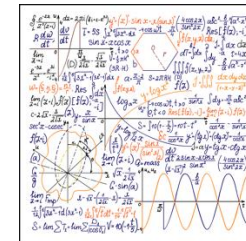
SPH

Hydrodynamics solver



Ryoanji

N-body gravity solver



Physics modules

*Radiative cooling
Nuclear reactions
Star formation
Stellar feedback*

SPH-EXA: Optimization Strategy



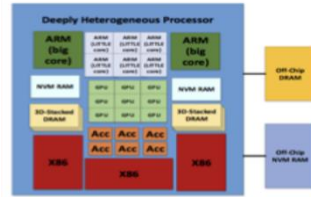
Scalability

weak, strong



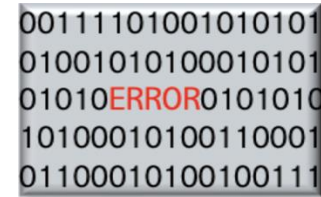
Scheduling & Load-balancing

*dynamic, adaptive
asynchronous
execution*



Heterogeneity

*portability on
various CPU and
GPU architectures*



Fault-tolerance

*silent error
detection and
recovery*



Energy

*measurement
reporting
efficiency*

Sustainable Computing Motivation – Astronomy and Astrophysics



⑦ C... ①A Computing: big data.

Carbon footprint = energy used x carbon intensity

gCO_2e

kWh

gCO_2e/kWh



$$E = t \times (P_c + P_m) \times PUE$$

Running time (h)

Power draw of processing cores (W)

Power draw from memory (W/GB)

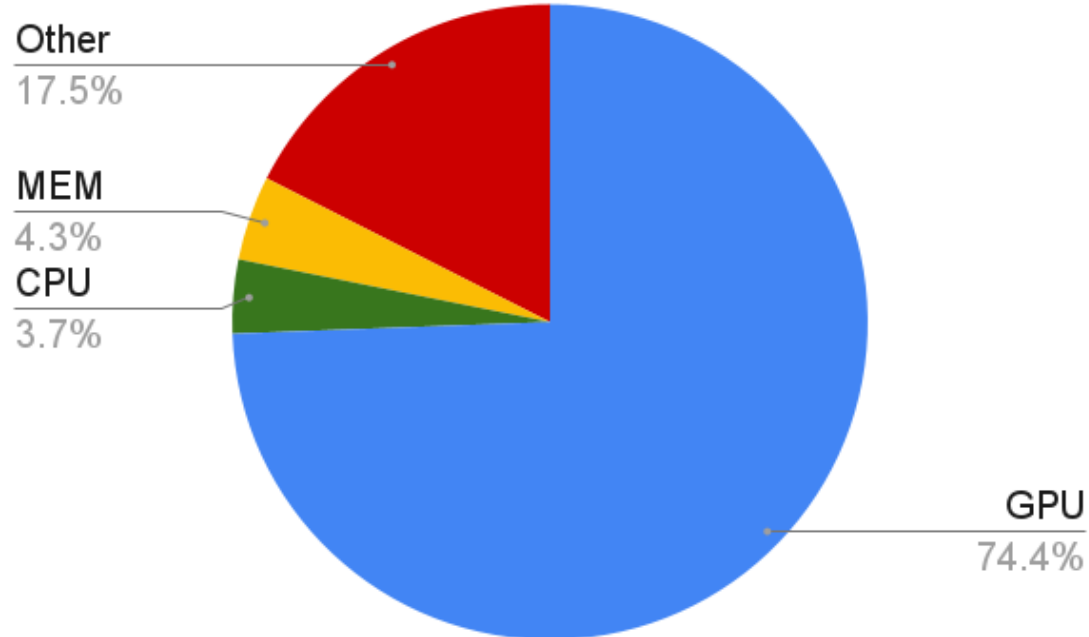
Efficiency of the data centre

ure
or
s the
tia
nst
mend
se of efficient
l architectures
s scientists
to ensure that all
to achieve the
minimising the

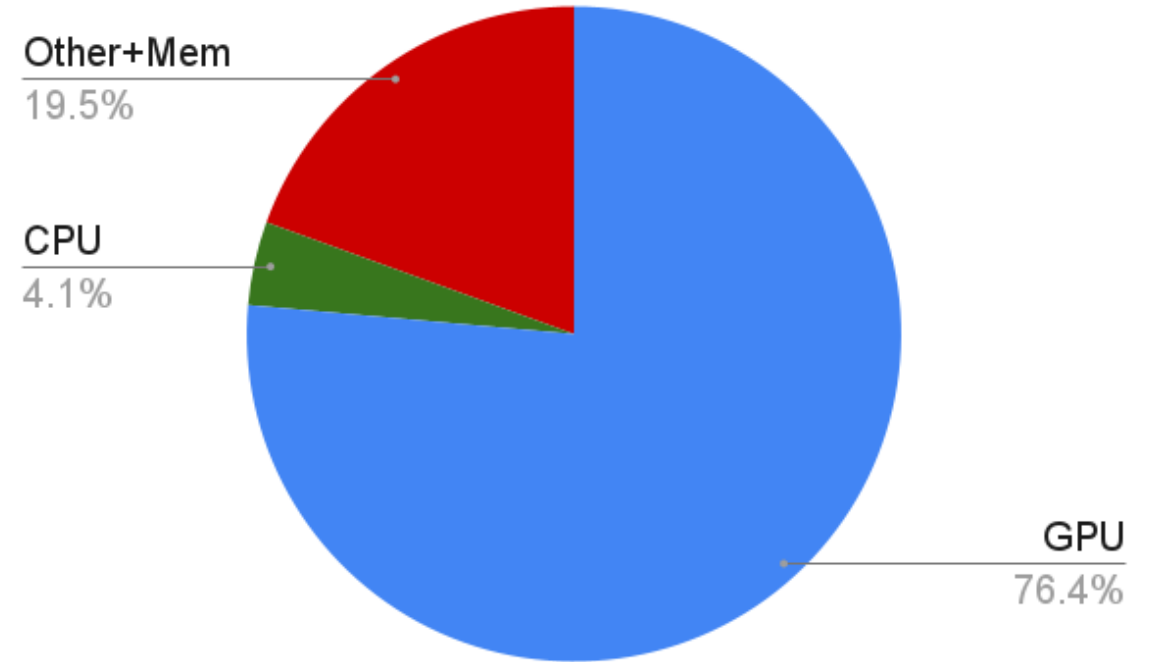
Source: <https://www.astronet-eu.org>

Energy Measurement and Reporting: Device Breakdown

Turbulence on LUMI-G AMD MI250X 32 MPI Ranks

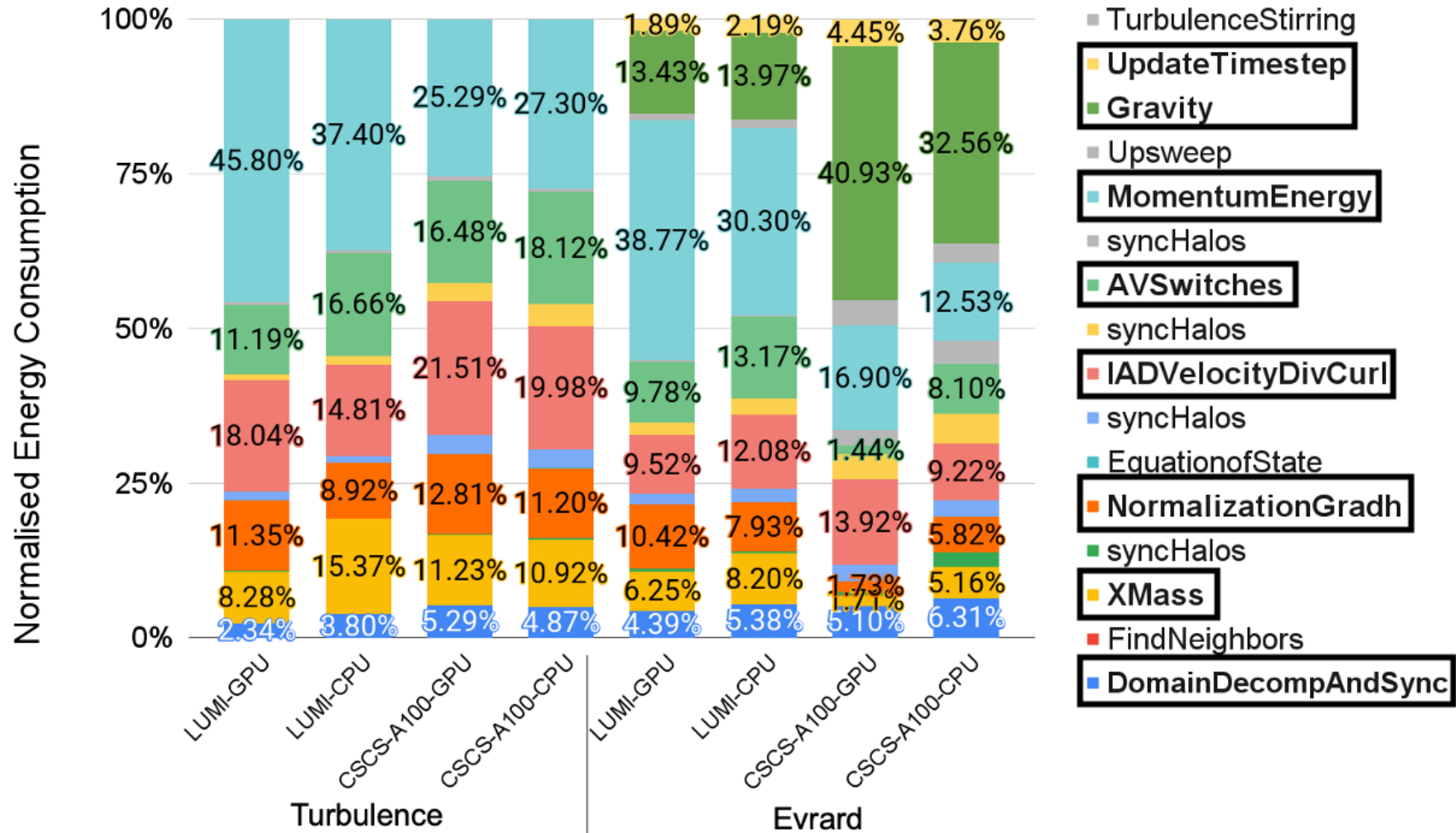


Turbulence on CSCS Nvidia A100 32 MPI Ranks

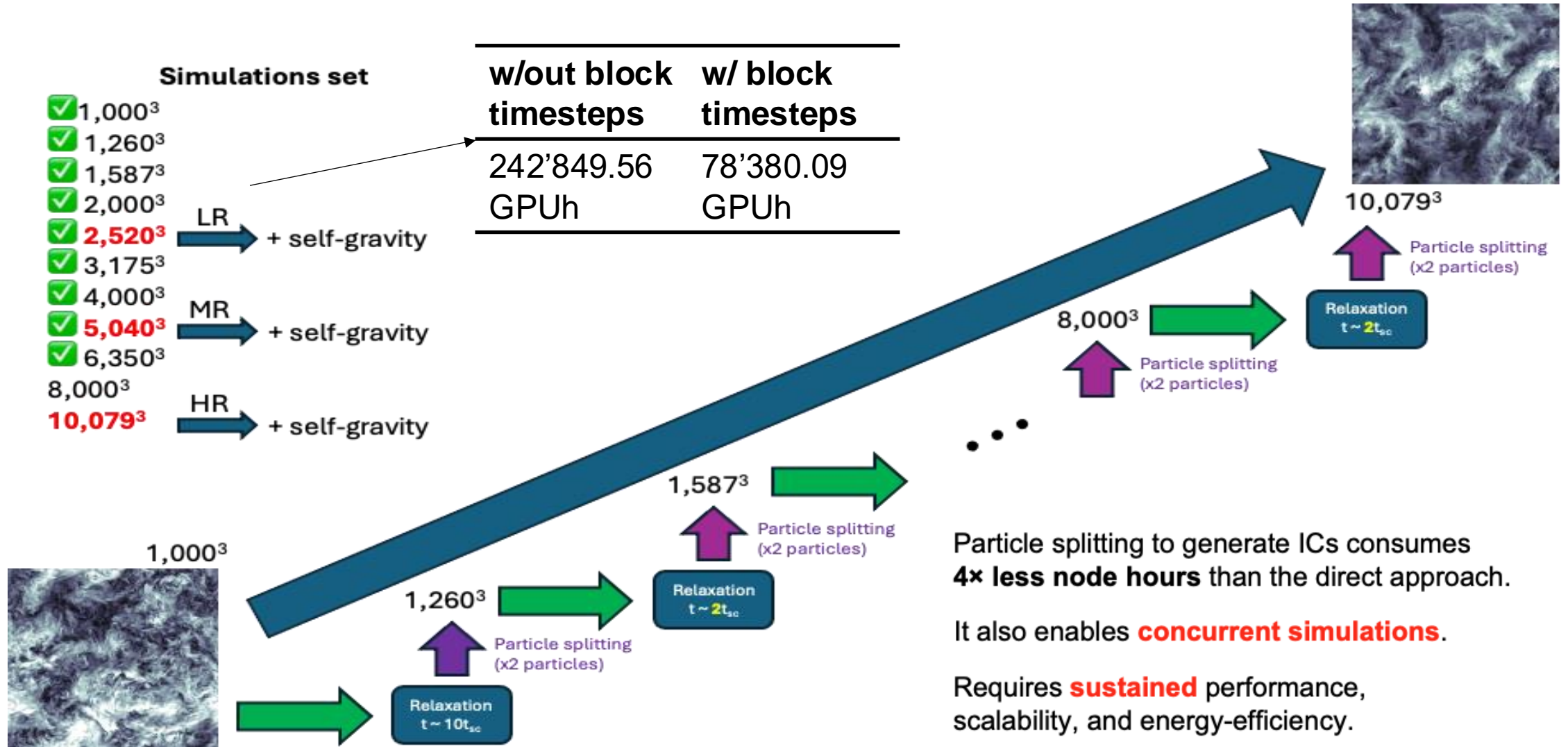


Energy Measurement and Reporting: Functional Breakdown

Code functional breakdown, 100 time-steps, 32 MPI Ranks on 4 LUMI-G Nodes and 8 CSCS-A100 Nodes



TGSF Simulation Plan

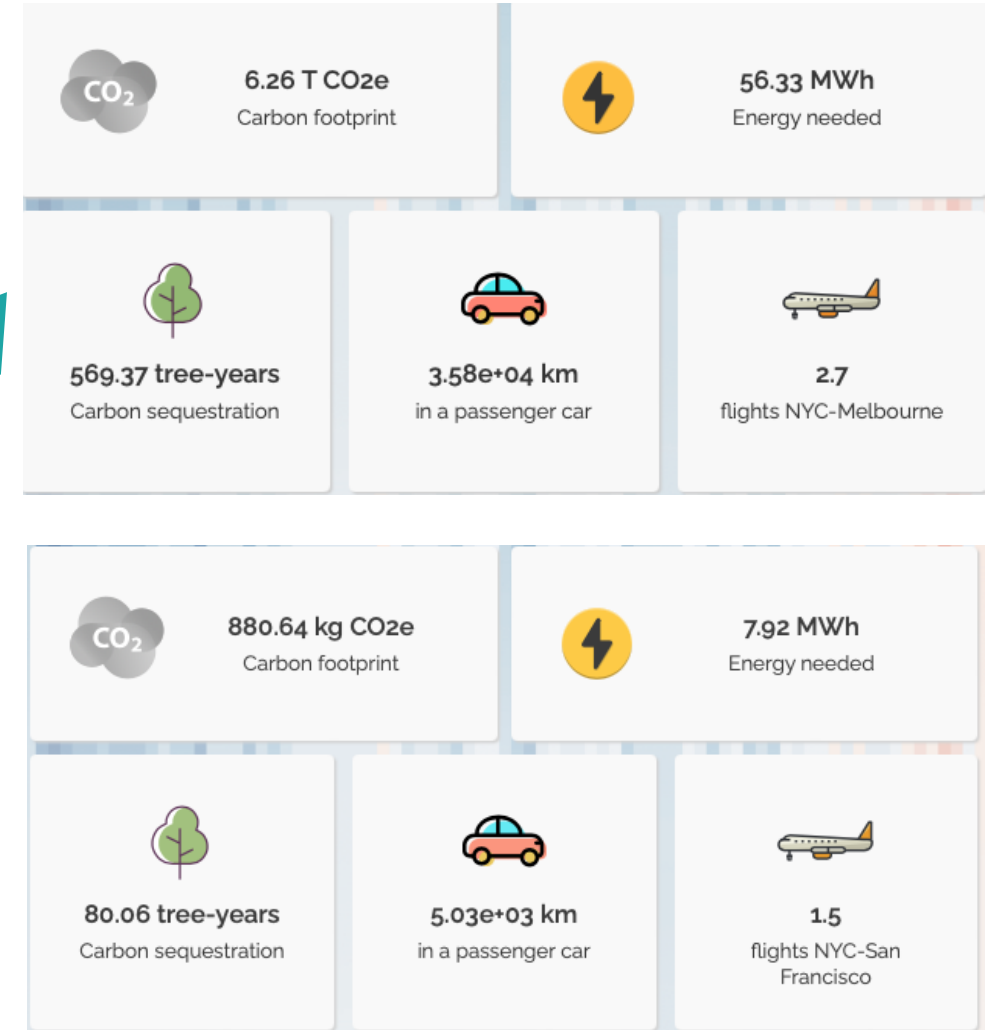


TGSF Low-Resolution Simulation Energy Measurements

Simulations set

✓ 1,000 ³	
✓ 1,260 ³	
✓ 1,587 ³	
✓ 2,000 ³	
✓ 2,520³	LR → + self-gravity
✓ 3,175 ³	
✓ 4,000 ³	
✓ 5,040³	MR → + self-gravity
✓ 6,350 ³	
8,000 ³	
10,079³	HR → + self-gravity

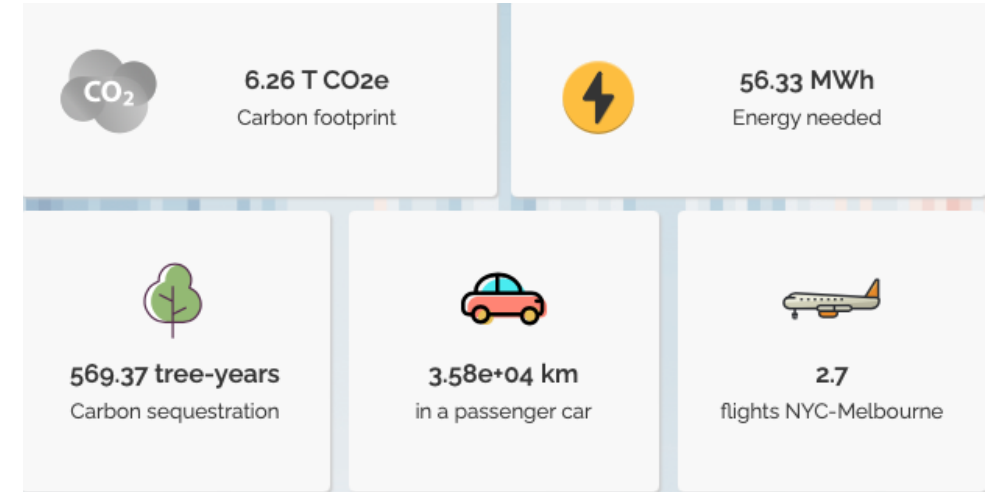
- Total of 35 jobs.
 - 6 failed and 29 successful.
- Successful jobs = 242'849.56 GPUh
 - Measured value.
- Failed jobs = 34'082.44 GPUh
 - Calculated value.



TGSF Low-Resolution Simulation Energy Measurements

Simulations set

✓ 1,000 ³	
✓ 1,260 ³	
✓ 1,587 ³	
✓ 2,000 ³	
✓ 2,520 ³	LR → + self-gravity
✓ 3,175 ³	
✓ 4,000 ³	
✓ 5,040 ³	MR → + self-gravity
✓ 6,350 ³	
8,000 ³	
10,079 ³	HR → + self-gravity



Measured energy breakdown of successful jobs per device

- Total of 35 jobs.
 - 6 failed and 29 successful.
- Successful jobs = 242'849.56 GPUh
 - Measured value.
- Failed jobs = 34'082.44 GPUh
 - Calculated value.

