

Radio galaxy image generation with the scattering transform

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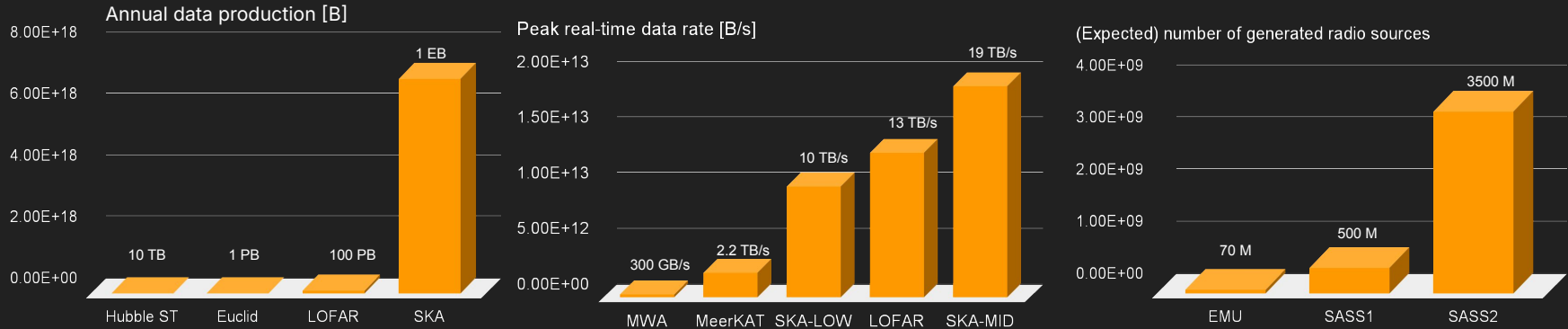
Supervisor: Emma Tolley

Co-supervisor: Rémi Poitevineau

3rd of September 2024



With great telescopes comes great data challenges



Credit: NASA, ESA, IRA, SKAO

Lonsdale et al., 2009; Boot & Jonas, 2012; Labate et al., 2022; van Haarlem et al., 2013; Swart et al., 2022

Norris et al., 2011; Norris et al., 2014;

Machines learn differently from humans

Interpretation



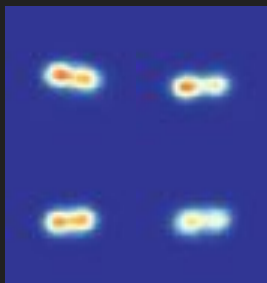
Biases



Generalisation



Reliability

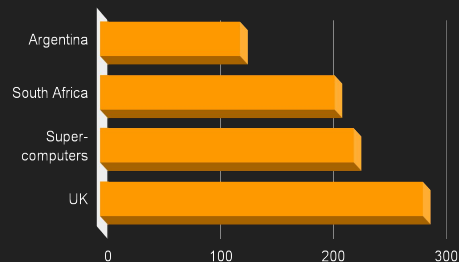


Scalability



Energy consumption

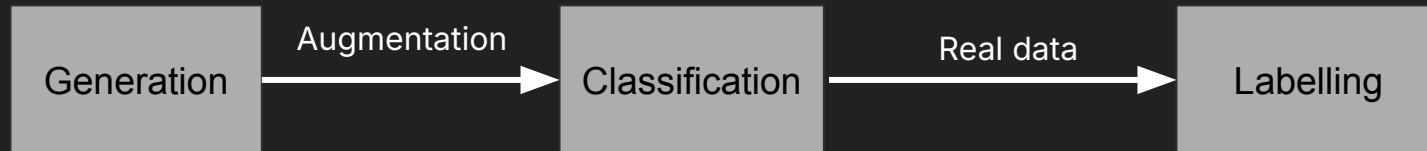
Electricity consumption in TWh



Credit: Enerdata, IEA, 2020

Content

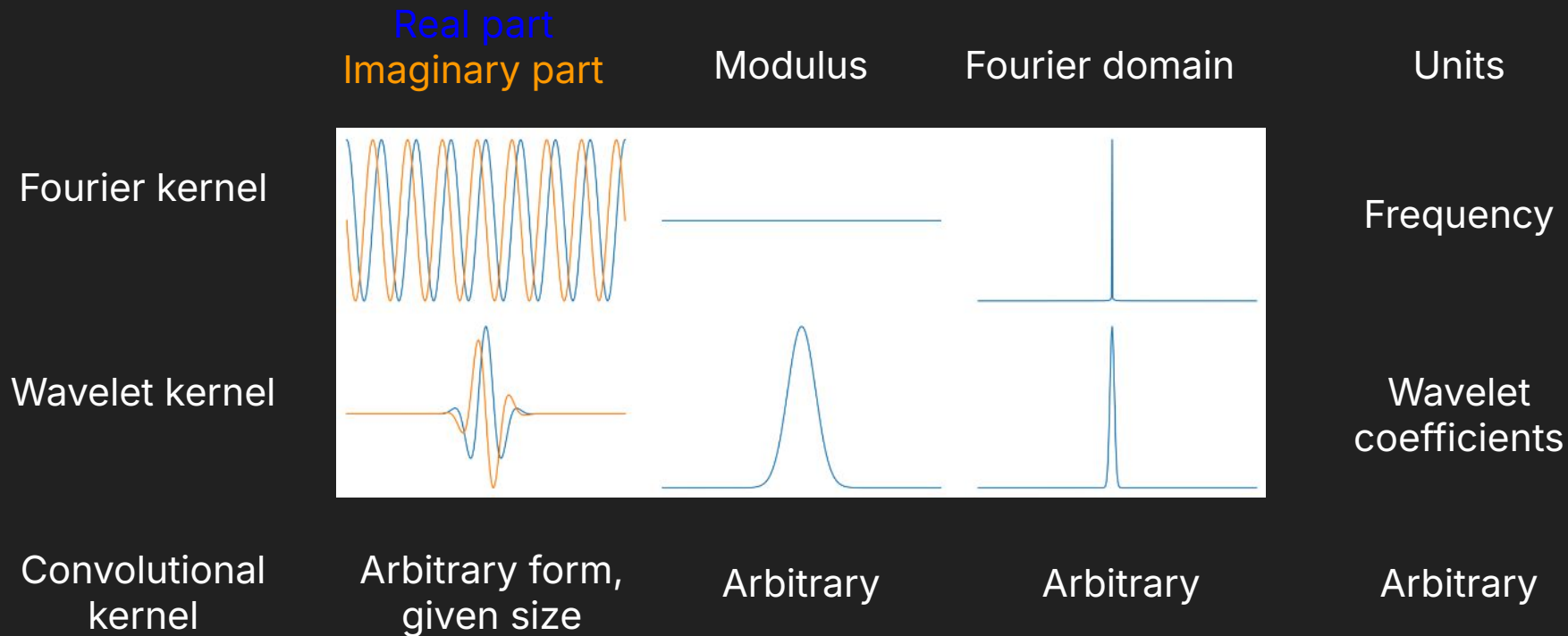
1. **The scattering transform** - A non-trainable CNN
2. **Generation** - Shallow and deep image generation
3. **Conclusions and outlook**



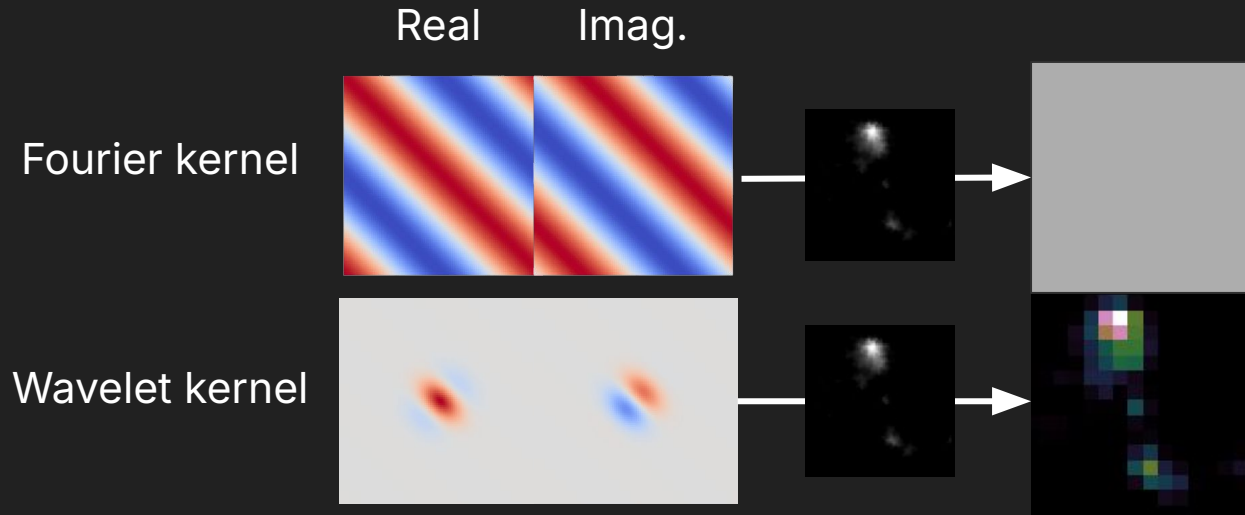
Scattering transform

A non-trainable convolutional neural network

The scattering transform uses wavelet kernels



Wavelet kernels are localised and extract features

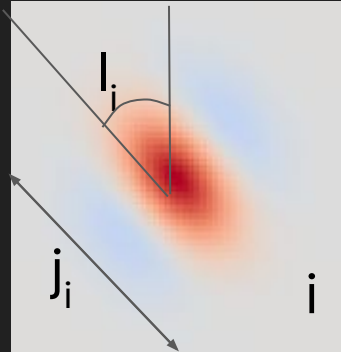


Wavelet kernels are localised and extract features



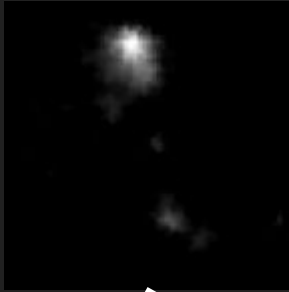
Morlet wavelets are sinusoids with Gaussian envelopes

$$\psi_{j,\ell}(u) = 2^{-2j} \psi(2^{-j} r_\theta u) \quad \text{for } 0 \leq \ell < L, \theta = \ell\pi/L$$

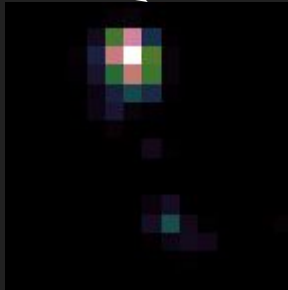


Scattering transform is a cascade of wavelet transforms

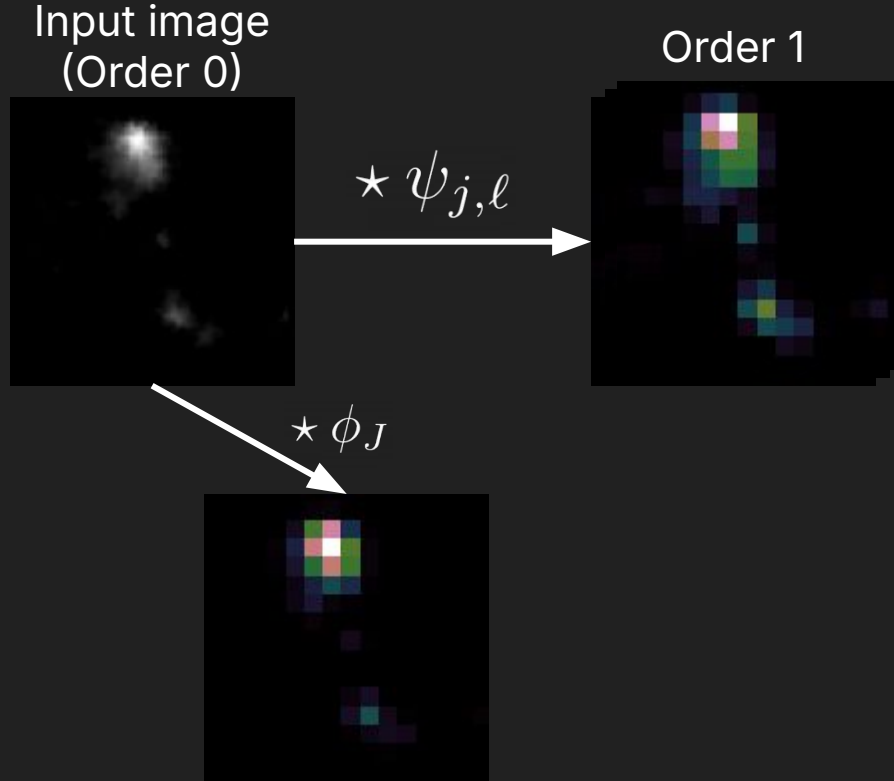
Input image
(Order 0)



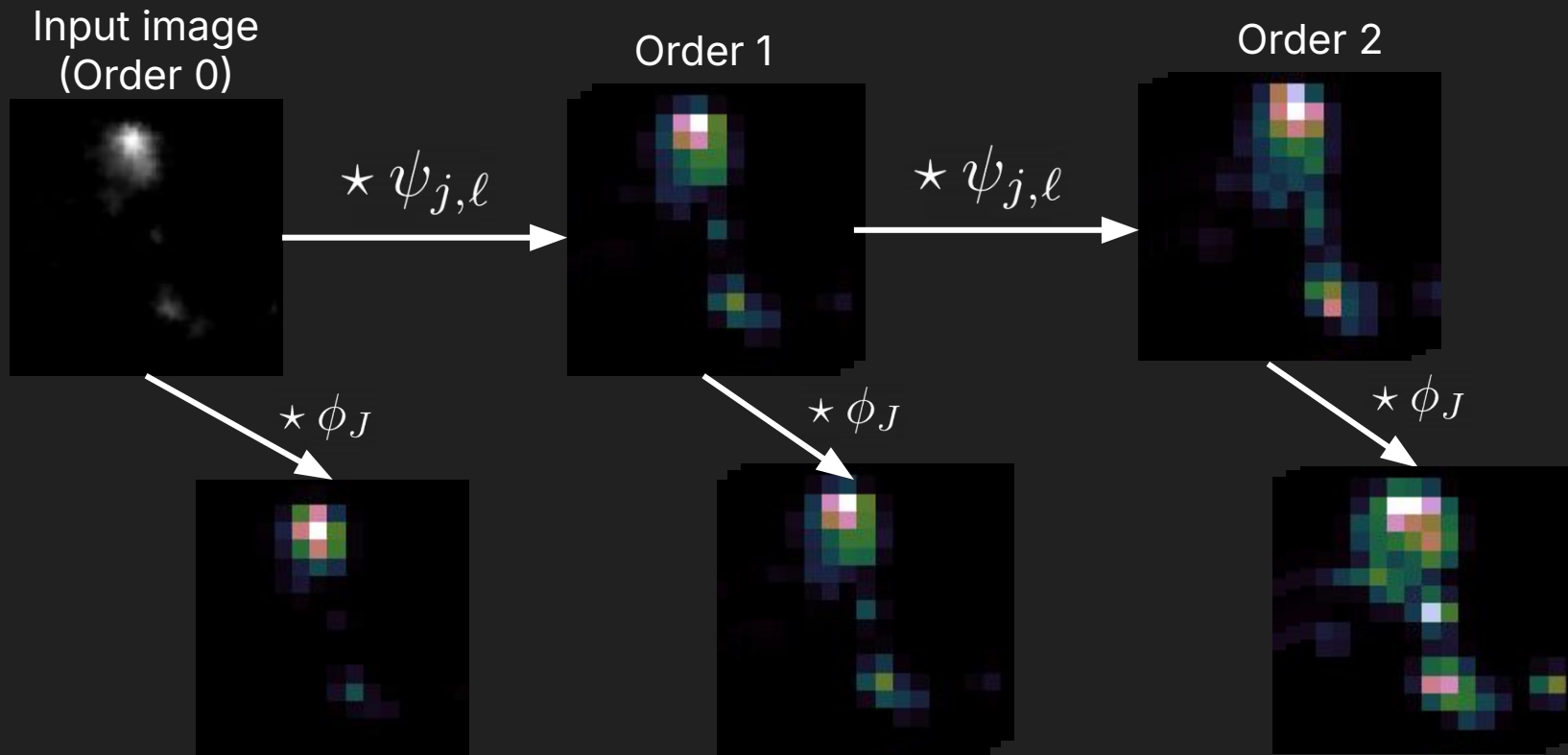
$\star \phi_J$



Scattering transform is a cascade of wavelet transforms



Scattering transform is a cascade of wavelet transforms



Scattering transforms are iterative wavelet transforms

$$\Psi_{J,L}(x) = [X, |X \star \psi_{j,\ell}|, ||X \star \psi_{j,\ell}| \star \psi_{j',\ell'}|]_{1 \leq \ell, \ell' \leq L, 1 \leq j < j' \leq J}$$

$$\phi_J(u) = 2^{-2J} \phi(2^{-J} u)$$

$$\begin{aligned} S_{J,L}(X) &= \Psi_{J,L}(X) \star \phi_J \\ &= [X \star \phi_J, |X \star \psi_{j,\ell}| \star \phi_J, ||X \star \psi_{j,\ell}| \star \psi_{j',\ell'}| \star \phi_J]_{1 \leq j < j' \leq J, 1 \leq \ell, \ell' \leq L} \end{aligned}$$

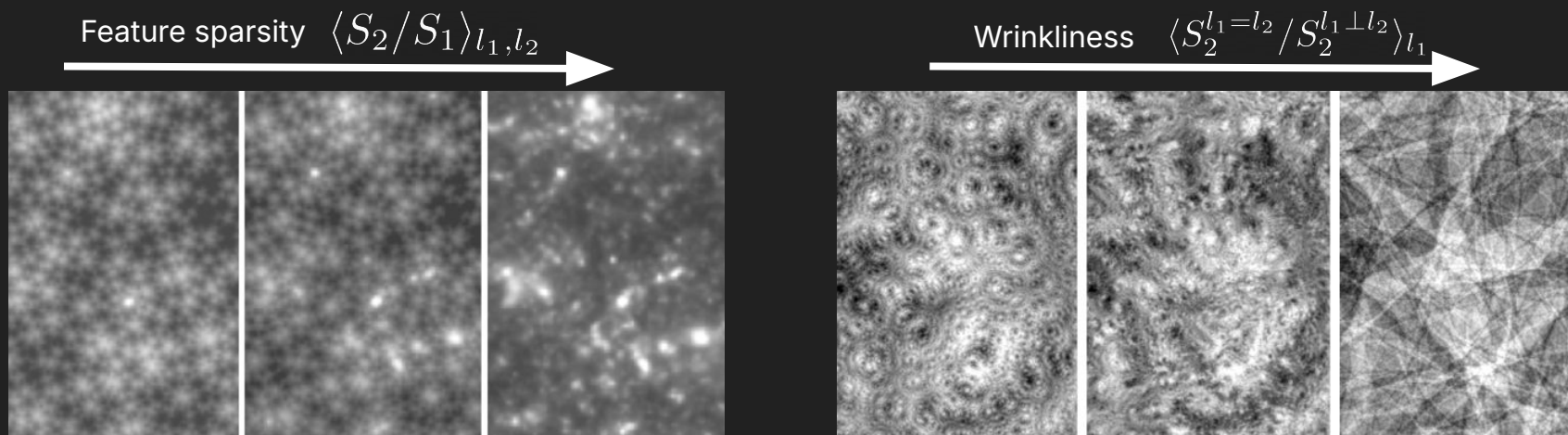
$$X \star \phi_2 \quad |X \star \psi_{0,0}| \star \phi_2 \quad |X \star \psi_{0,1}| \star \phi_2 \quad |X \star \psi_{1,0}| \star \phi_2 \quad |X \star \psi_{1,1}| \star \phi_2 \quad ||X \star \psi_{0,0}| \star \psi_{1,0}| \star \phi_2 \quad ||X \star \psi_{0,0}| \star \psi_{1,1}| \star \phi_2 \quad ||X \star \psi_{0,1}| \star \psi_{1,0}| \star \phi_2 \quad ||X \star \psi_{0,1}| \star \psi_{1,1}| \star \phi_2$$



The scattering transform is a simplified CNN or extended power spectrum

	<u>CNN</u>	<u>Power Spectrum</u>	<u>Scattering Transform</u>
Convolutions	Feature convolutions	Filters	Localised kernels
Non-linear function	Activation function	Modulus squared	Modulus
Average	Pooling	Global average	Global average
Repetition	Multi-layer	Single layer	Iterations

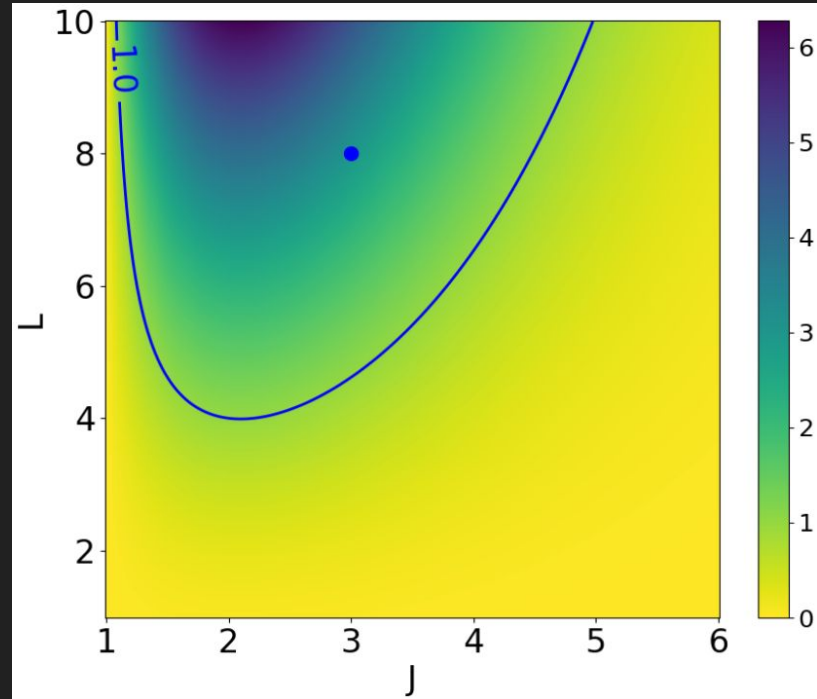
The scattering transform extracts features



Cheng & Ménard, 2021, arXiv:2112.01288v1

The scattering transform could be invertible

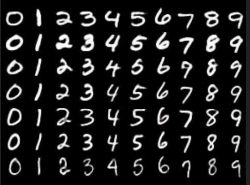
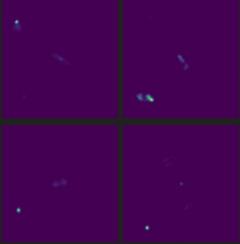
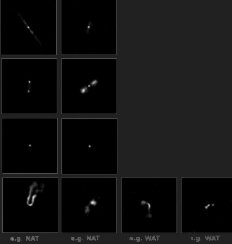
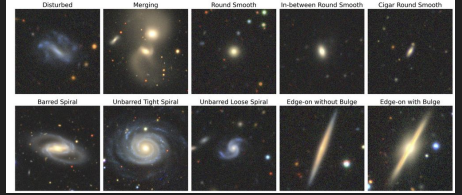
$$\frac{n_{\text{ST}}}{n_{\text{pixels}}} = 2^{-2J} \left(1 + LJ + L^2 \frac{J(J-1)}{2} \right)$$



Generation


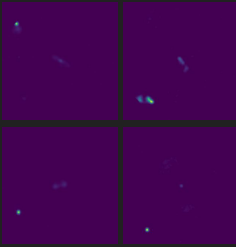
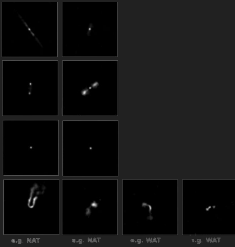
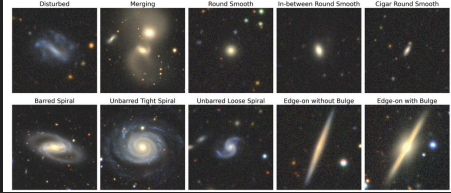
Shallow and deep image generation

Datasets used for this work

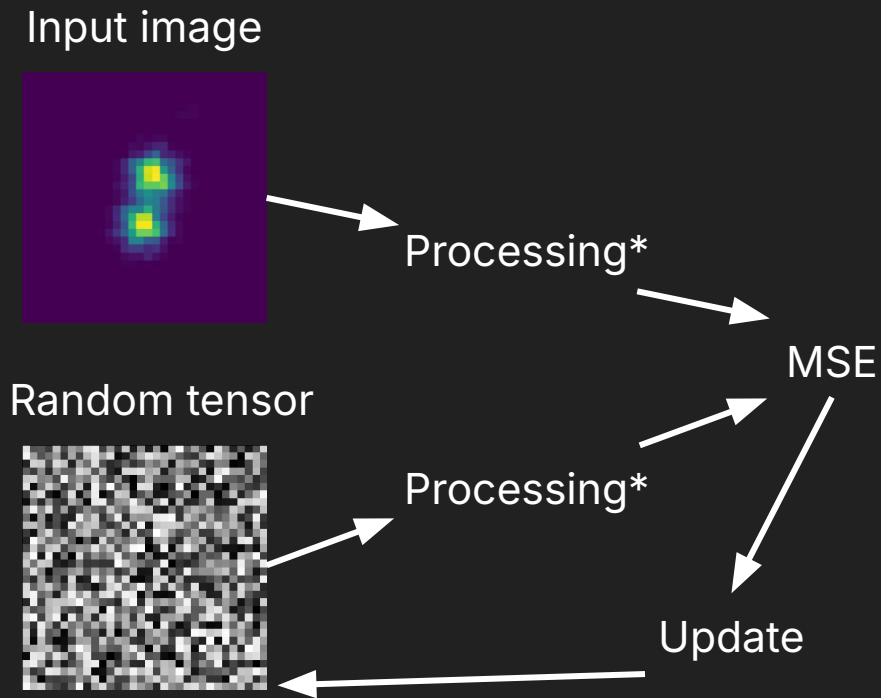
Input data	MNIST	Mirabest	FIRST	Galaxy10
Image size	28x28	150x150	300x300	3x256x256
Number of samples per class	6000	~500	~500	~1000
Example images				
# of classes	10	2	4	10
Source	Deng 2012, doi.org/10.1109/MSP.2012.2211477	Porter & Scaife 2023, arXiv:2305.11108 v1	Griese et al. 2023, doi.org/10.1016/j.ib.2023.108974	Lintott et al. 2011, arXiv:1007.3265

Datasets used for this work

For generation:

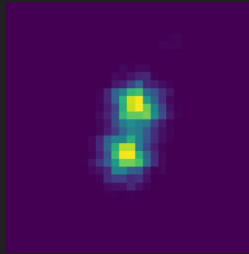
Input data	MNIST	Mirabest	FIRST	Galaxy10
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Example images				
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Reconstruction from scattering coefficients without DL



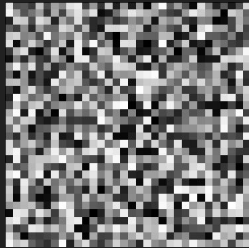
Reconstruction from scattering coefficients without DL

Input image



Processing*

Random tensor



Processing*

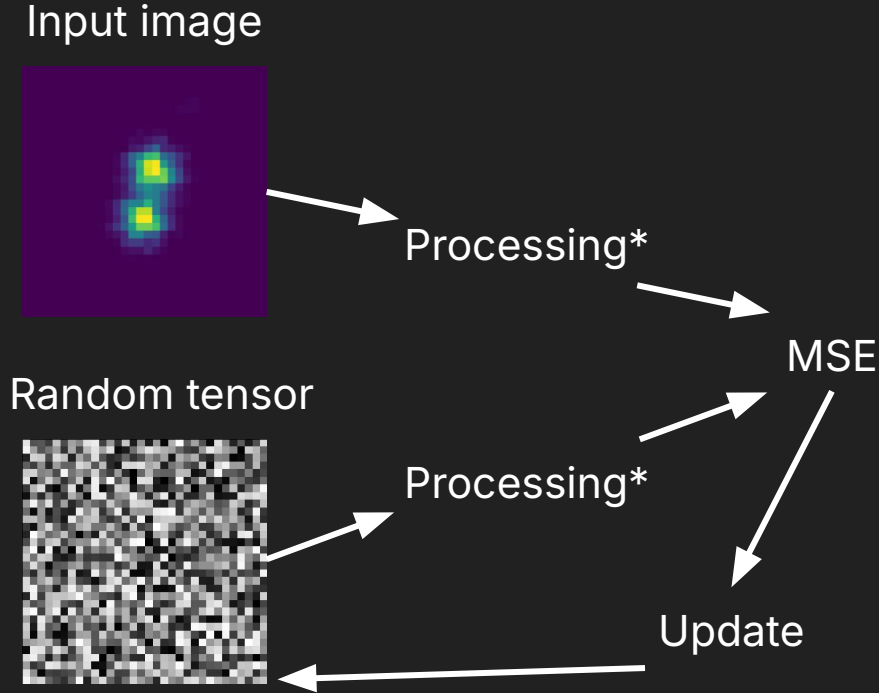
MSE

Update

* Processing:

1. Pixels: no processing
2. Full: scattering transform (ST)
3. lavg: ST + average across direction
4. Idiff: ST + subtract adjacent directions
5. javg: ST + average across scale

Reconstruction from scattering coefficients without DL



* Processing:

1. Pixels: no processing
2. Full: scattering transform (ST)
3. lavg: ST + average across direction
4. Idiff: ST + subtract adjacent directions
5. javg: ST + average across scale

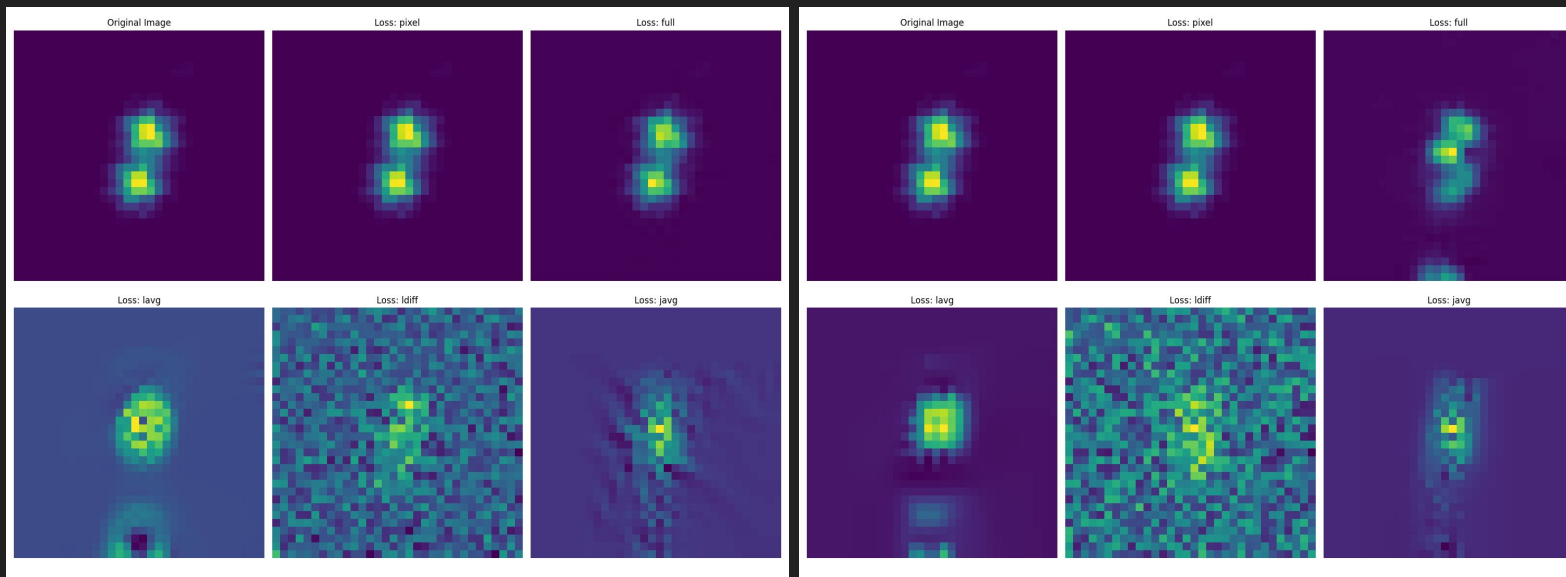


- PCA
- Average across space
- Average across space + direction
- Average across adjacent directions
- ...

Numerical reconstruction is poor

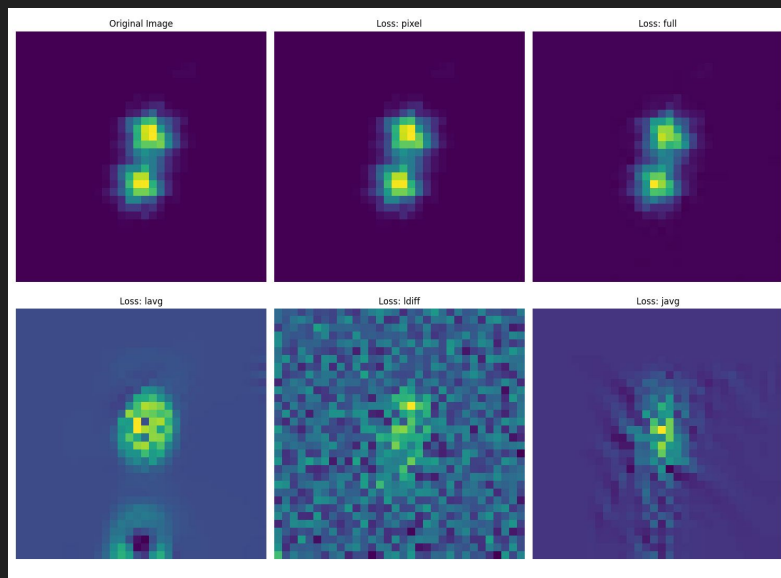
$J=3, L=8, 15,000$ iterations, $T=6h$

$J=3, L=2, 15,000$ iterations, $T=0.5h$

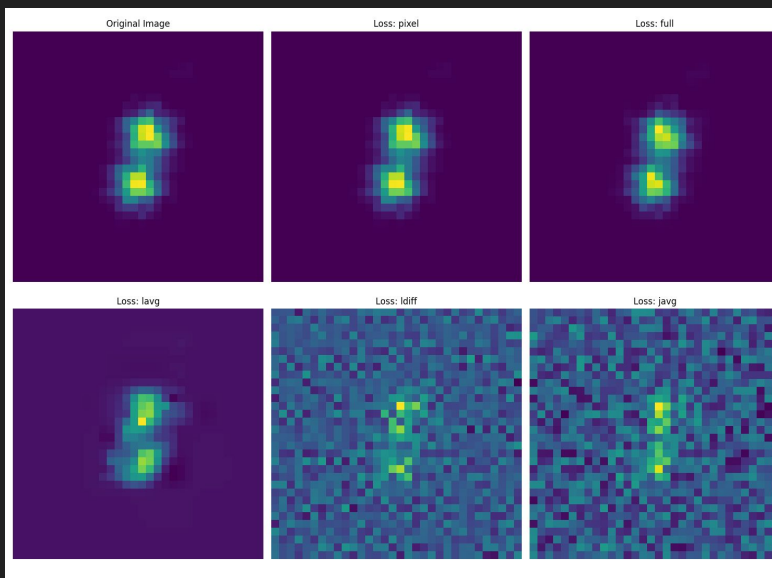


Numerical reconstruction is poor

$J=3, L=8, 15,000$ iterations, $T=6h$



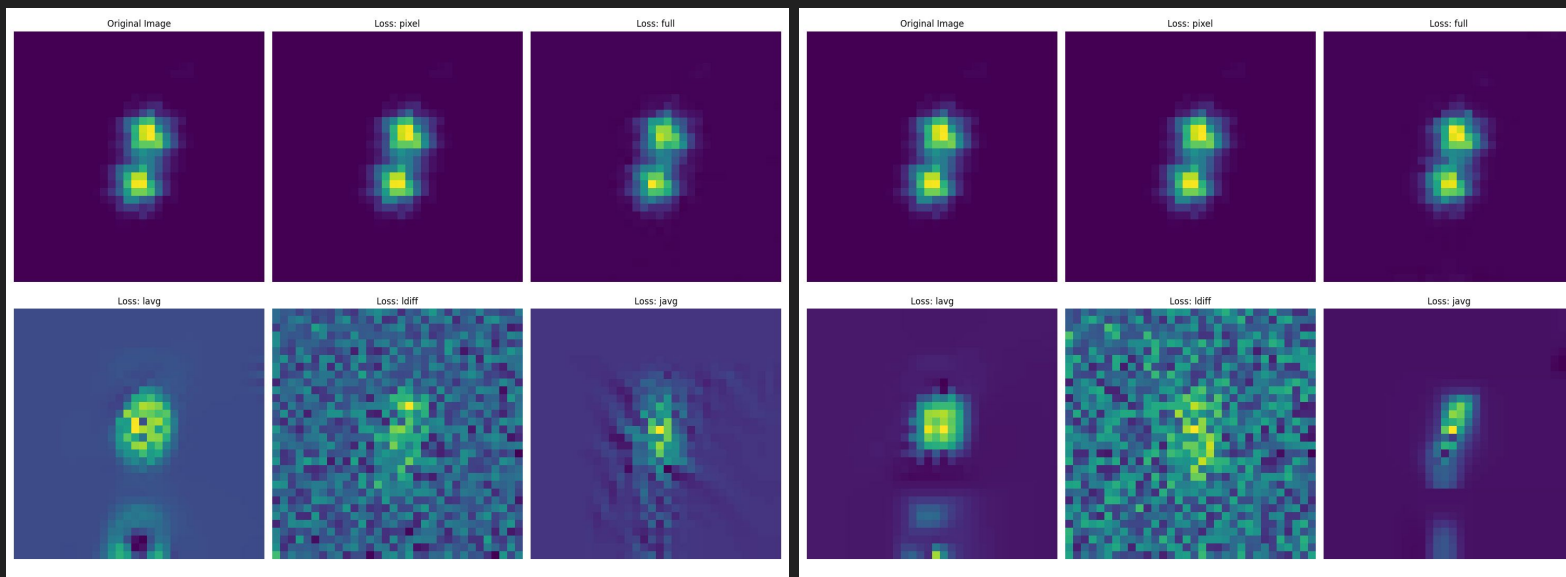
$J=3, L=2, 50,000$ iterations, $T=2h$



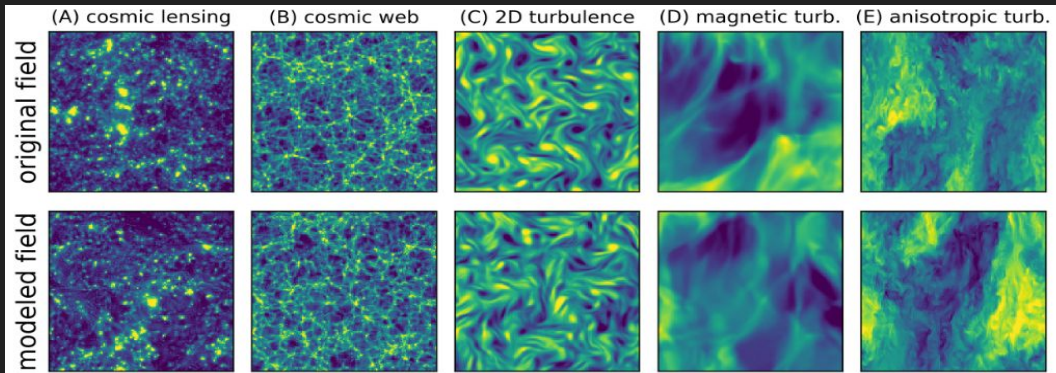
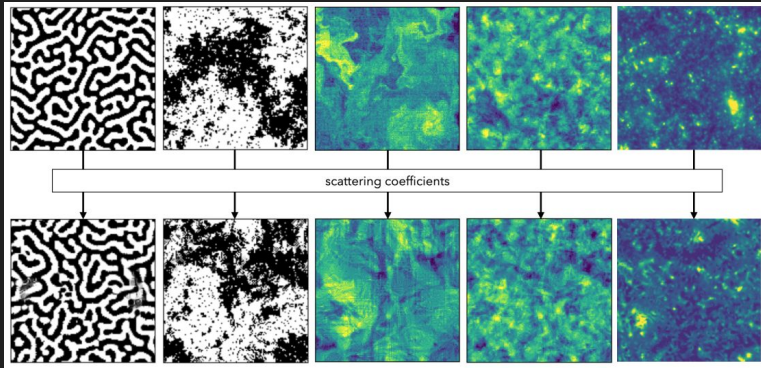
Numerical reconstruction is poor

$J=3, L=8, 15,000$ iterations, $T=6h$

$J=3, L=2, 100,000$ iterations, $T=4h$

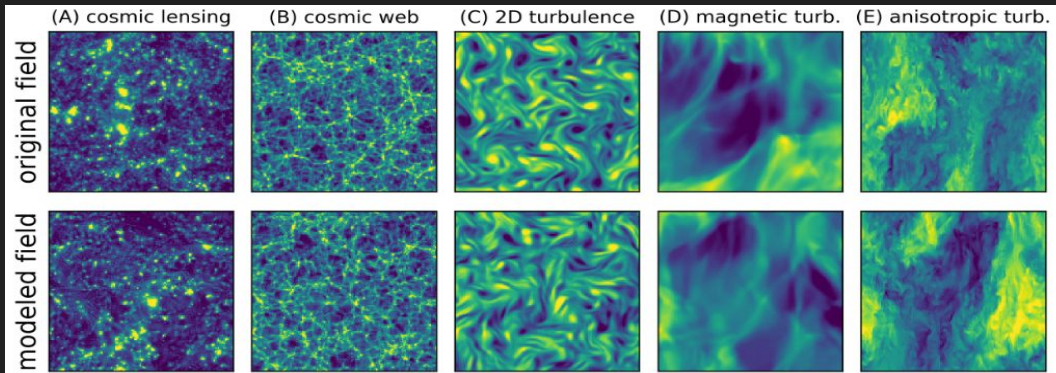
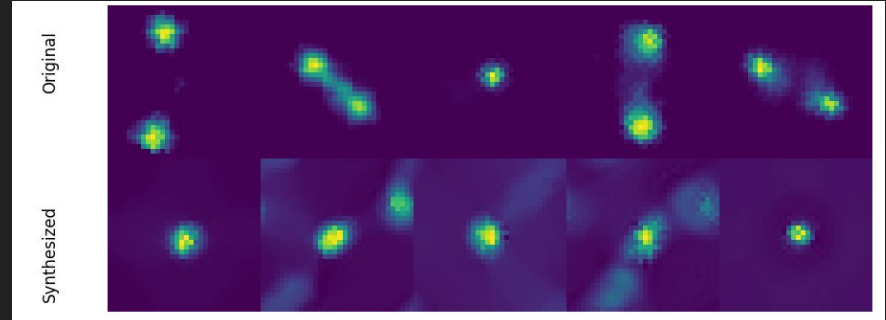
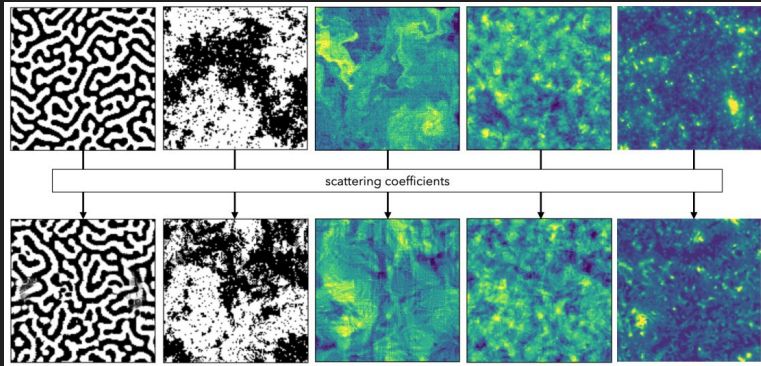


The scattering transform reproduces fields and textures



Cheng & Ménard, 2021,
arXiv:2112.01288v1

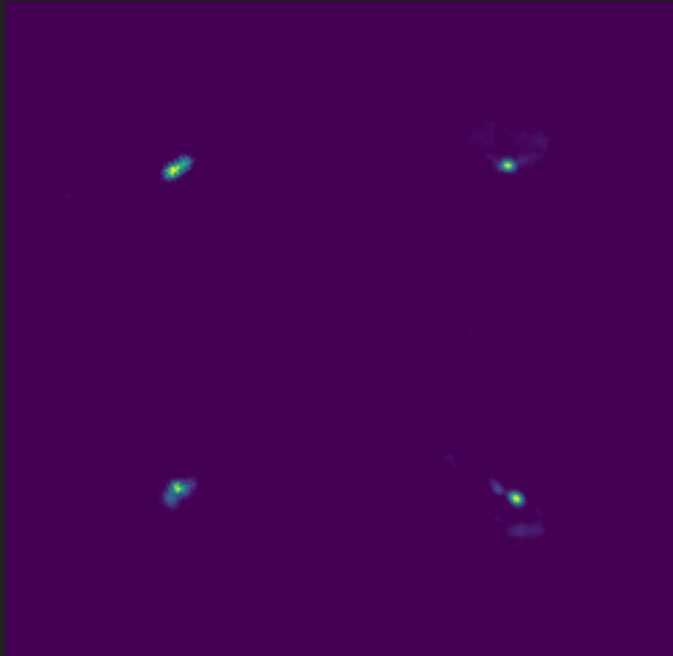
The scattering transform reproduces fields and textures but not sparse sources



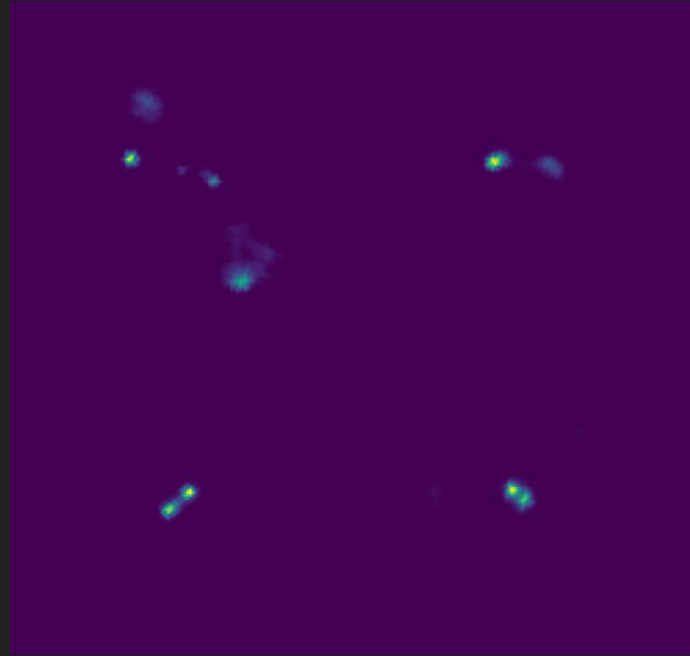
Cheng & Ménard, 2021,
arXiv:2112.01288v1

Dimensional correlations demand variable mixing

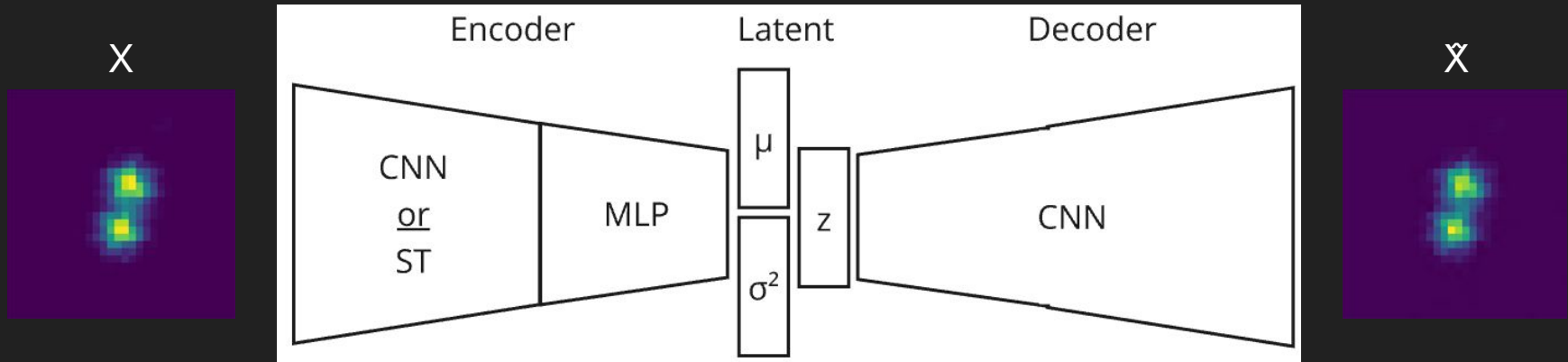
FRI



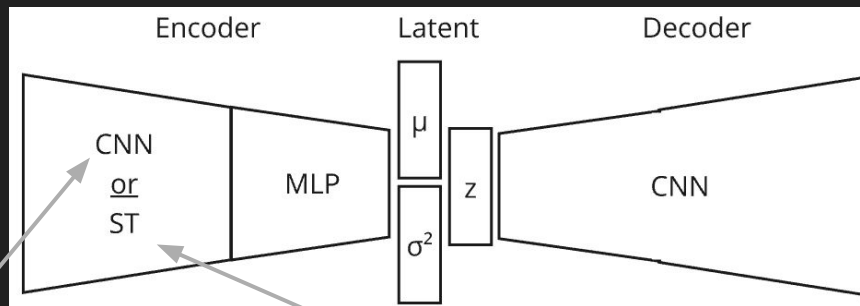
FRII



Variational autoencoders sample features and construct images



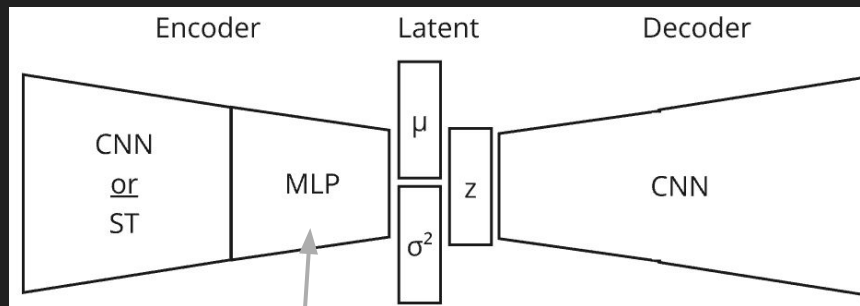
The encoder learns spatial features, by first extracting shapes



$$[X \star \phi_J, |X \star \psi_{j,e}| \star \phi_J, ||X \star \psi_{j,e}| \star \psi_{j',e'}| \star \phi_J]_{1 \leq j < j' \leq J, 1 \leq e, e' \leq L}$$

Layer	Component	Depth	Activation	Regulariser	Pooling	Parameters
1-3	5x5 Convolution	32	Leaky ReLU	Batch norm.	2x2 Max pooling	19,520
4-6	5x5 Convolution	64	Leaky ReLU	Batch norm.	2x2 Max pooling	125,504
7-9	5x5 Convolution	128	Leaky ReLU	Batch norm.	2x2 Max pooling	353,244
10-12	5x5 Convolution	256	Leaky ReLU	Batch norm.	2x2 Max pooling	2,001,152
13-15	5x5 Convolution	256	Leaky ReLU	Batch norm.	2x2 Max pooling	2,820,352
						5,319,772

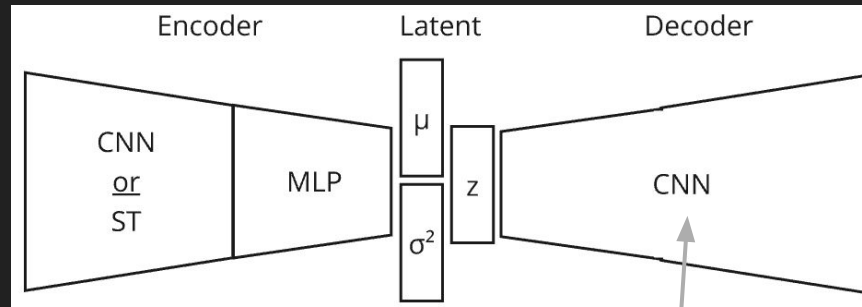
The encoder learns spatial features, by first extracting shapes, then learning their dependencies



Layer type	Kernel size	Stride	Padding	Depth	Activation	Regulariser	Pooling	Parameters
Convolutional	3x3	2	1	256	Leaky ReLU	Batch norm.	2x2 Max pooling	590,592
Convolutional	3x3	2	1	384	Leaky ReLU	Batch norm.	2x2 Max pooling	885,888
Convolutional	2x2	-	0	384	-	-	-	590,208
								2,066,688

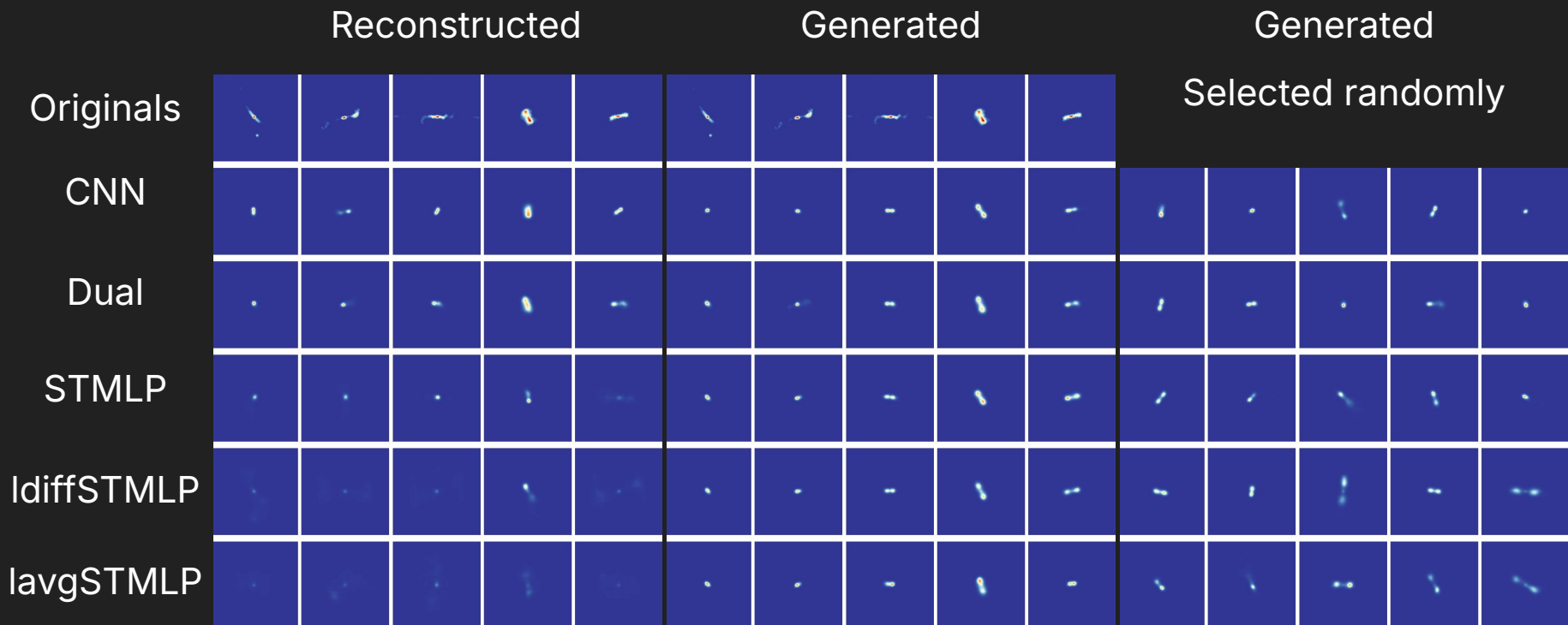
Layer type	Depth	Activation	Regulariser	Parameters
Linear	384	Leaky ReLU	Layer norm.	148,608
Linear	256	Leaky ReLU	Layer norm.	99,072
Linear	2*32	-	-	25,700
				273,380

The encoder learns spatial features, by first extracting shapes, then learning their dependencies, and then upscaling

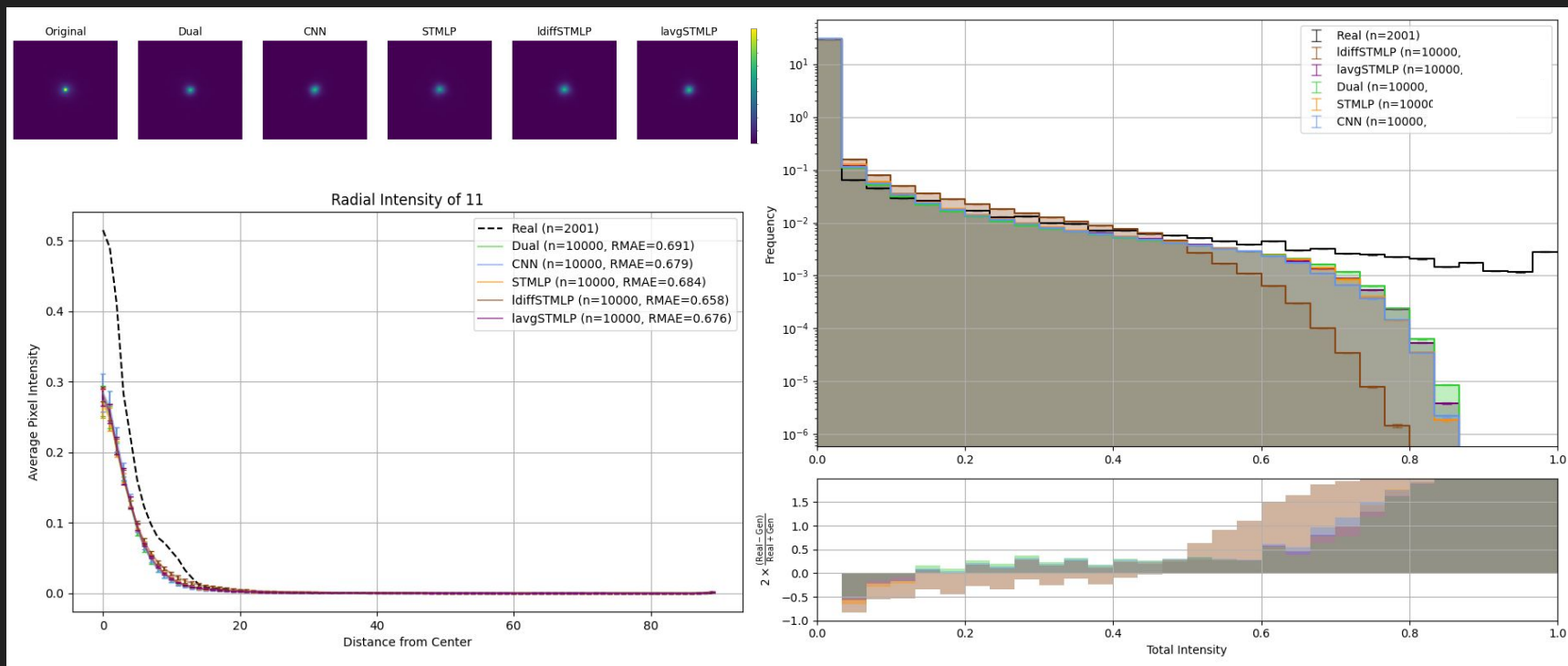


Layer	Component	Depth	Activation	Regulariser	Parameters
1	Linear	2048	Leaky ReLU	-	67,584
2-4	4x4 Convolutional transpose	64	Leaky ReLU	Batch norm.	205,376
5-7	4x4 Convolutional transpose	128	Leaky ReLU	Batch norm.	427,136
8-10	4x4 Convolutional transpose	64	Leaky ReLU	Batch norm.	205,376
11-13	4x4 Convolutional transpose	32	Leaky ReLU	Batch norm.	51,488
14	4x4 Convolutional transpose	1	Sigmoid	-	513
					957,473

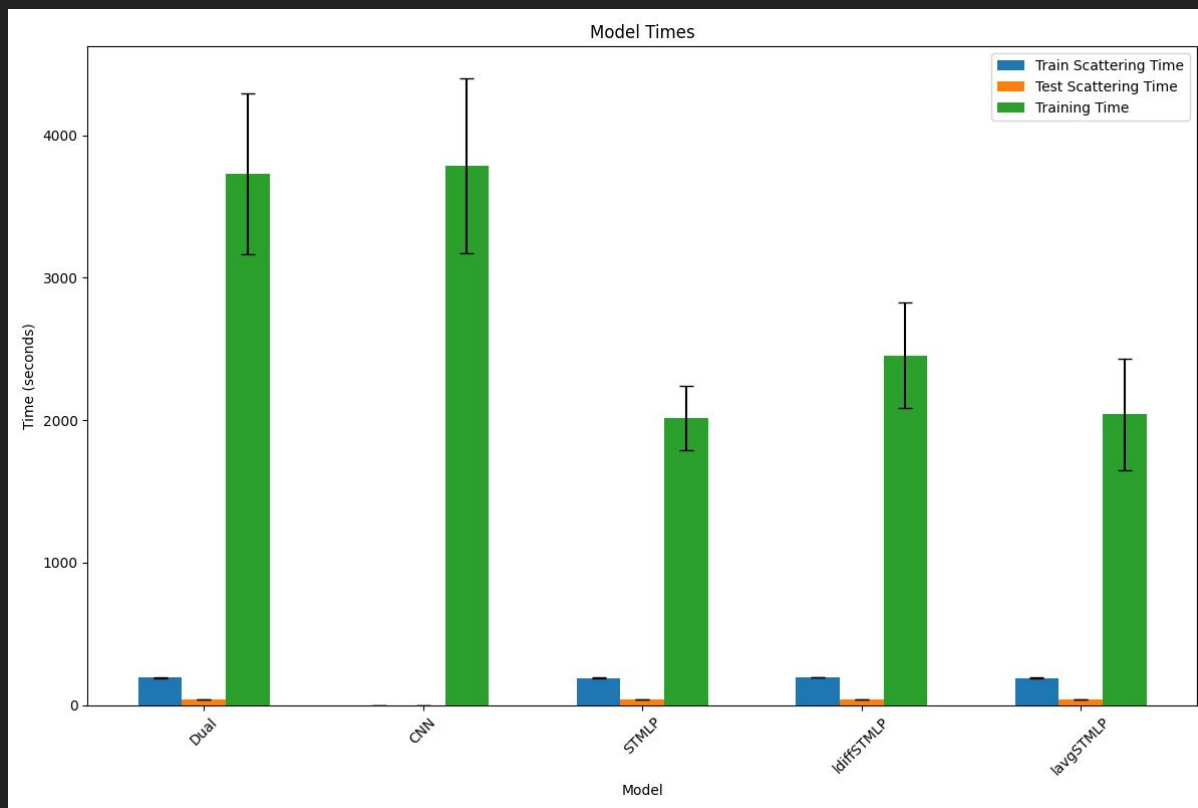
Model reproduces smoother sources than the originals



Model reproduces fainter sources than the originals



But the scattering transform is faster



Conclusions and outlook

Conclusions

1. The scattering transform (ST) is computationally efficient and interpretable
2. Scattering-based generative modelling of radio galaxies require multivariate learning
3. My variational autoencoders produces overly smooth and faith images

Future plans

1. More complex generative models
 - a. Generative Adversarial Networks (GANs)
 - b. Diffusion models
 - c. Normalising flows
2. Apply on diffuse cluster radio emission

Markus Bredberg

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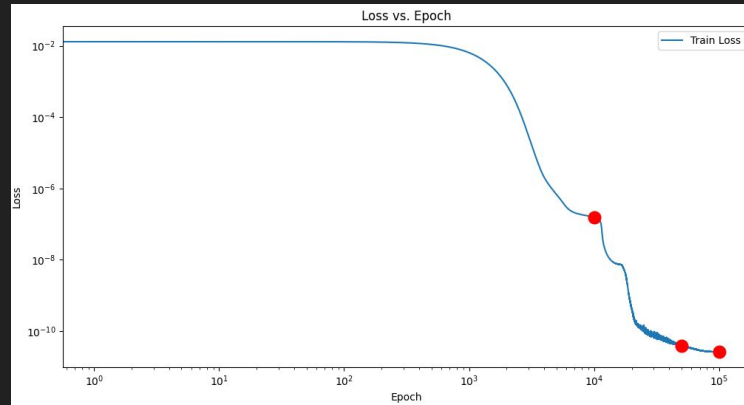
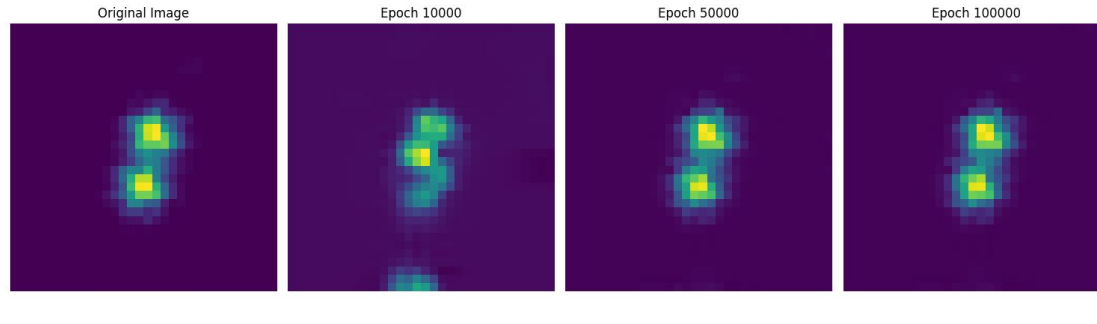
Bonus slides

Clarifications

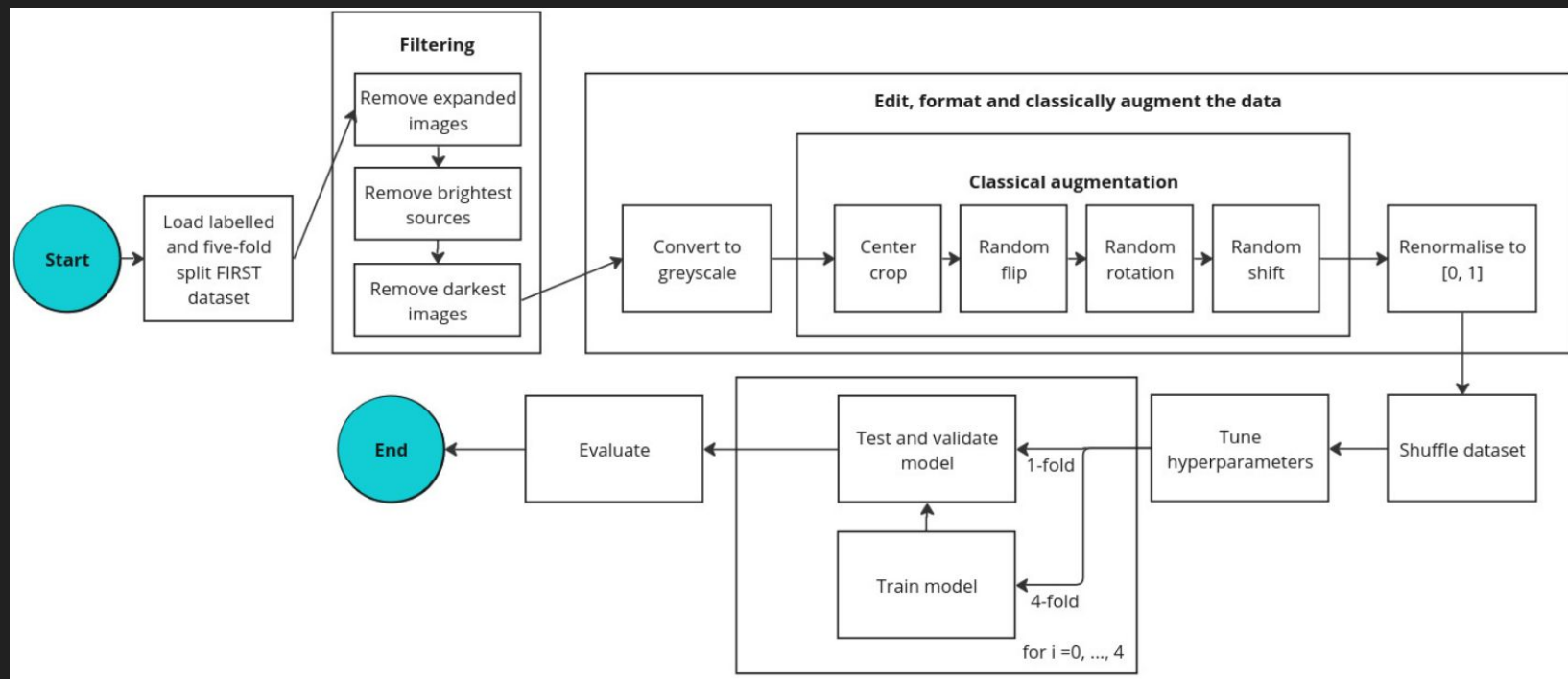
Classification for labelling and evaluation

Shallow generation in epochs

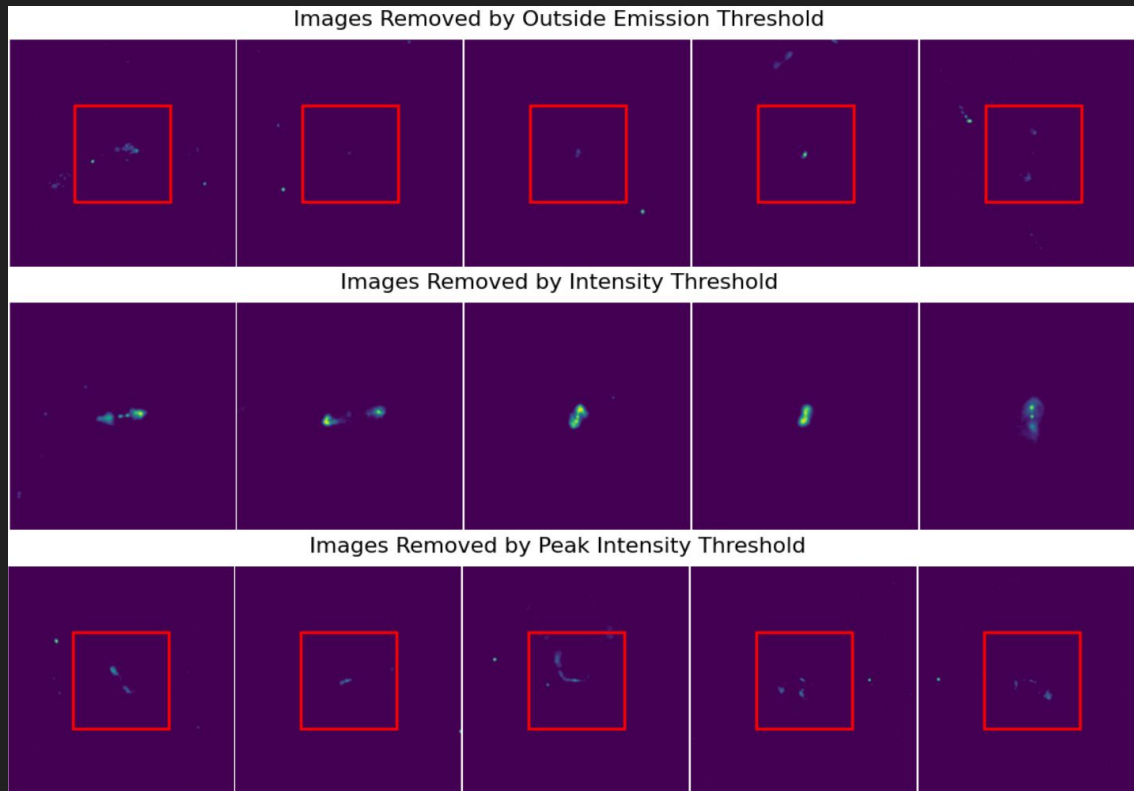
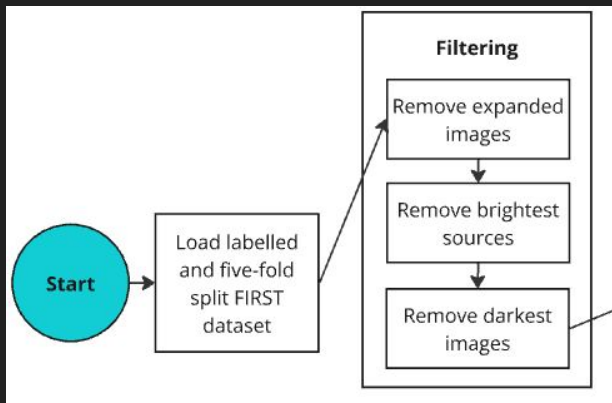
Generated image with full loss, for one galaxy of type Confidently classified FRIIs and 100000 epochs.



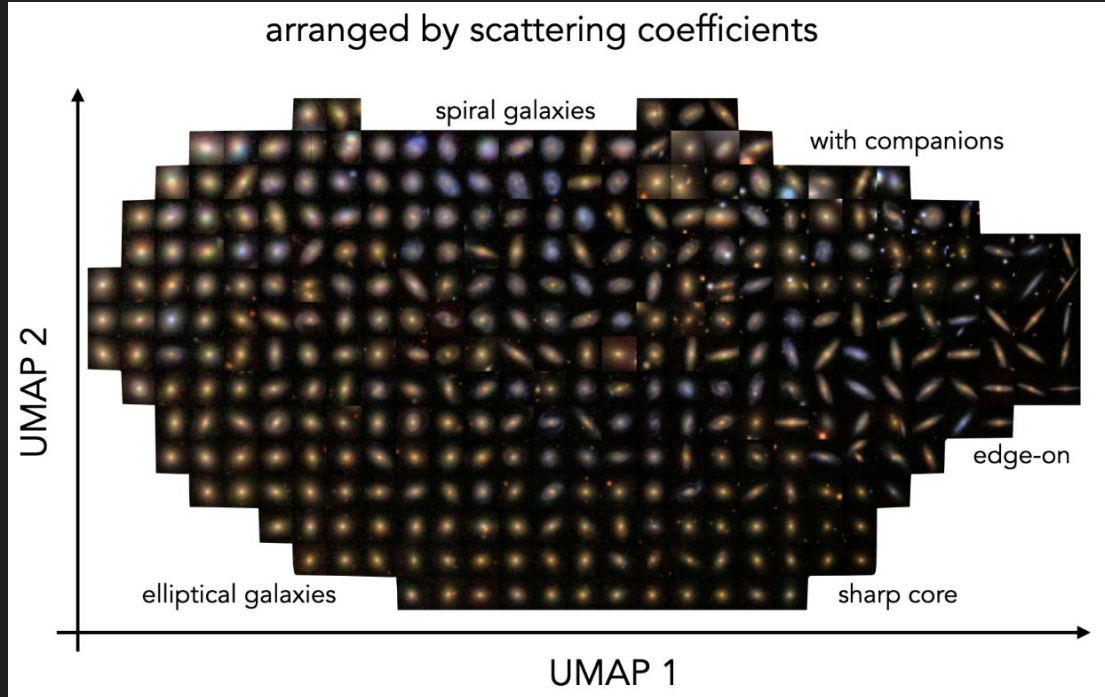
Running the generative script: the algorithm schematically



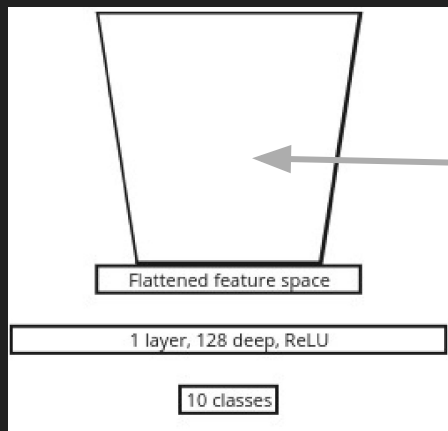
Running the generative script: the filtering



The scattering transform can categorise



Classifying with the scattering transform



CNN

Layer type	Depth	Kernel size	Activation	Pooling type	Pooling size
Conv2D	32	5x5	ReLU	MaxPool2D	2x2
Conv2D	64	5x5	ReLU	MaxPool2D	2x2

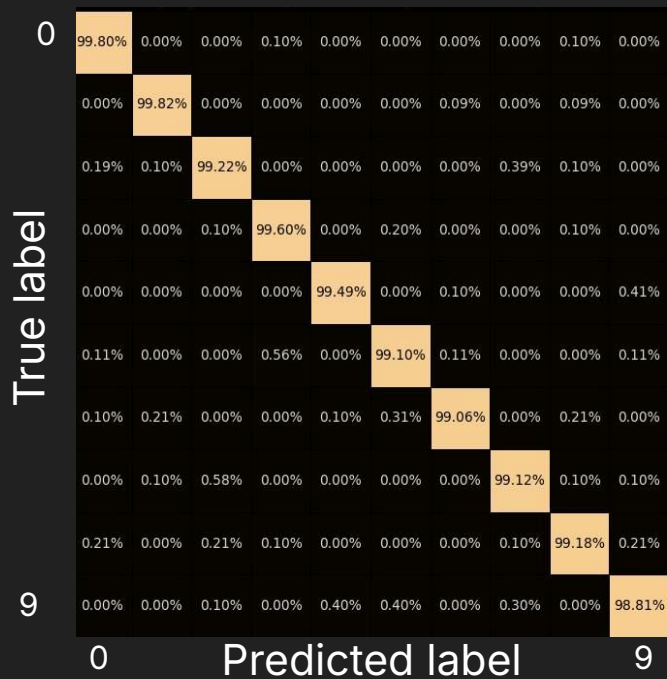
or

Scattering Transform

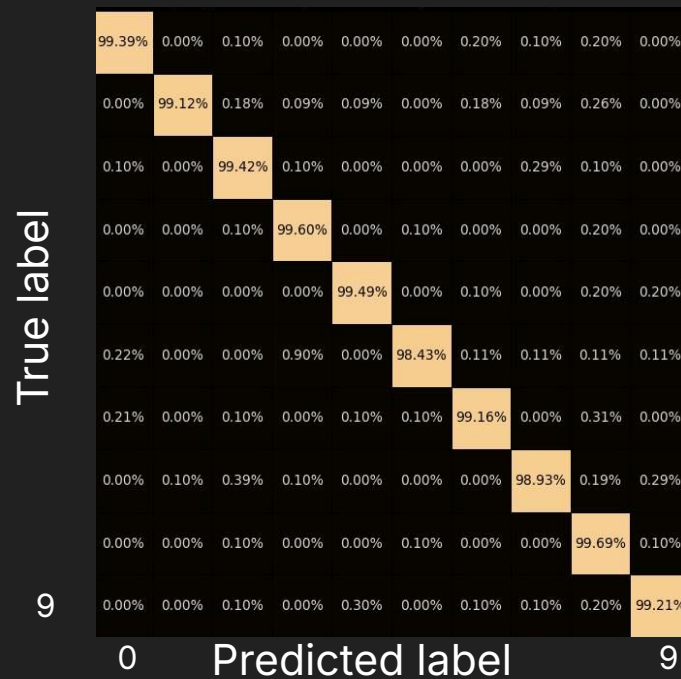
$$S_{3,8}(X) = \Psi_{3,8}(X) \star \phi_3$$

Scattering transform classification speeds up

CNN: 99.33%



Scattering Transform Network:
99:25%



Markus Bredberg, Jiaxin Guo, Han Zhang (2024)

The VAEs do not learn the feature representation

