

# HIRAX - Overview and Status

**Jennifer Studer, ETHZ Cosmology Group  
SKA Days 2024**

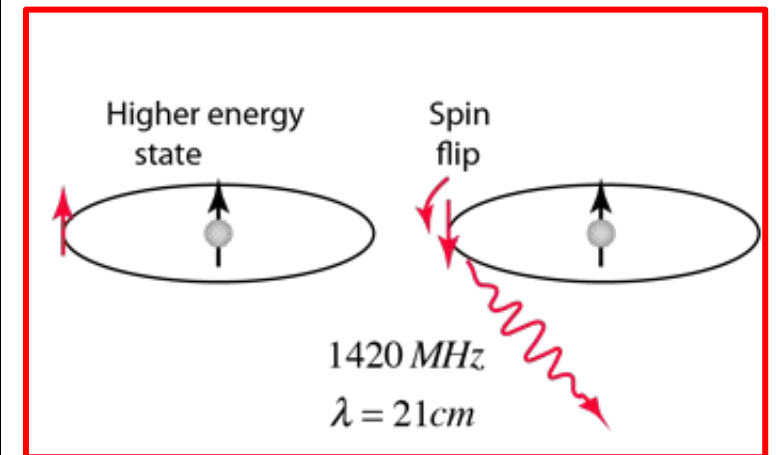
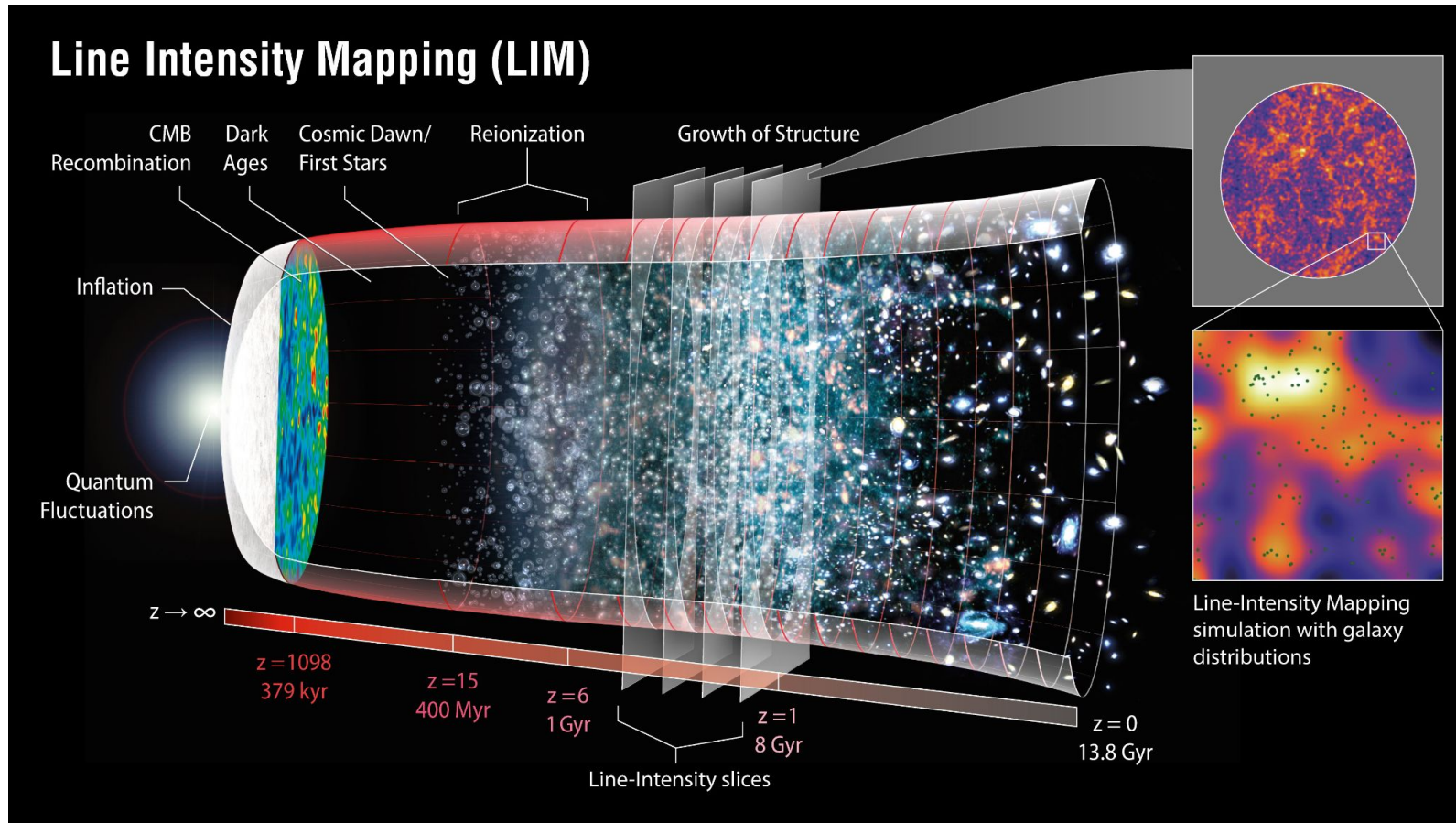
Alexandre Refregier, Devin Crichton, Thierry Viant,  
Corrie Ungerer, Kavilan Moodley



# HIRAX Overview



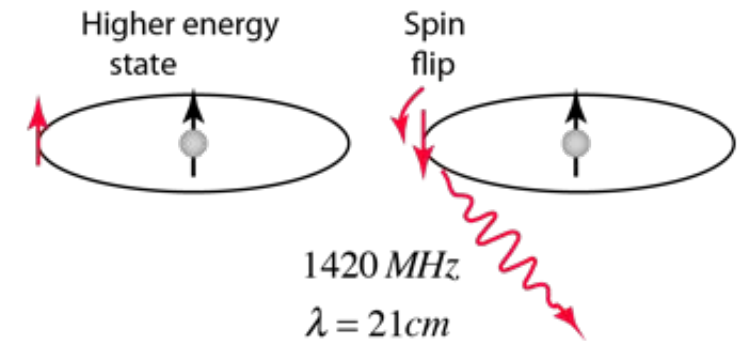
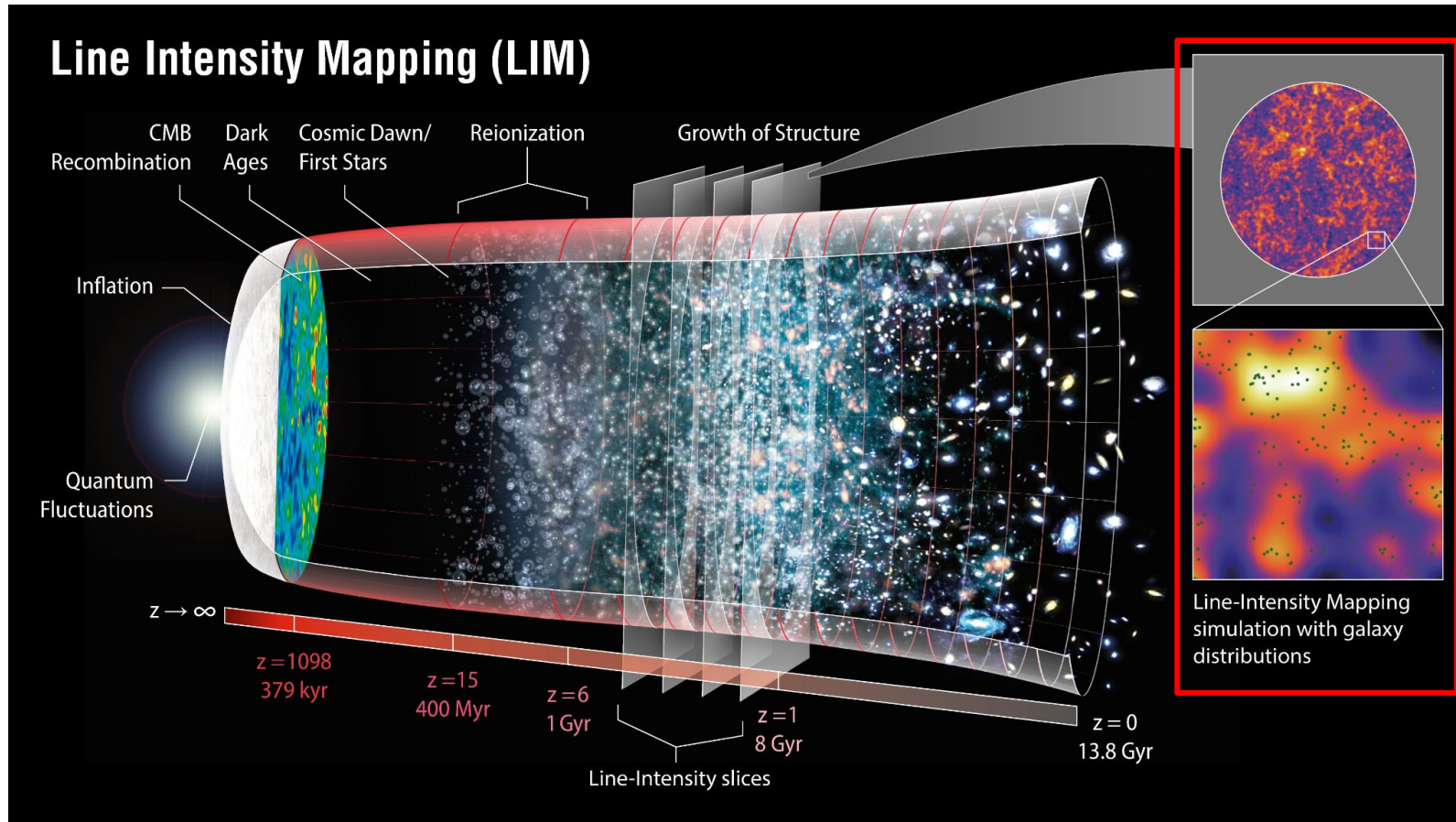
- **H**ydrogen **I**ntensity and **R**eal-time **A**nalysis e**X**periment
- Radio interferometer with a compact, redundant layout



# HIRAX Overview



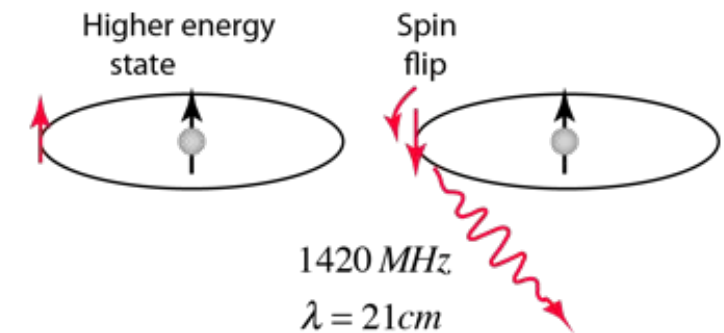
- **H**ydrogen **I**ntensity and **R**eal-time **A**nalysis e**X**periment
- Radio interferometer with a compact, redundant layout



# HIRAX Overview

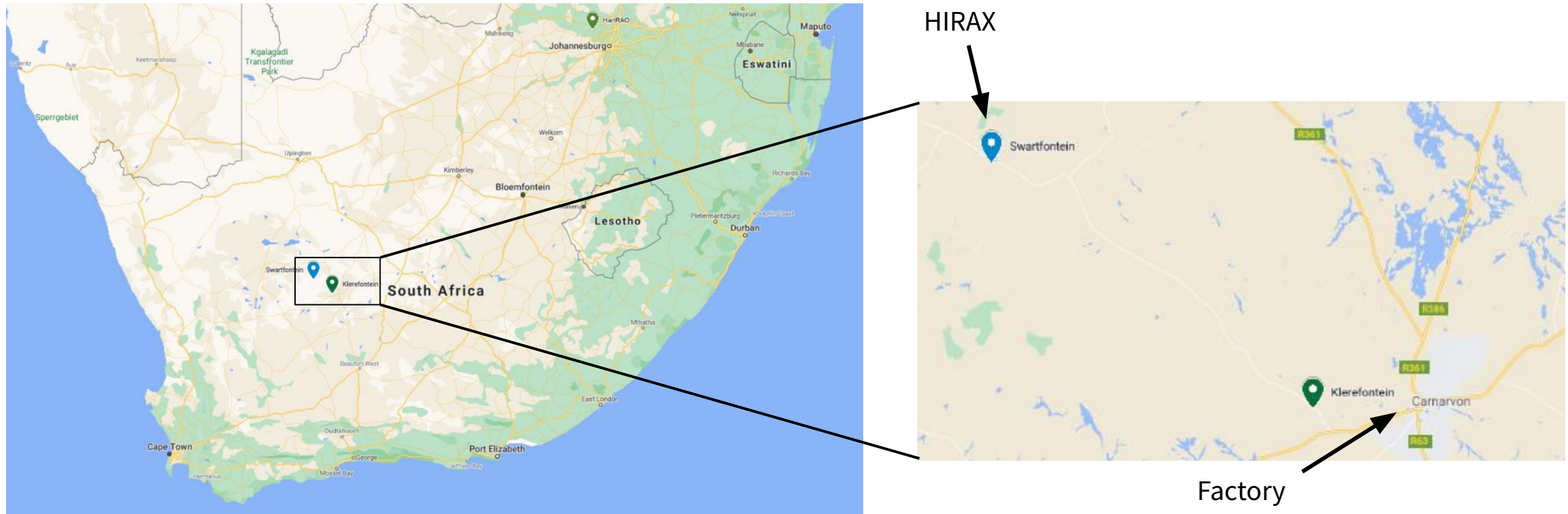


- **H**ydrogen **I**ntensity and **R**ead-time **A**nalysis e**X**periment
- Radio interferometer with a compact, redundant layout
- To be co-located with SKA in the Karoo, South Africa
- Funded up to 256 element deployment. Plans to extend to 1024.
- 6 m diameter dishes instrumented to operate between 400–800 MHz /  $z = 0.8-2.6$
- Intensity mapping survey of  $\sim 1/3$  of the sky over 4 years
- Field of view:  $5^\circ-10^\circ$
- Primary Science Goals:
  - Observationally probe the evolution of dark energy
  - Survey the transient radio sky



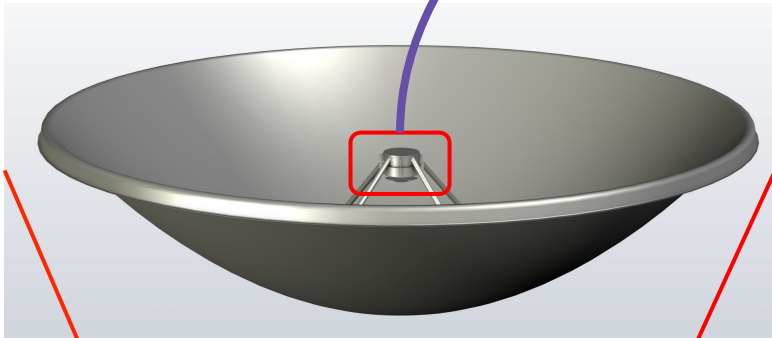
# HIRAX Site

- Guest instrument on SKA site in the Karoo, South Africa
- Low RFI (radio frequency interference) site - protected by government regulations
- Access to roads, power supply, external network connection, and SKA infrastructure

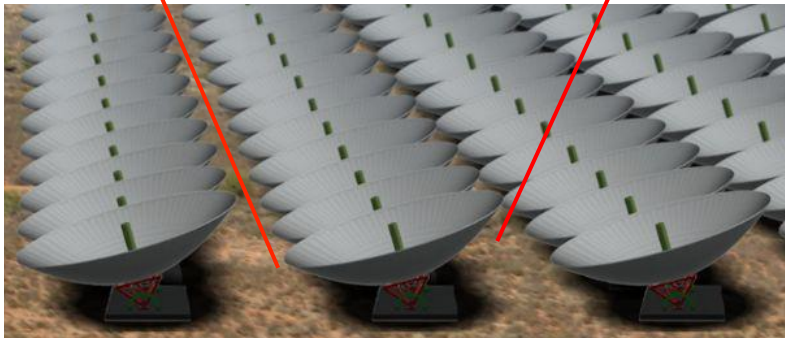


# HIRAX Schematic

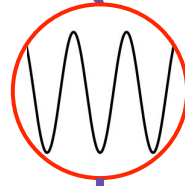
**RF Frontend**  
dual-polarized feed  
RFOF transmitter



dish diameter: 6 m  
 $f/D = 0.21$



256 antennas



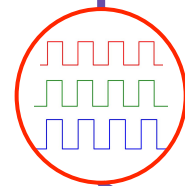
**RFOF receiver**

**F-Engine**



ICEBoard

32 boards:  
1 board per 4 antennas



**X-Engine**



Integrated Correlator System  
8 x nodes

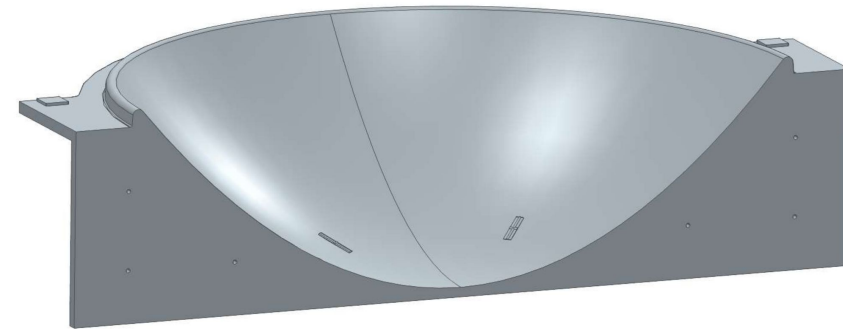
Process about 1.6  
Tb/s of raw data



# HIRAX Dish Production

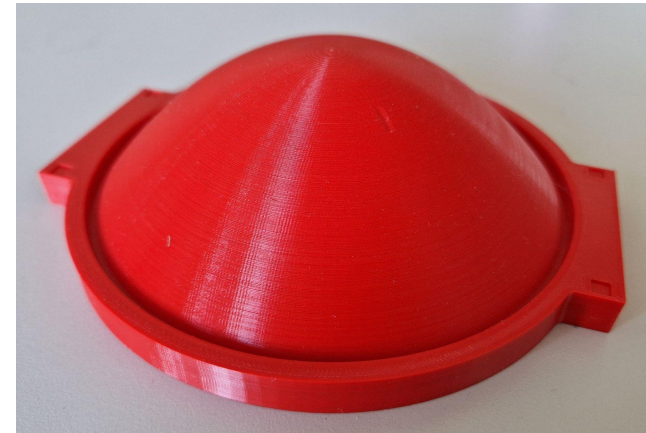
## Reflector Plug

- Manufactured in two halves
- Manufactured and measured in Cape Town
- Combined, measured and finished in Carnarvon



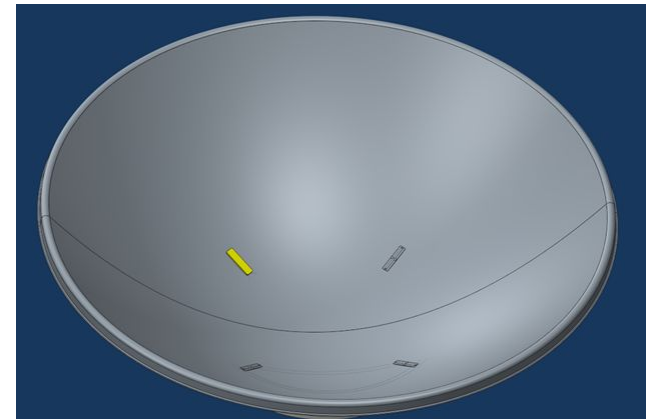
## Reflector Mold

- Half molds manufactured and measured in Cape Town
- 4 molds for main dishes

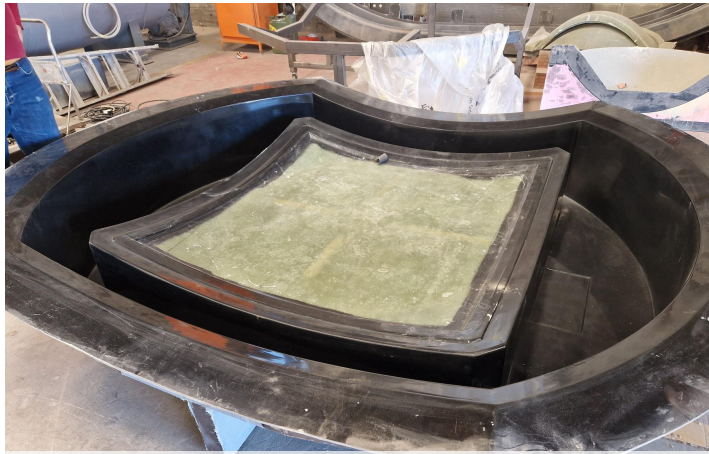


## Reflector Dish

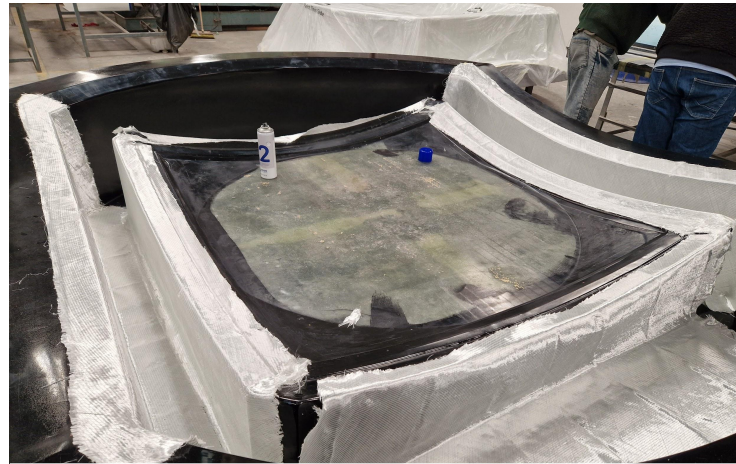
- First half dish prototype measured
- First 2 prototypes planned in October/November
- Fiberglass with an embedded aluminium mesh



# Telescope Mechanical Assembly - Fibreglass



Backing ring mould



Backing ring built on a mould



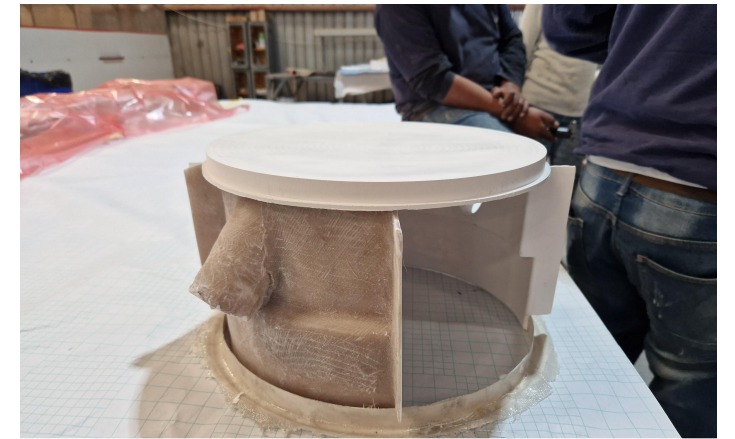
Backing ring mould infusion



Scalar model of plug and mould



Feedleg assembly jig



Feed assembly jig



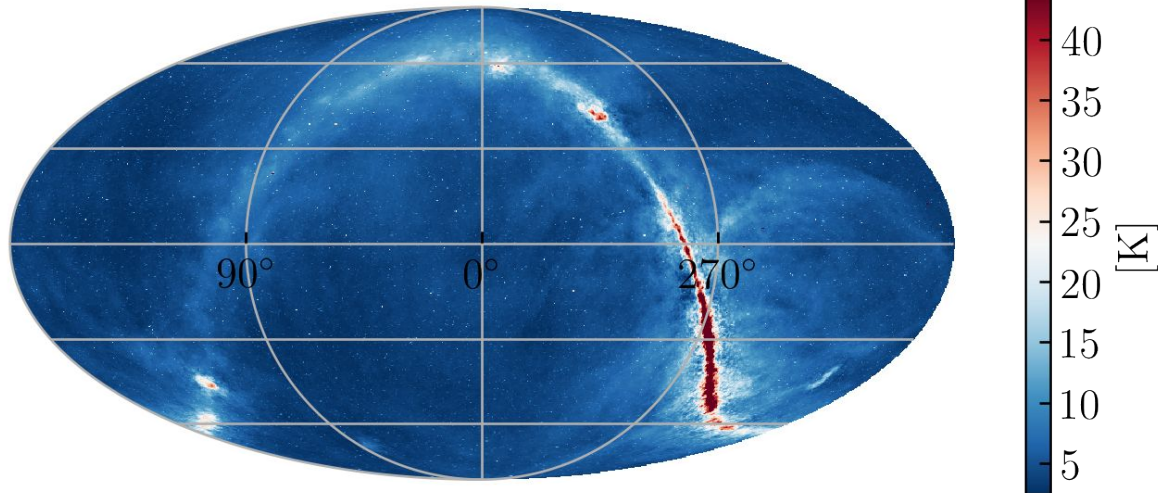
# First assembled Dish



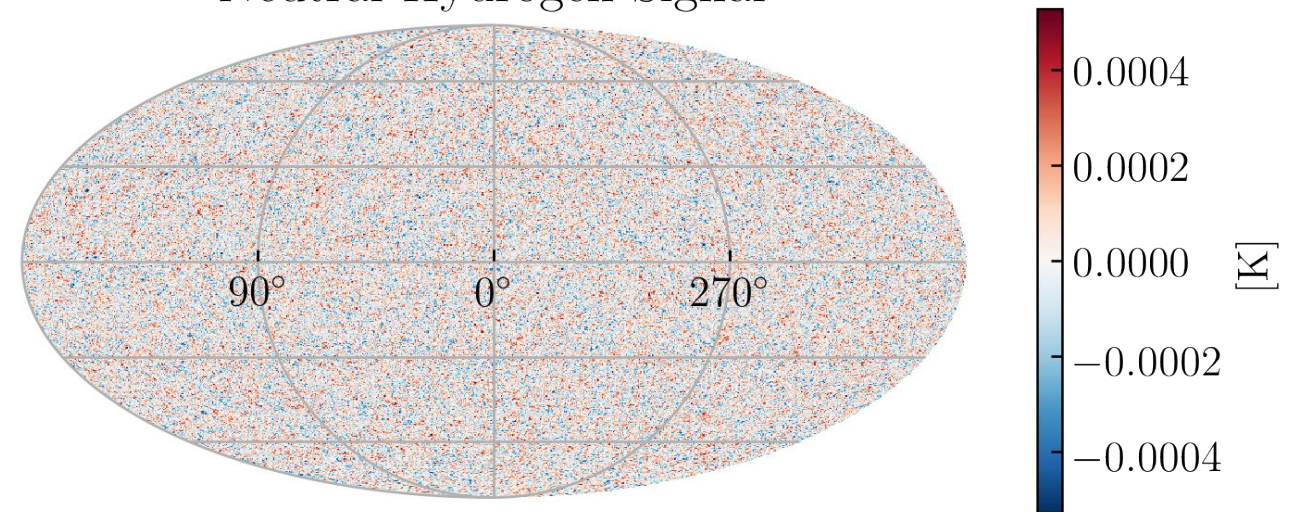
# Systematics

- Foreground highly dominates
  - Smooth in frequency
  - HI is correlated over small ranges in frequency
  - Instrument systematics are frequency dependent
- Imperfect knowledge of the instrument leads to foreground leakage
- Need instrument with very low systematics and a very good understanding of it

Galactic Foreground

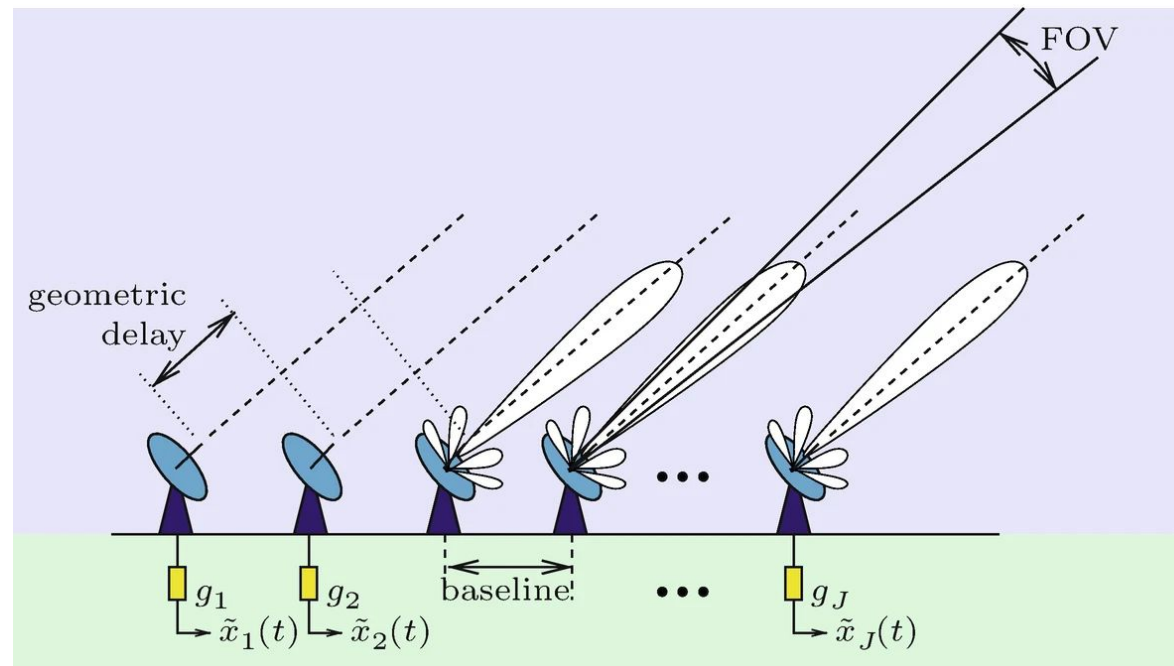


Neutral Hydrogen Signal



# Systematics

- Foreground highly dominates
  - Smooth in frequency
  - HI is correlated over small ranges in frequency
  - Instrument systematics are frequency dependent
- Imperfect knowledge of the instrument leads to foreground leakage
- Need instrument with very low systematics and a very good understanding of it



<b>Telescope mechanical parameter</b>	<b>Target precision (RMS)</b>
Receiver position relative to focus	0.5 mm
Receiver orientation relative to boresight vector	2.5' polar and azimuthal
Dish surface deviations	1 mm
Dish vertex position relative to elevation axis	1 mm
Orthogonality of boresight vector and elevation axis	1'
Elevation axis position within the array	0.5 mm in array plane 1 mm out of array plane
Elevation axis alignment within the array	1'
Elevation pointing angle	1'

**Table 4** Target precision values for HIRAX telescope mechanical structure

# Measurement Equipment to Characterize the Dish

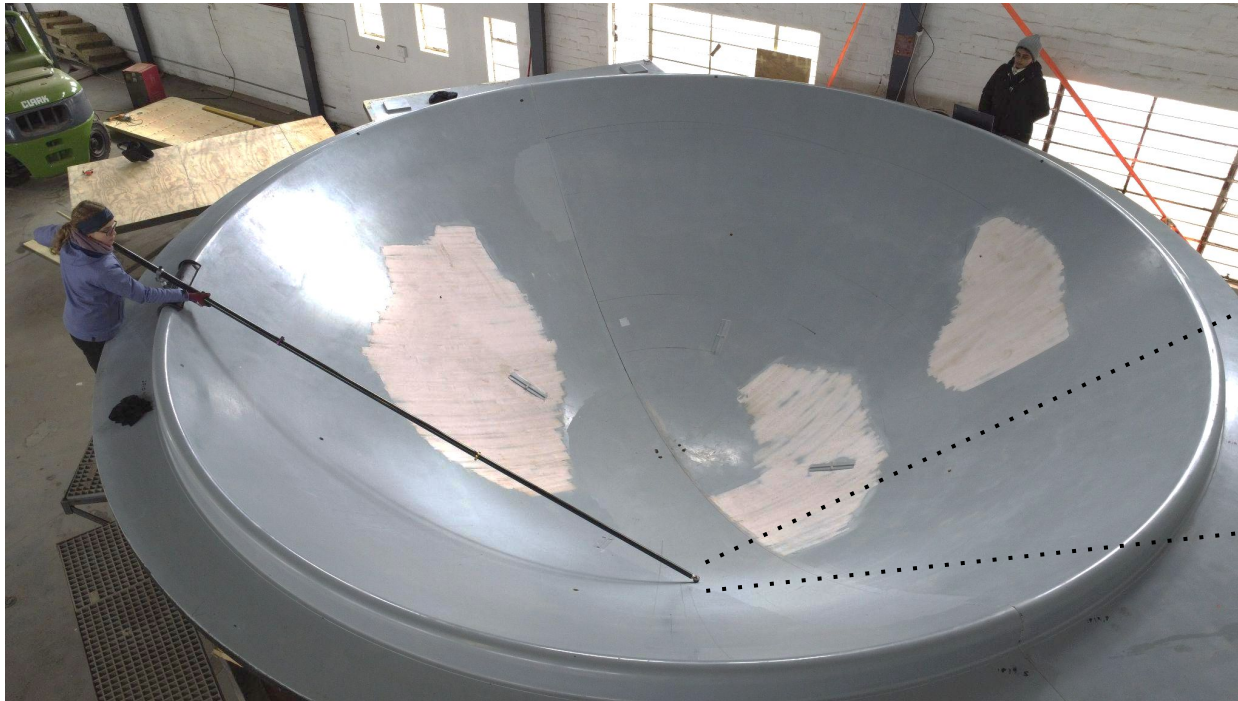
- Laser tracker
  - Very dense point cloud
  - ~ 1-2 hours/dish
- Photogrammetry system
  - Medium dense point cloud
  - < hour/dish
  - Survey over time
- Reflectometer
  - Very sparse point density
  - ~ 3 hours/dish
  - Measures the actual EM surface relative to the dish surface



# Laser Tracker

Measurement procedure:

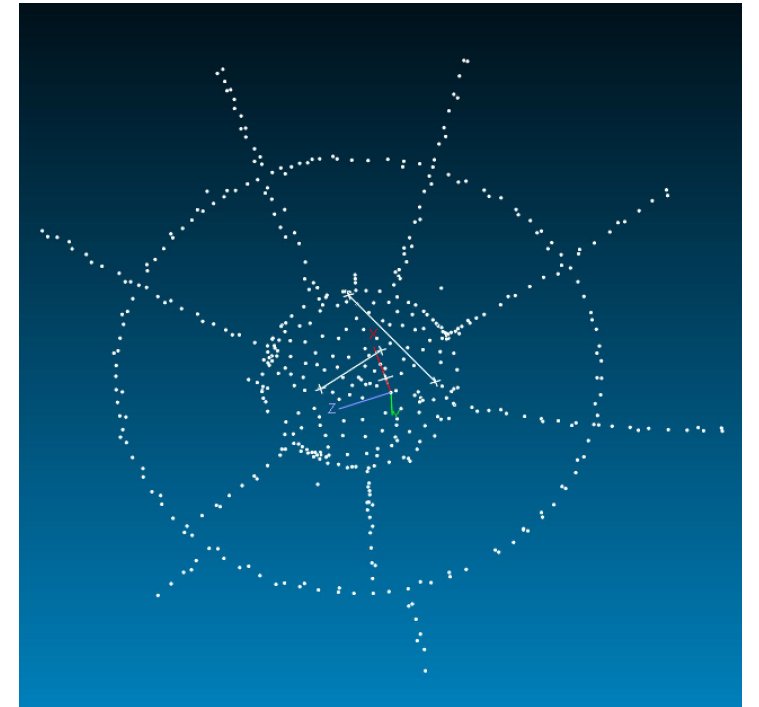
1. Sweep the reflector over the surface of the device under test while the laser tracker tracks it
2. Analyze the resulting 3D point cloud



# Photogrammetry System

Measurement procedure:

1. Glue coded and uncoded targets
2. Take pictures from different angles
3. Feed the pictures into the software to get the 3D point cloud

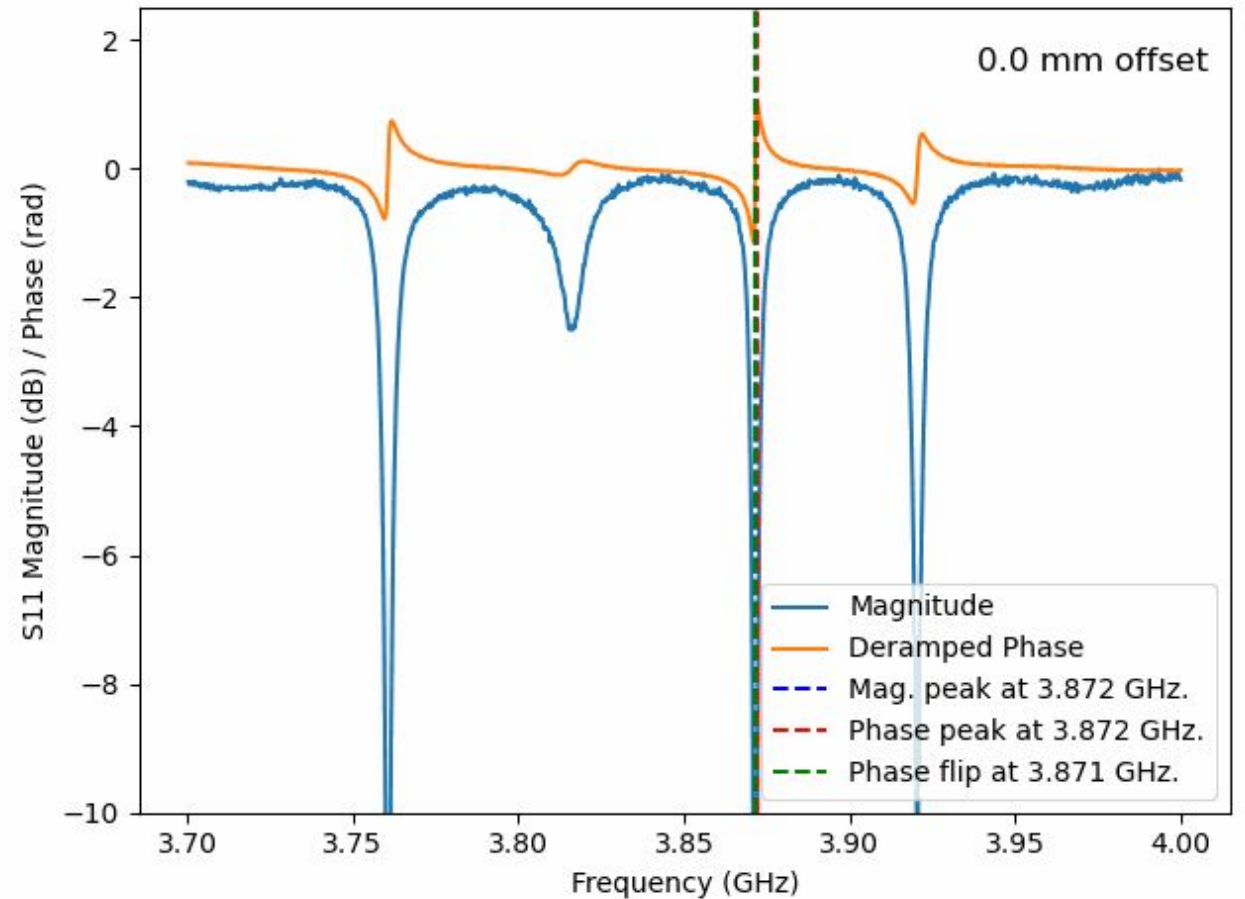
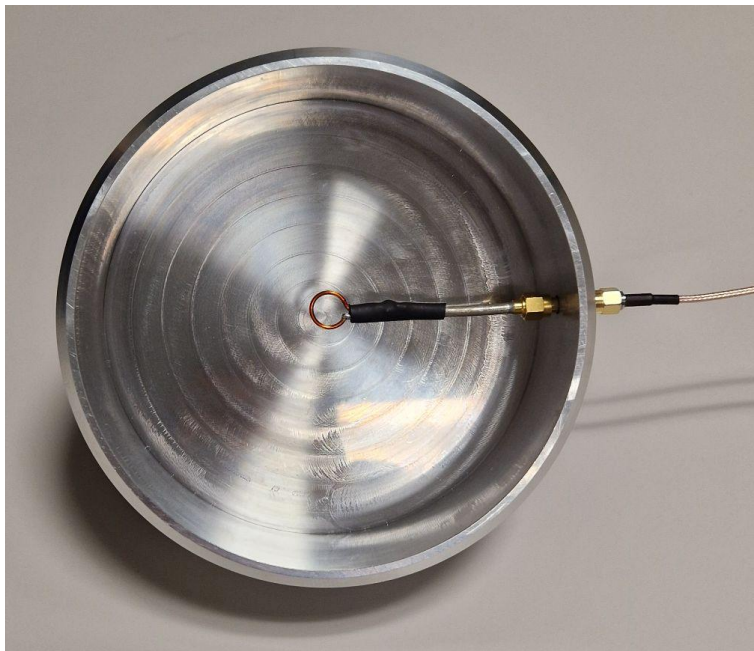


# Reflectometer

Tracks distance offset through shifting resonant frequency

Measurement procedure:

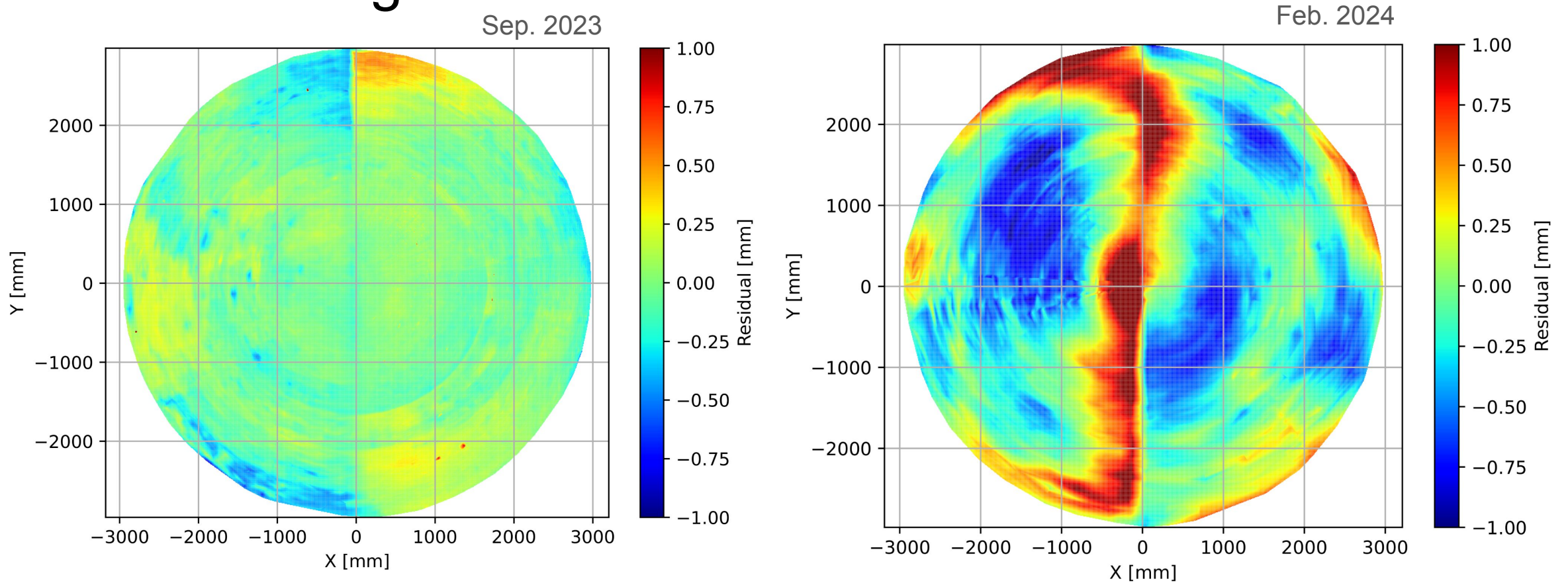
1. Calibrate with aluminium mesh
2. Take measurements at locations of interest
  - a. Measure the arc length/height
3. Analyze data



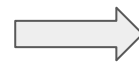


# Transport from OmegaVerse to Carnarvon

## Effect on Plug



The **RMS** value is **0.123 mm < 0.6 mm**  
The **focal length** is **1260.0 mm**

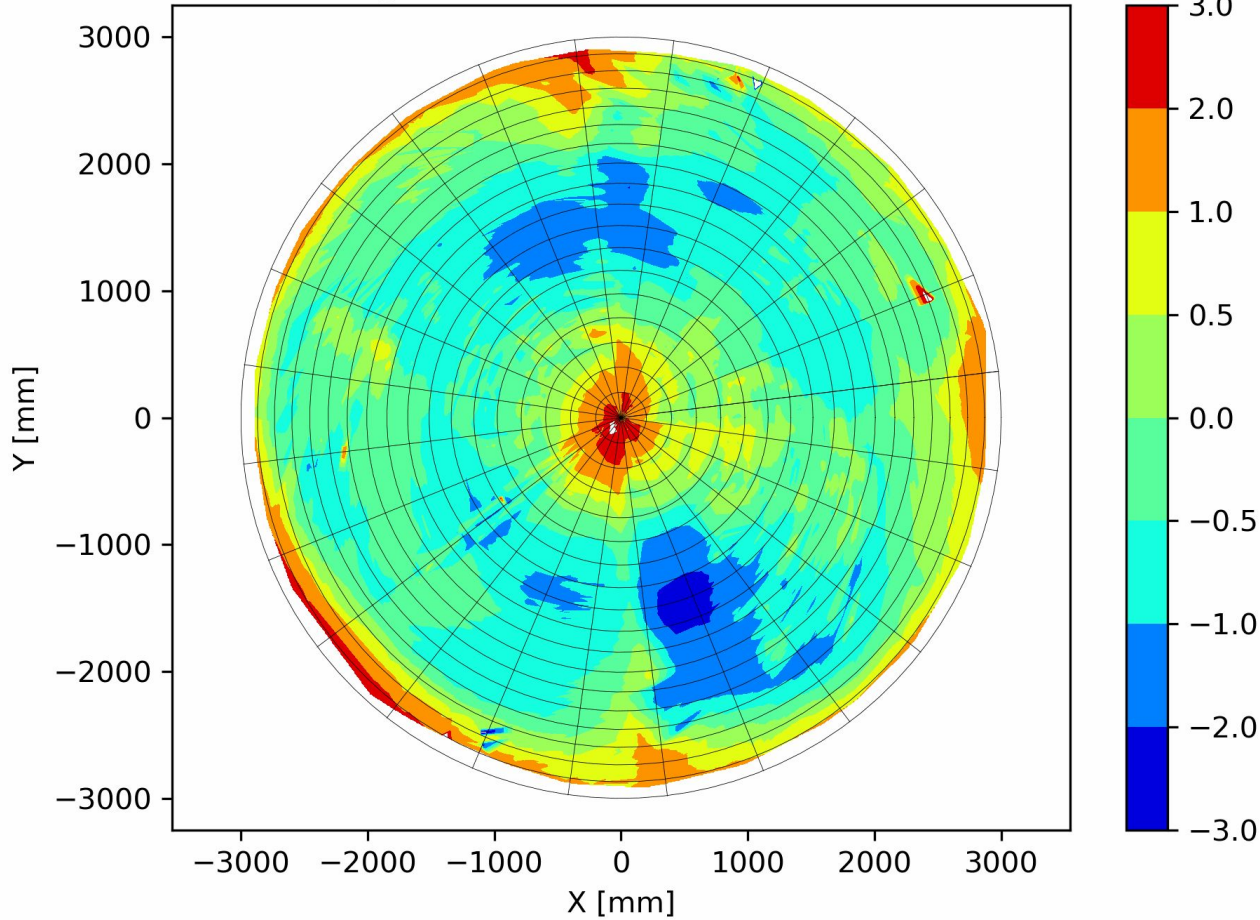


The **RMS** value is **0.510 mm < 0.6 mm**  
The **focal length** is **1260.7 mm**

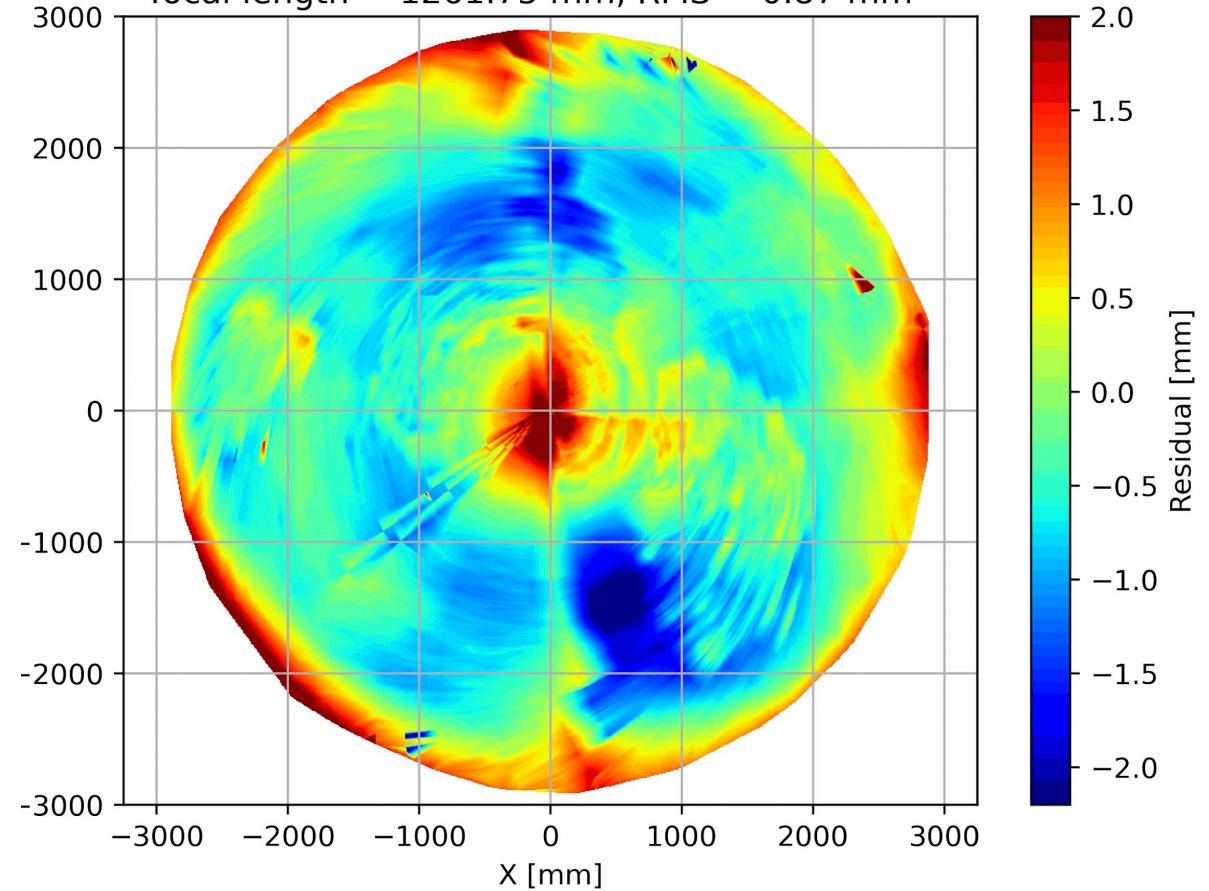
# Plug Improvement

June 2024

focal length = 1261.75 mm, RMS = 0.87 mm



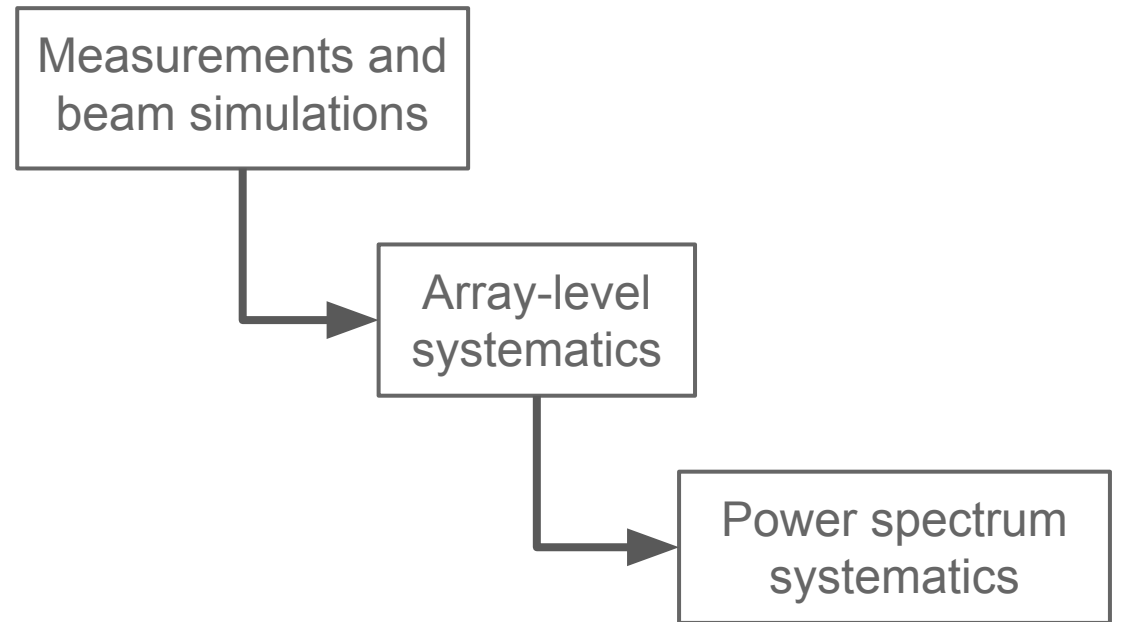
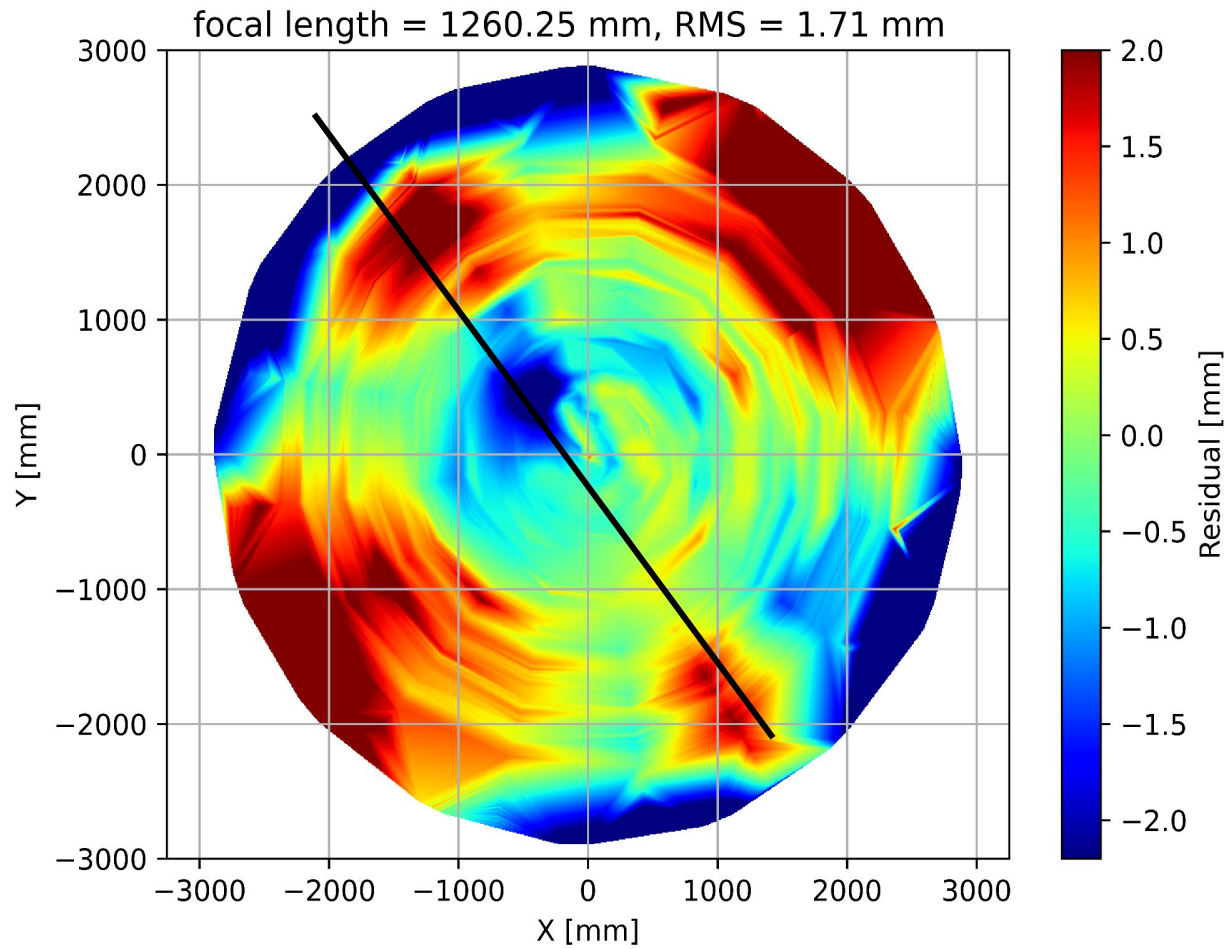
focal length = 1261.75 mm, RMS = 0.87 mm



with the support of Keshav Bechoo,  
Tasmiya Papiiah, Thierry Viant, and others

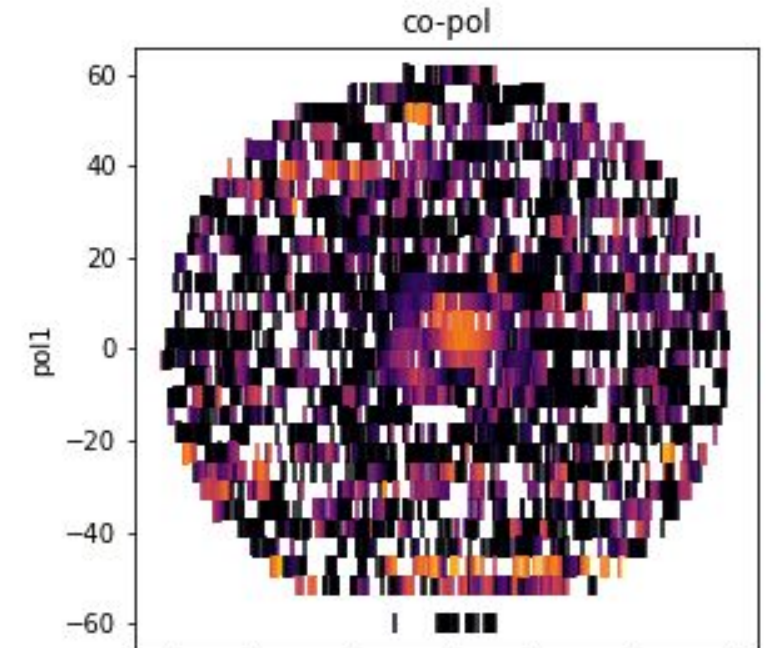
# Calibration

June 2024



# Drone Beam Mapping and Holography

- Drone mounted transmitter for direct beam mapping of
- Test flights at Bleien Observatory in Switzerland
  - Also with other groups at Green Bank CHIME outrigger and DRAO
- Multiple parallel efforts
- Exploring feasibility of flights at Karoo site
  - RFI characterisation and testing
- Comparing with holography and metrology based reconstruction



# Klerefontein test side ready in October!



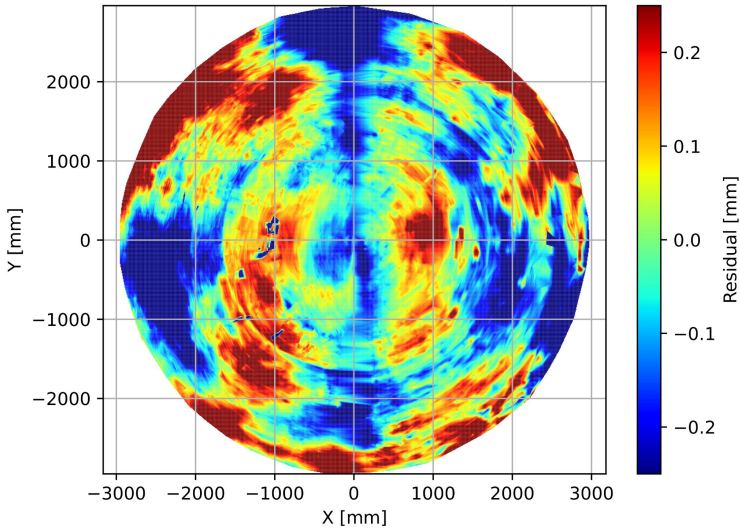
# Conclusions

- HI IM provides access to large cosmological volumes
- Systematics need to be controlled
  - Focus on the dish surface / primary beam
- Manufacturing process is in progress
  - Plug is ready for production
- Propagate dish surface deviations into beams
  - Understand the effects and combine with other methods
  - Enable the mitigation of systematics in the HI IM measurement to constrain cosmological parameters
- HIRAX will have first light in 2025



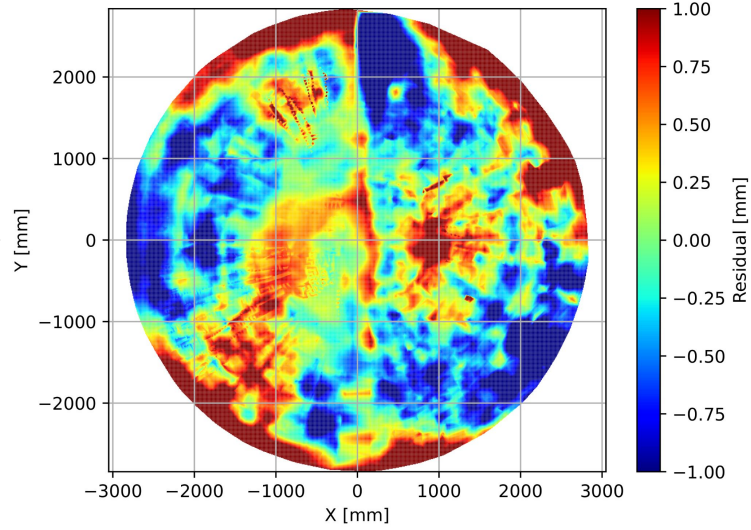
# RMS Error Propagation

## Plug (Split, Cape Town)



RMS = **0.205 mm**  
focal length = 1260.3 mm

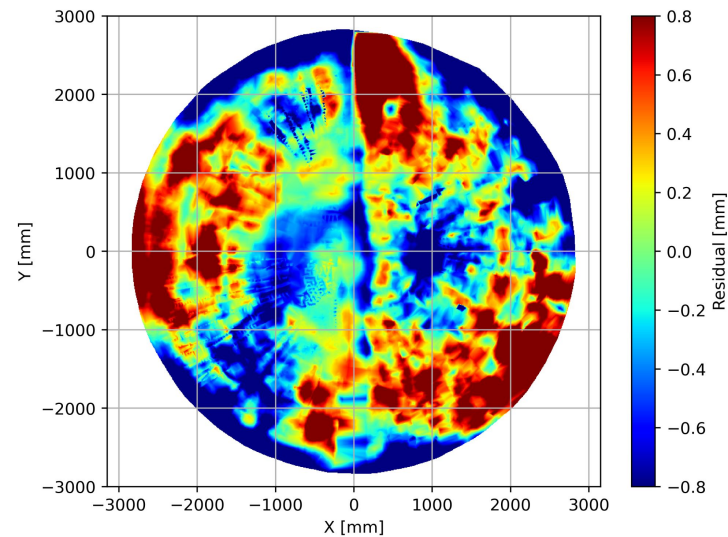
## Mold (Split)



RMS = **0.7 mm**  
focal length = 1258.6 mm

x4

## Split Dish 1



RMS = **1.7 mm**  
focal length = 1260.25 mm

x2.4

