

A large radio telescope dish is positioned on a hill in the foreground. In the background, there are snow-capped mountains. The sky is dark with numerous white star trails, indicating a long-exposure photograph. A bright red light is visible at the top of the telescope dish. The overall scene is a mix of natural and technological elements.

# Protecting the Quiet Sky

*Why Satellites Matter for Radio Astronomy*

Gregory Hellbourg  
California Institute of Technology

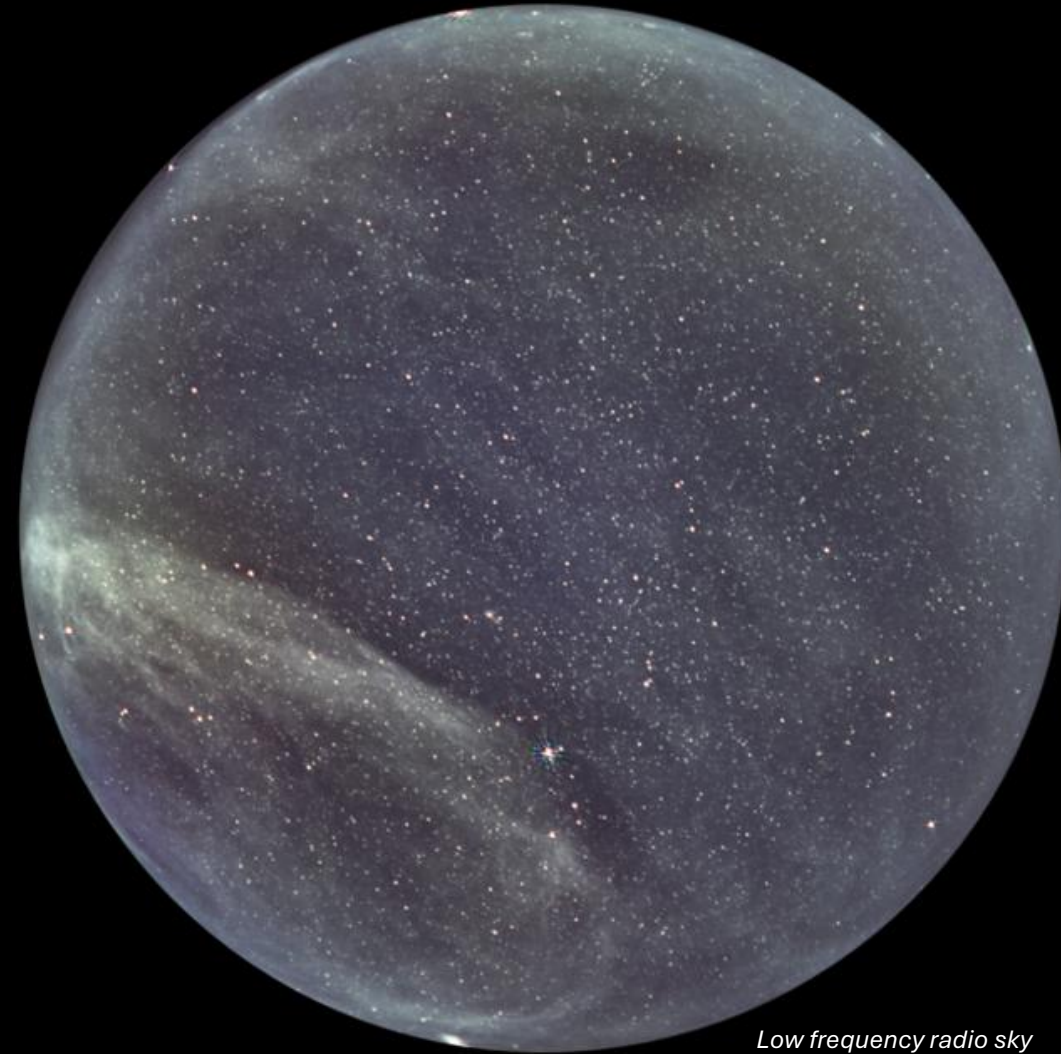
# What is Radio Astronomy?

Radio astronomy detects natural radio waves from the Universe

These waves are astonishingly faint, often **10 trillion times weaker** than a phone signal received on Earth

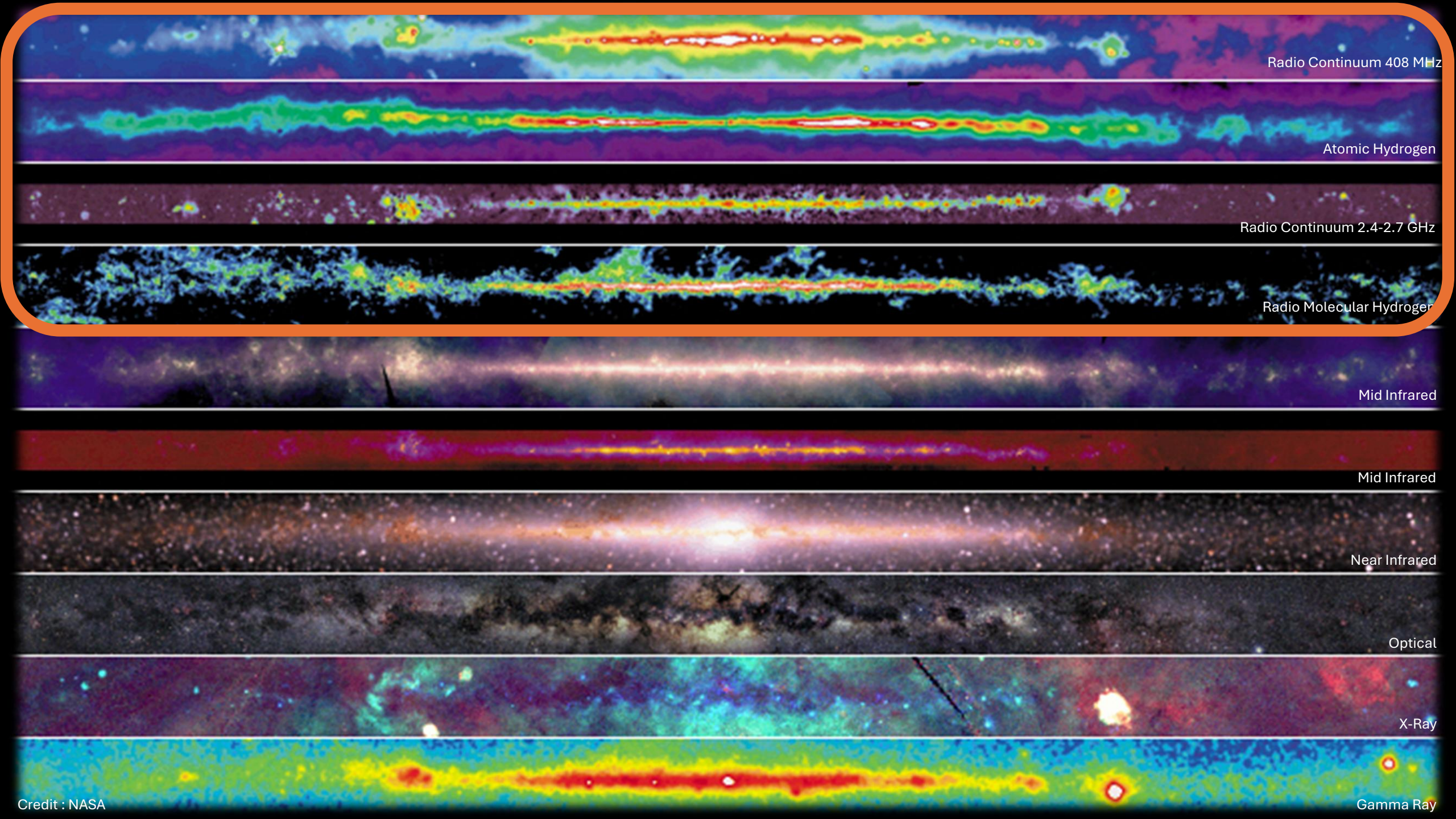
We use them to study:

- hydrogen in galaxies
- black holes
- planetary atmospheres
- cosmic magnetism
- the early Universe



*Low frequency radio sky*  
Credit : G.Hallinan  
Caltech/OVRO-LWA





Radio Continuum 408 MHz

Atomic Hydrogen

Radio Continuum 2.4-2.7 GHz

Radio Molecular Hydrogen

Mid Infrared

Mid Infrared

Near Infrared

Optical

X-Ray

Gamma Ray



# Why Is Radio Astronomy Valuable to Society?

## *Navigation and GPS stability*

- VLBI networks anchor the reference frame used by GNSS.
- Without radio astronomy, global navigation would slowly drift.

## *Space weather / solar bursts monitoring*

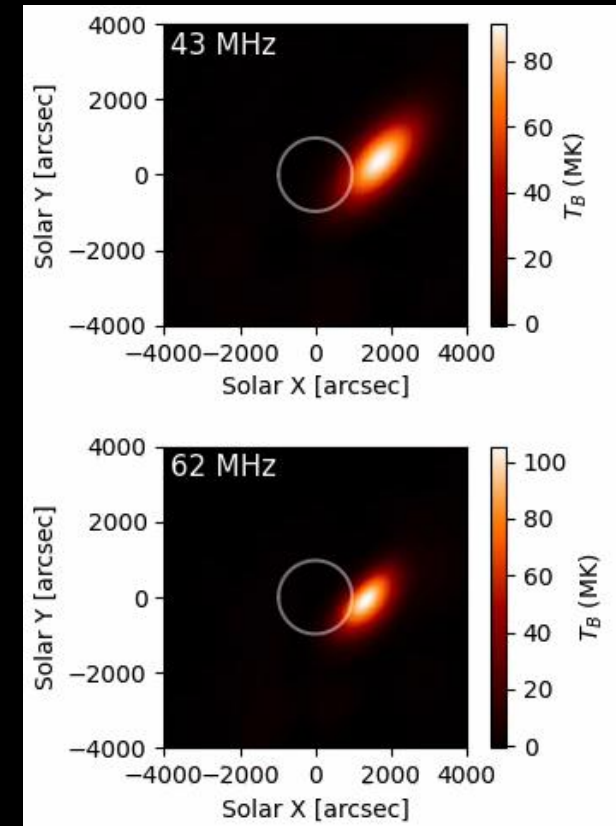
- satellites
- power grids
- aviation
- GNSS

## *Technology spin-offs*

- low-noise amplifiers → used in telecom
- fast signal processing → Wi-Fi
- imaging algorithms → MRI
- timing → finance, navigation

## *Global scientific prestige*

Radio astronomy is part of the world's scientific identity.



X2.8 solar flare  
2023-12-14  
Caltech/OVRO-LWA/NJIT

# Why Are Radio Telescopes So Vulnerable?

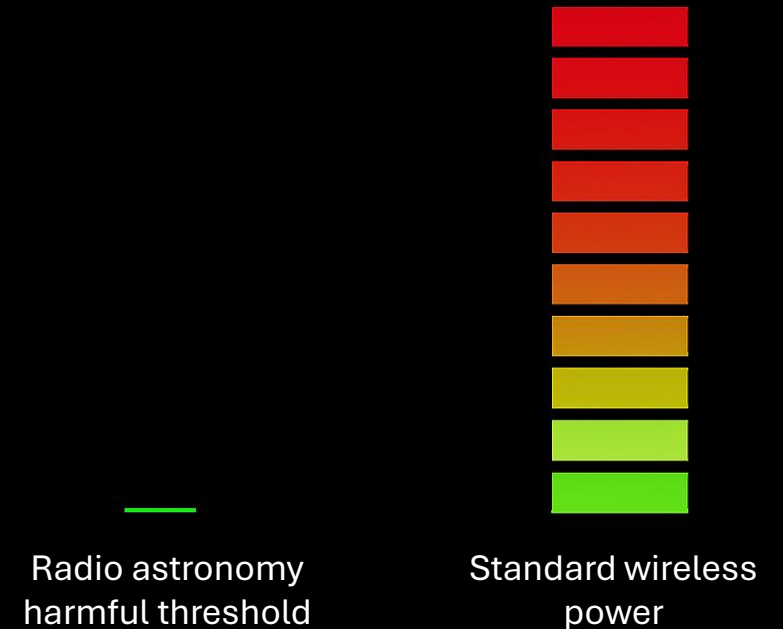
Radio telescopes are designed to hear extraordinary silence.

A harmful interference threshold may be as low as  $-250$  dBW, essentially zero power.

These thresholds are **1 billion times lower** than levels relevant to telecom systems.

Why?

- Large collecting area
- Large frequency bandwidths
- Low noise receivers
- Coherence between antennas kilometers apart



# What Has Changed: The Satellite Boom

Historically:

- < 1,000 active satellites

- Rarely aligned with telescope beams

Now:

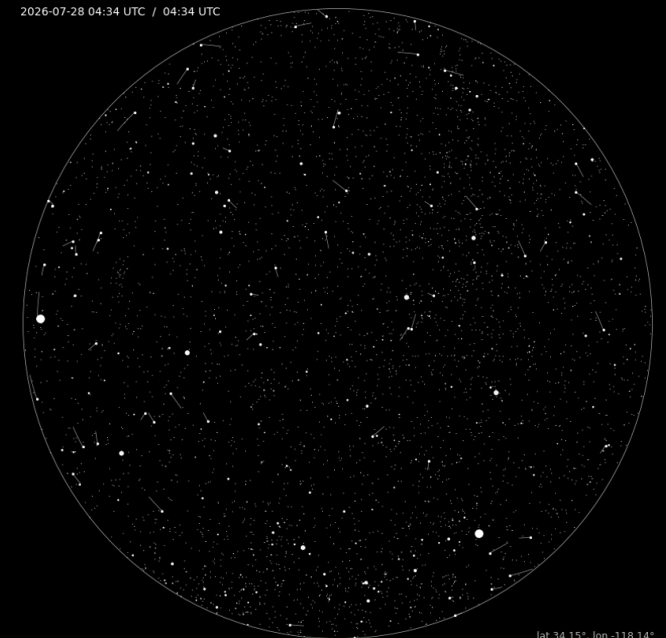
- 12,000 satellites already

- Constellations may reach 30,000–100,000

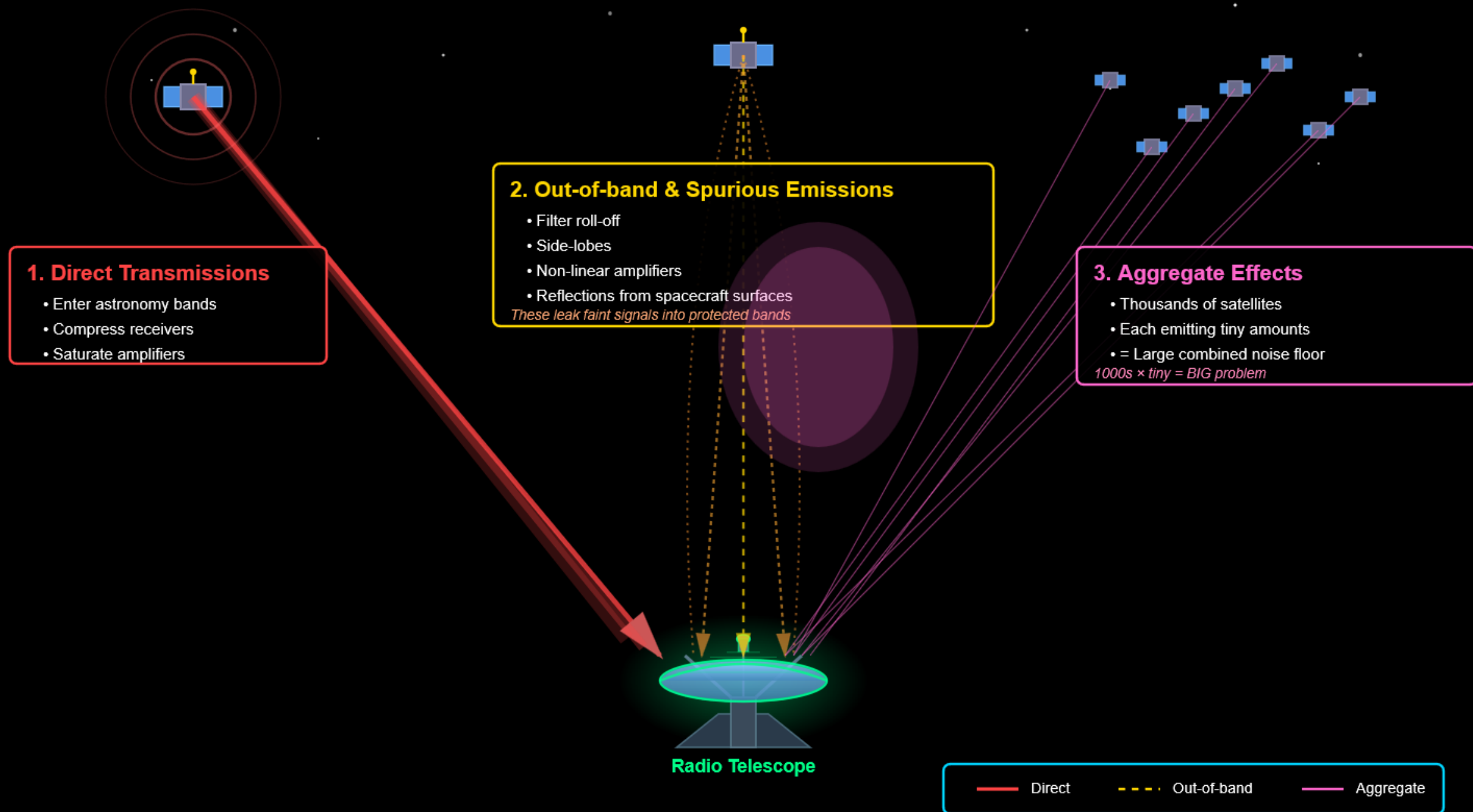
- GNSS, broadband, Earth observation, IoT

Result:

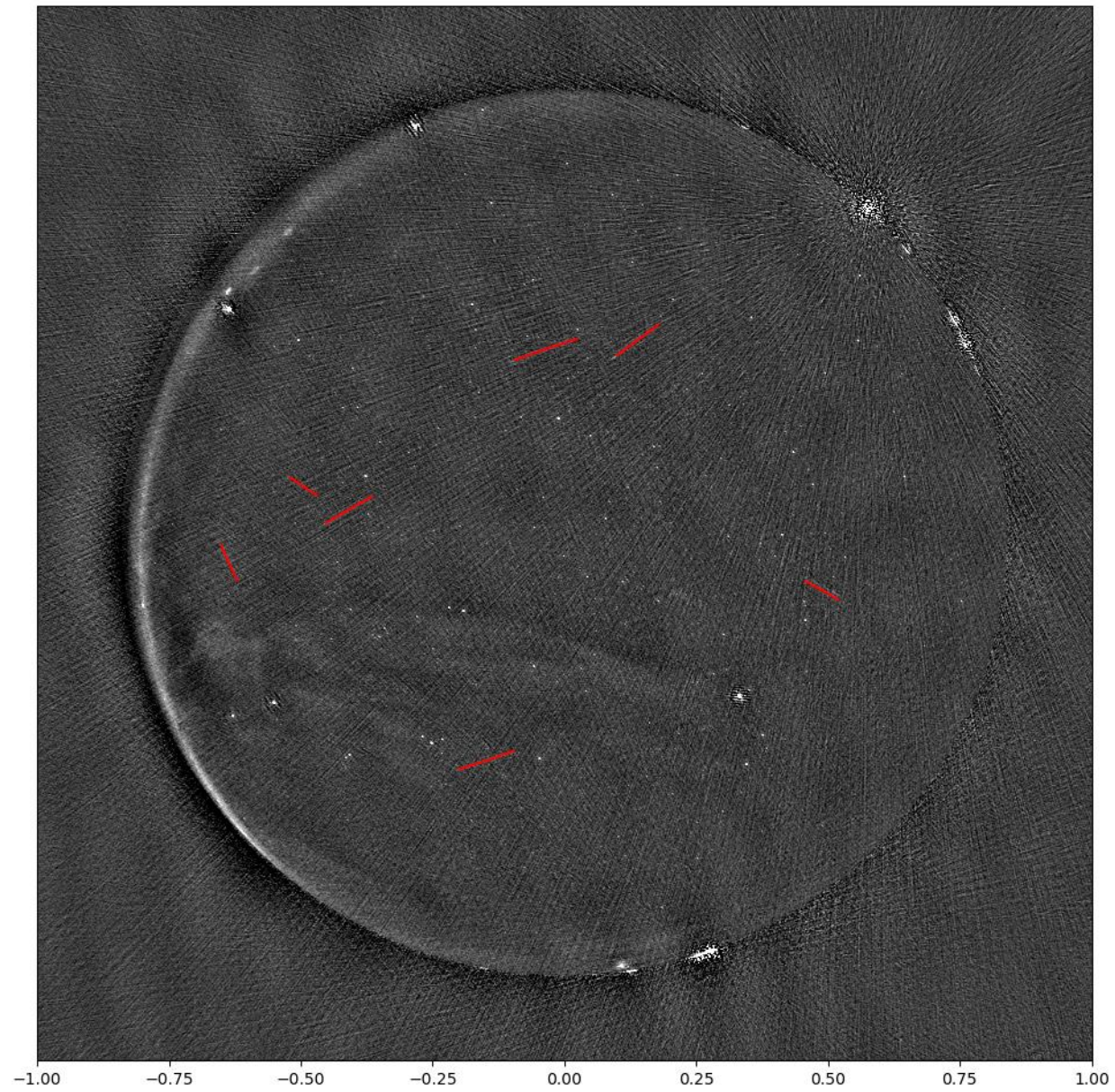
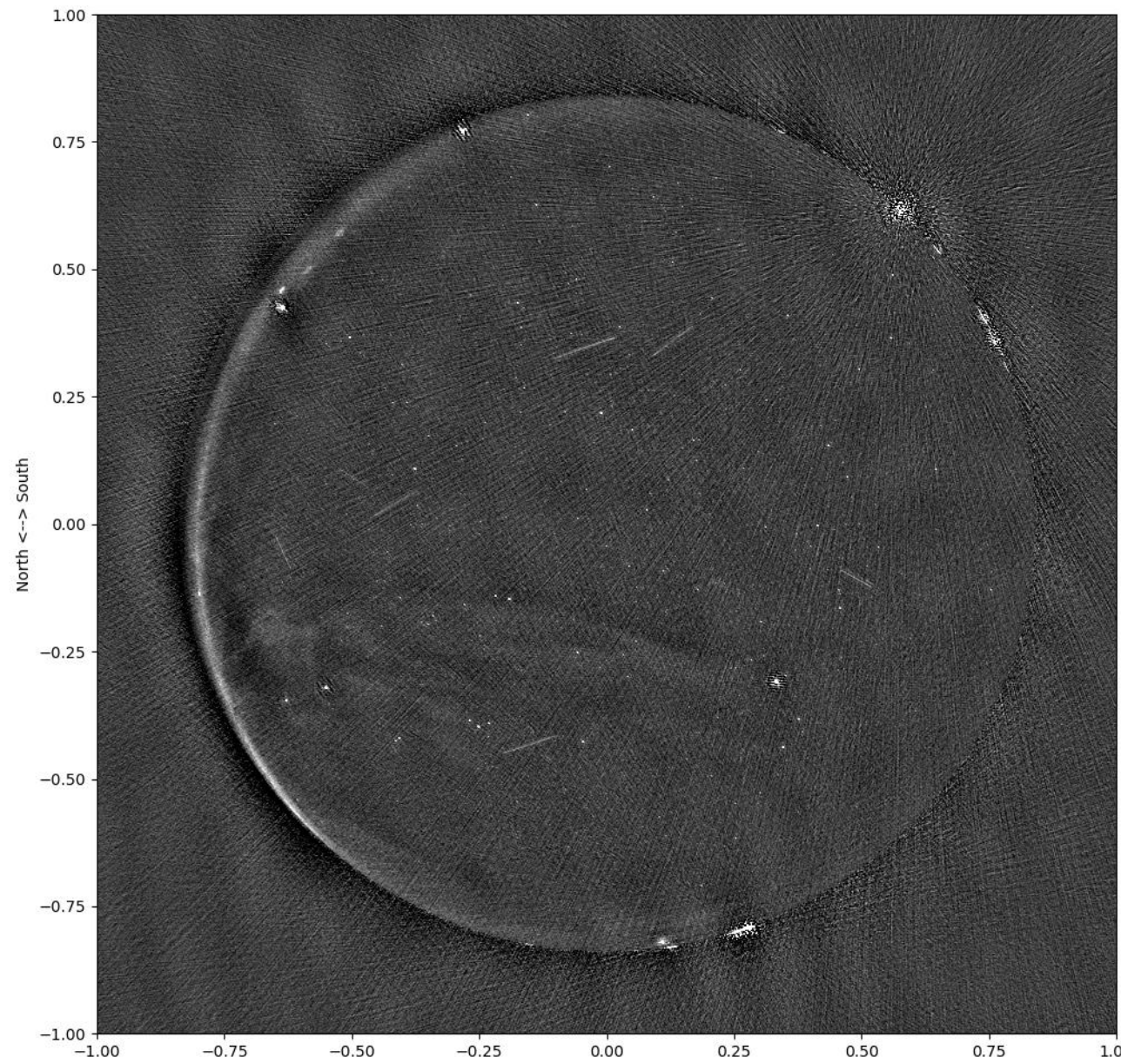
- The probability a satellite crosses a telescope beam becomes **near-continuous**.



# How Satellites Interfere with Radio Astronomy









# The Economic Cost

## Cost per hour of observing time

- Green Bank Telescope: \$1,400–\$2,000/hr
- VLA: \$1,700/hr
- ALMA: \$12,000/hr
- SKA (projected): €10,000+/hr

## If interference raises system temperature 2–3×

- observing time increases 4–9×
- direct economic losses = hundreds of thousands to millions per project

## Other costs:

- extra staff
- more computing
- longer calibration
- regulatory engagement
- lost scientific opportunities

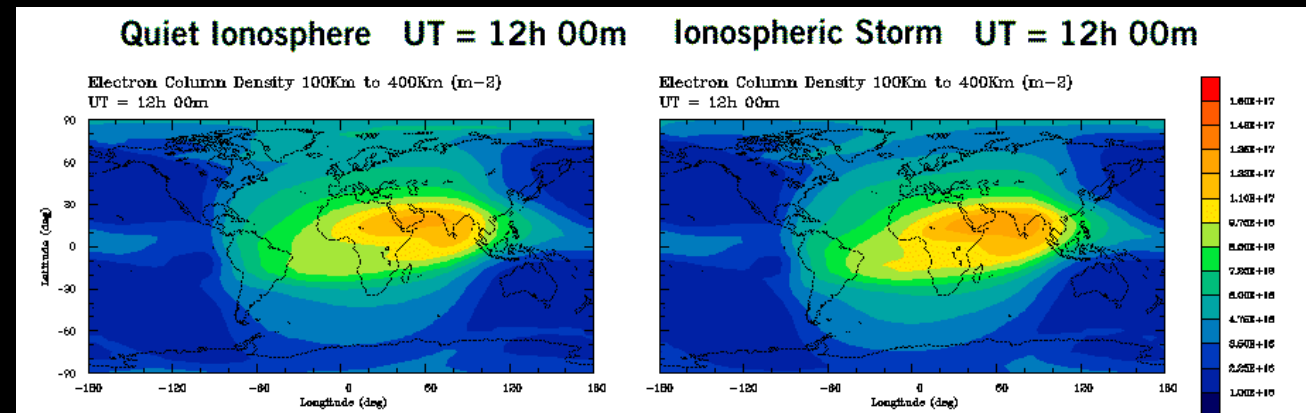
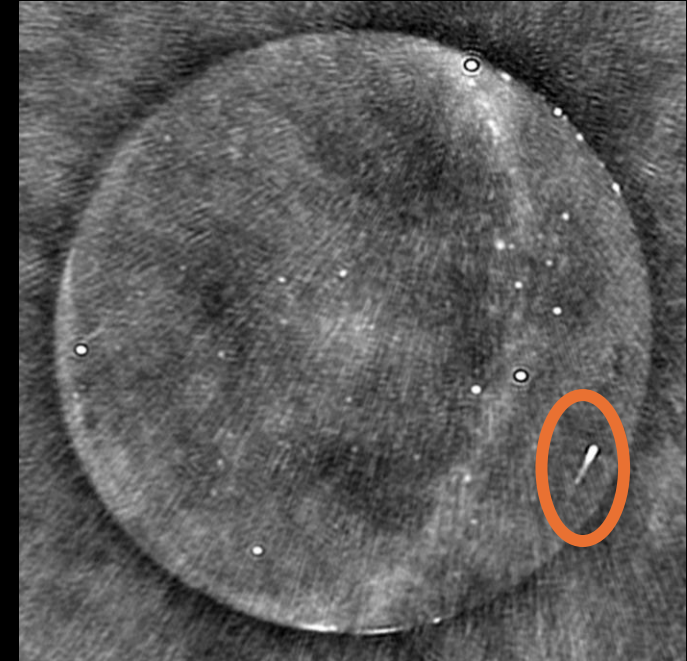
# Atmospheric & Launch Effects

Rocket launches and satellite re-entries create:

- plasma trails
- shock-induced electron variations
- ionospheric disturbances

These disrupt:

- radio cosmology
- Earth rotation measurements
- 21-cm experiments





# Conclusion : The radio sky is no longer quiet

## Astronomical signals remain extraordinarily faint

- often far below thermal noise
- requiring extreme sensitivity and stability

## Satellite presence is now continuous

- thousands of spacecraft across all orbits
- beams, reflections, and spurious emissions are always present

## Interference accumulates

- even tiny out-of-band levels become significant when thousands of satellites contribute
- system temperatures rise, sensitivity decreases

## Scientific impact is already measurable

- longer observing times
- reduced data quality
- increased operational load
- some experiments (e.g. 21-cm cosmology) becoming more difficult

The Quiet Sky Report:

