

# Rubin Observatory and Satellite Constellations

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*with contributions from many wonderful colleagues!*

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**#DarkAndQuietSkies**  
for Science and Society



U.S. National  
Science Foundation



U.S. DEPARTMENT  
of ENERGY

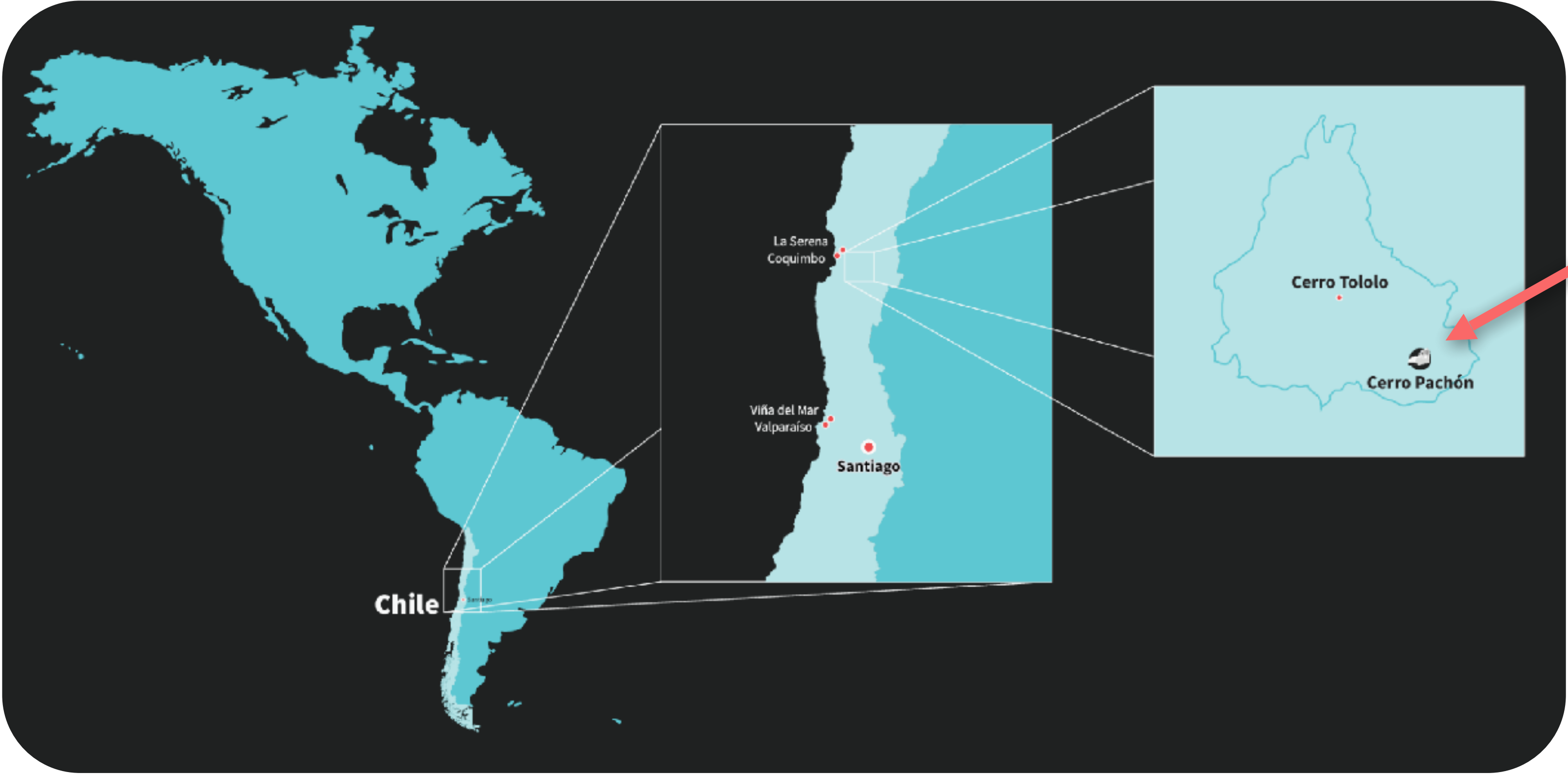
Office of  
Science



**SKAO**



# Meet Rubin Observatory



Located on **Cerro Pachón**, in the Coquimbo region of Chile

Jointly funded by the **U.S. National Science Foundation** and the **U.S. Department of Energy, Office of Science**



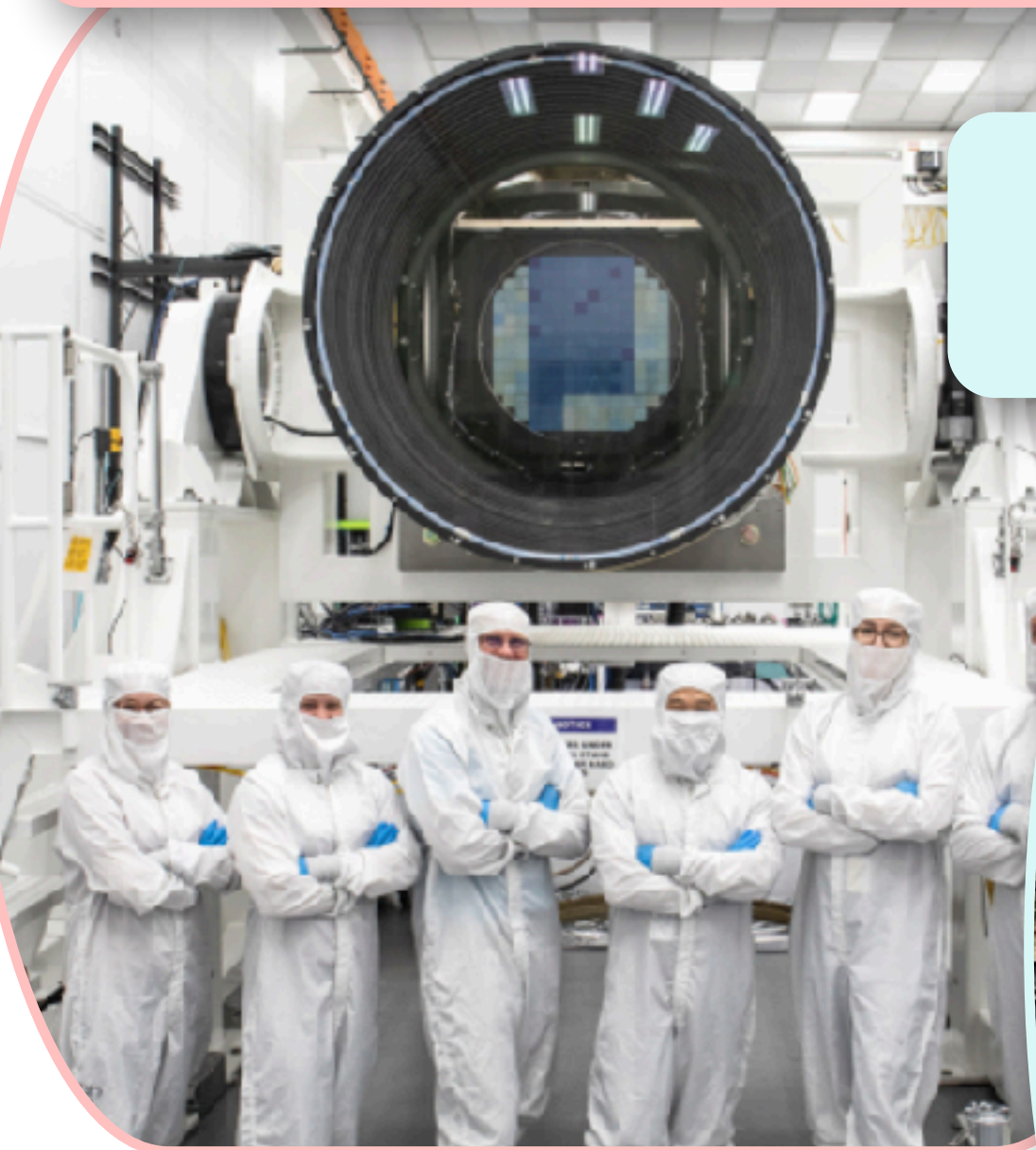


# Mission: Capture the Cosmos

The greatest astronomical movie of all time

## Wide Field of View

Largest digital camera ever built



J. Orrell/SLAC National Accelerator Lab



## Speed

Novel three-mirror design



RubinObs/NOIRLab/SLAC/NSF/DOE/AURA

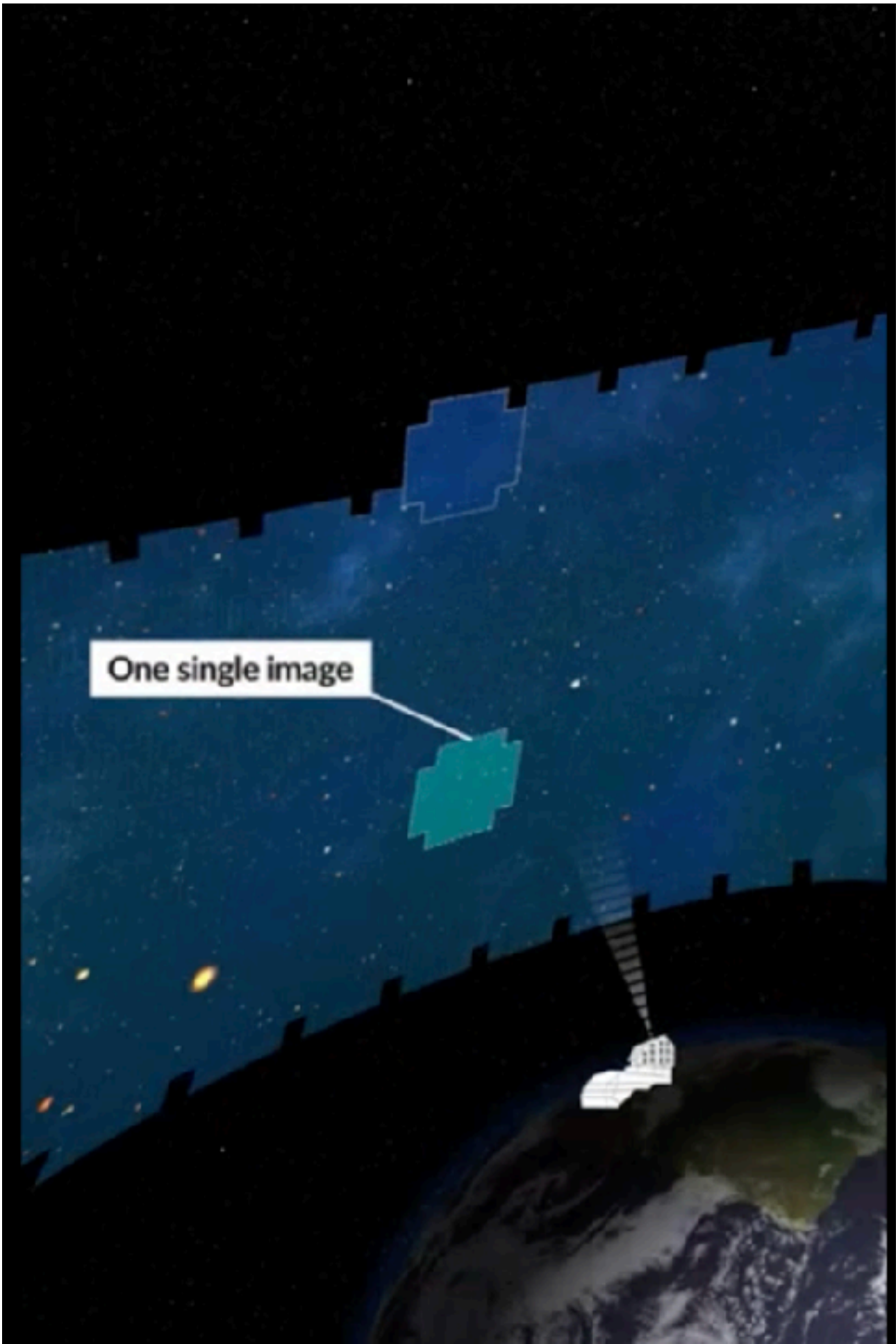


## Ability to see faint objects

Across the entire Southern sky



RubinObs/NOIRLab/SLAC/NSF/DOE/AURA





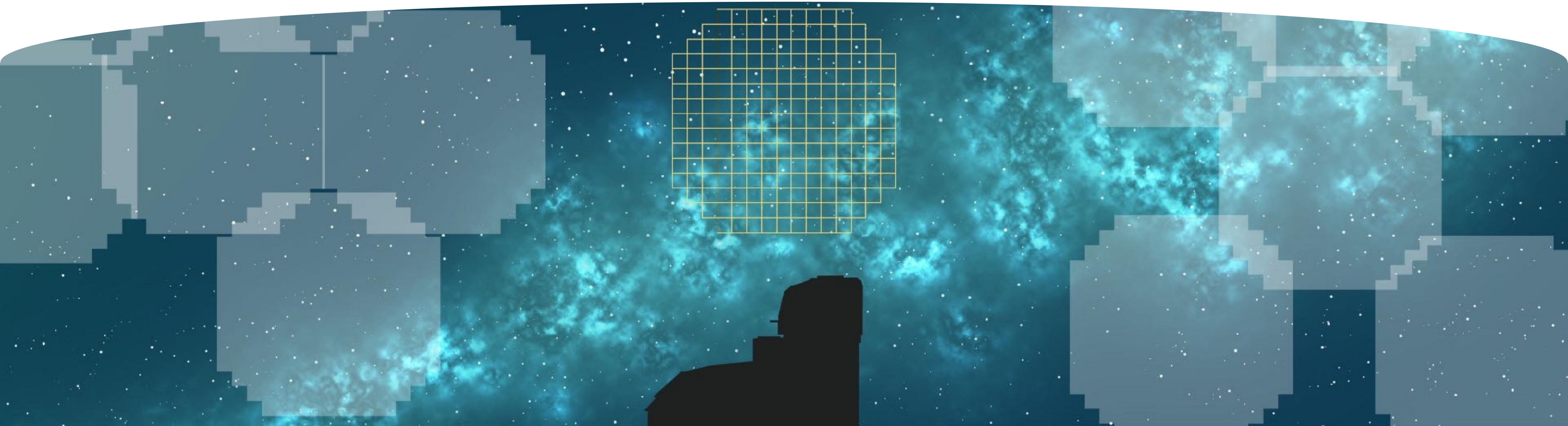
# Rubin's Legacy Survey of Space and Time (LSST)

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Repeatedly scan the southern sky every ~3 nights for 10 years

Wide-field survey with a fast cadence in 6 color filters

LSST will build the most detailed time-lapse view of the cosmos ever generated





# Four Science Areas

## Milky Way Structure & Formation

Rubin will help us make the best map of our home galaxy yet.



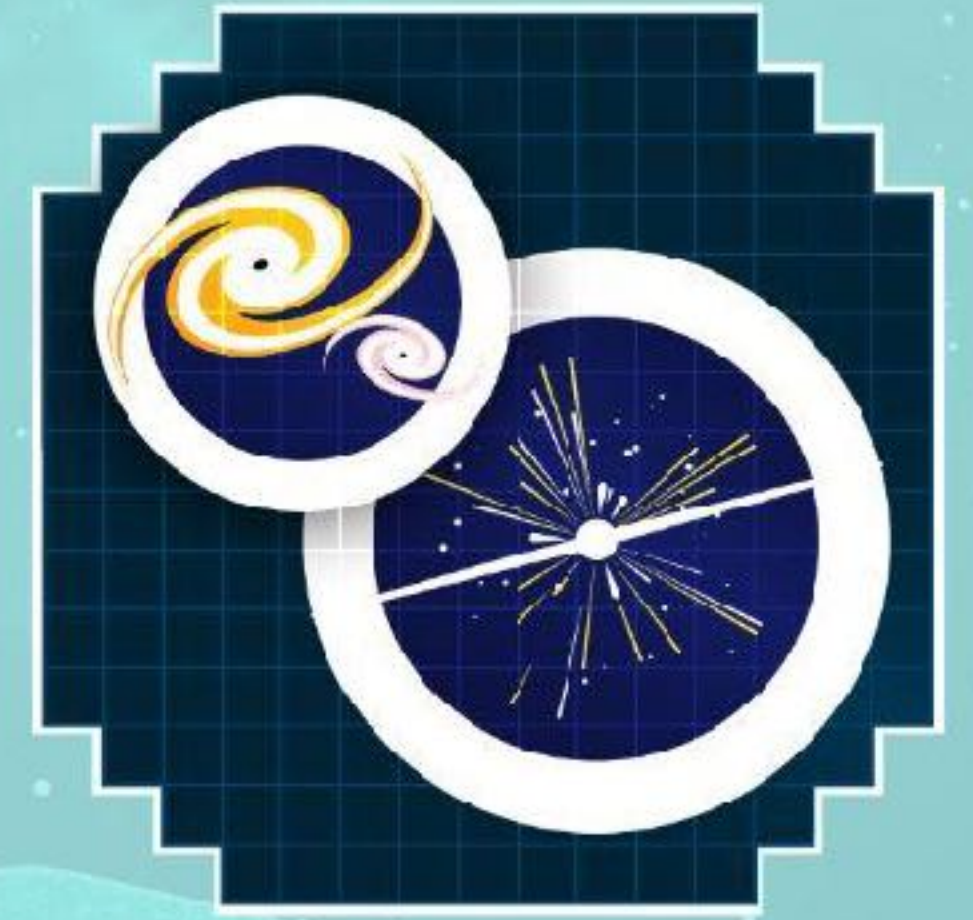
## Dark Matter & Dark Energy

Rubin is a brand new tool to help us learn more about their nature & behavior.



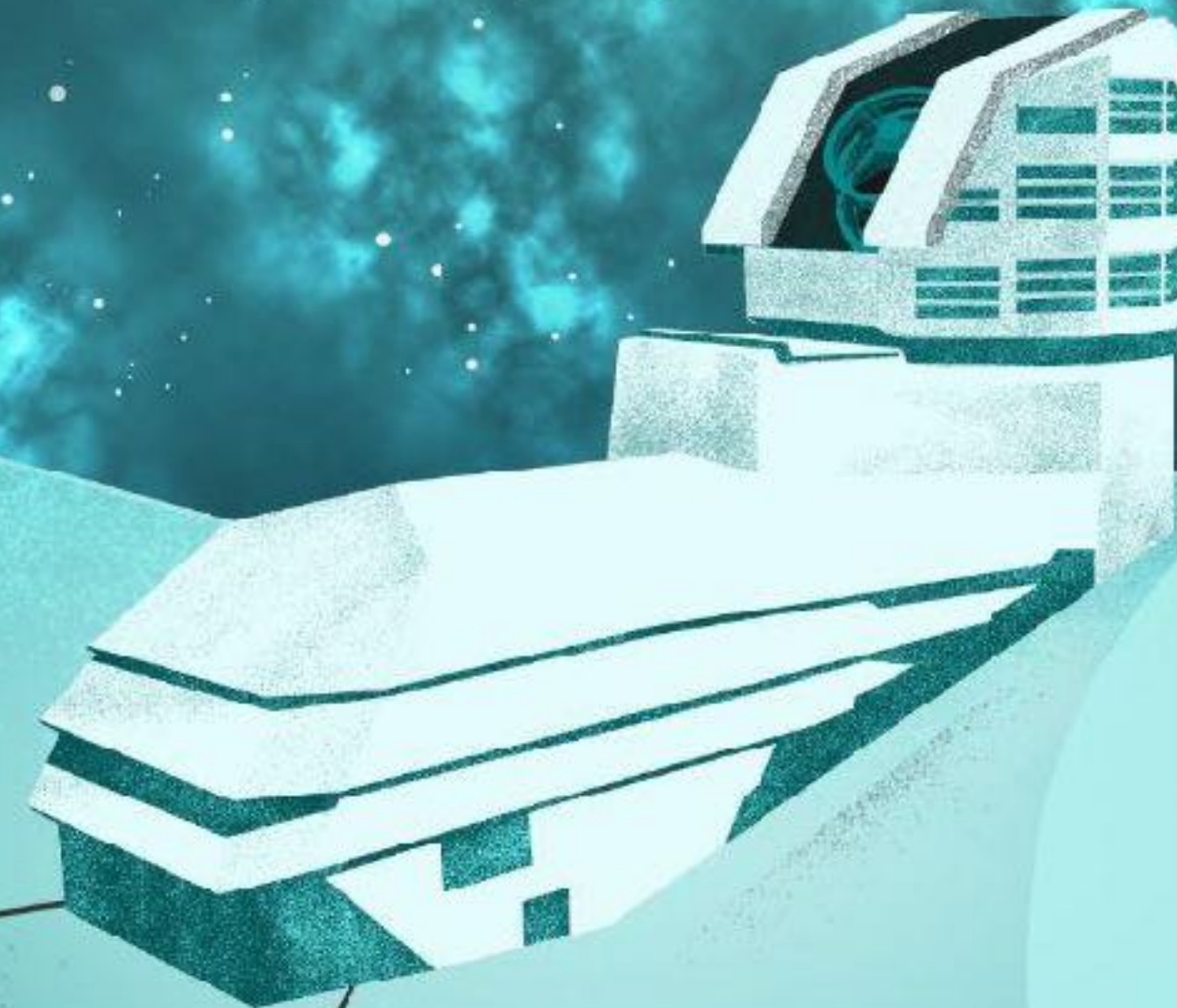
## Solar System Census

Rubin will show us millions of new asteroids and comets, and so much more.

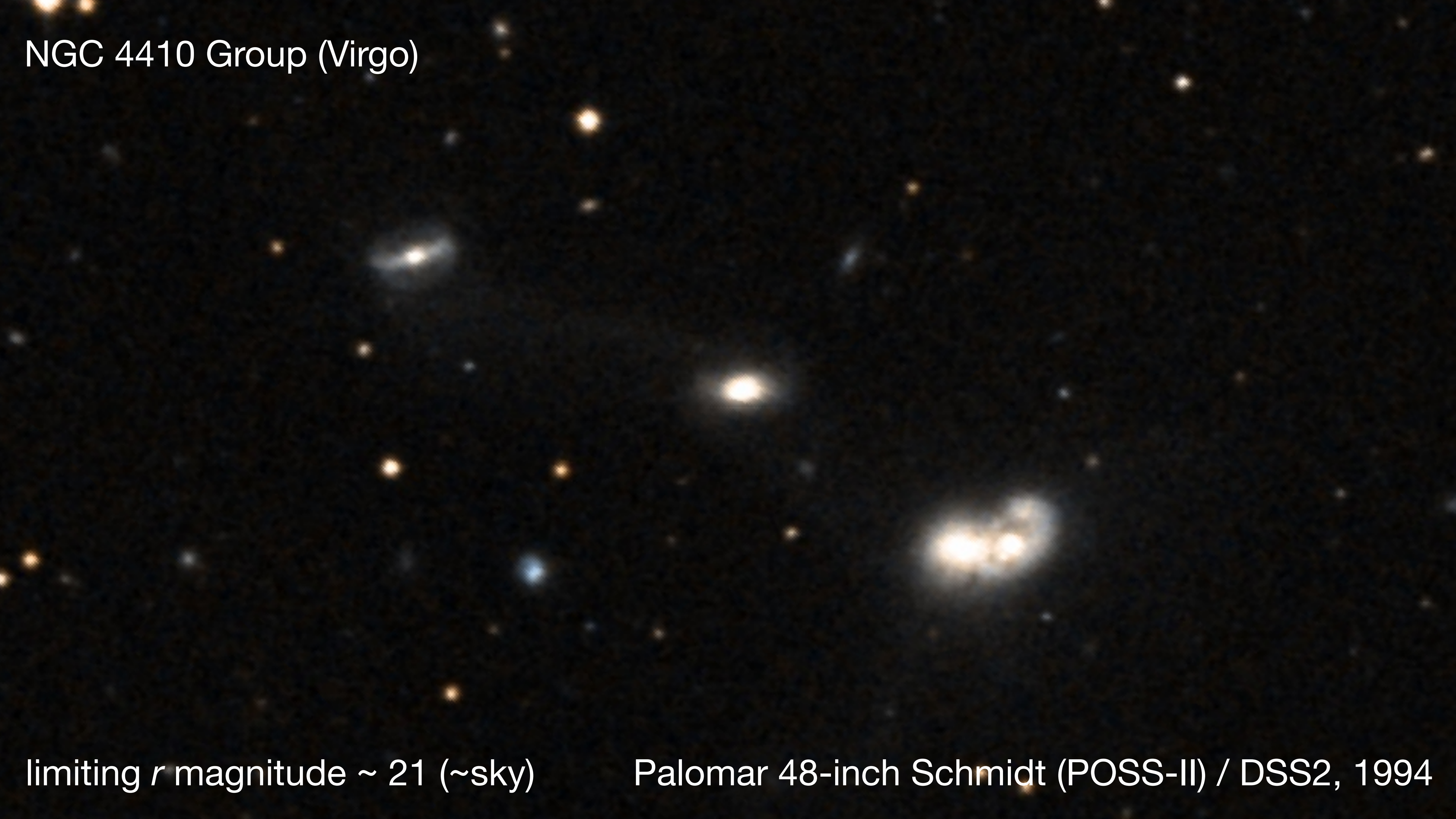


## The Changing Sky

Rubin will bring the night sky to life, yielding a treasure trove of discoveries.







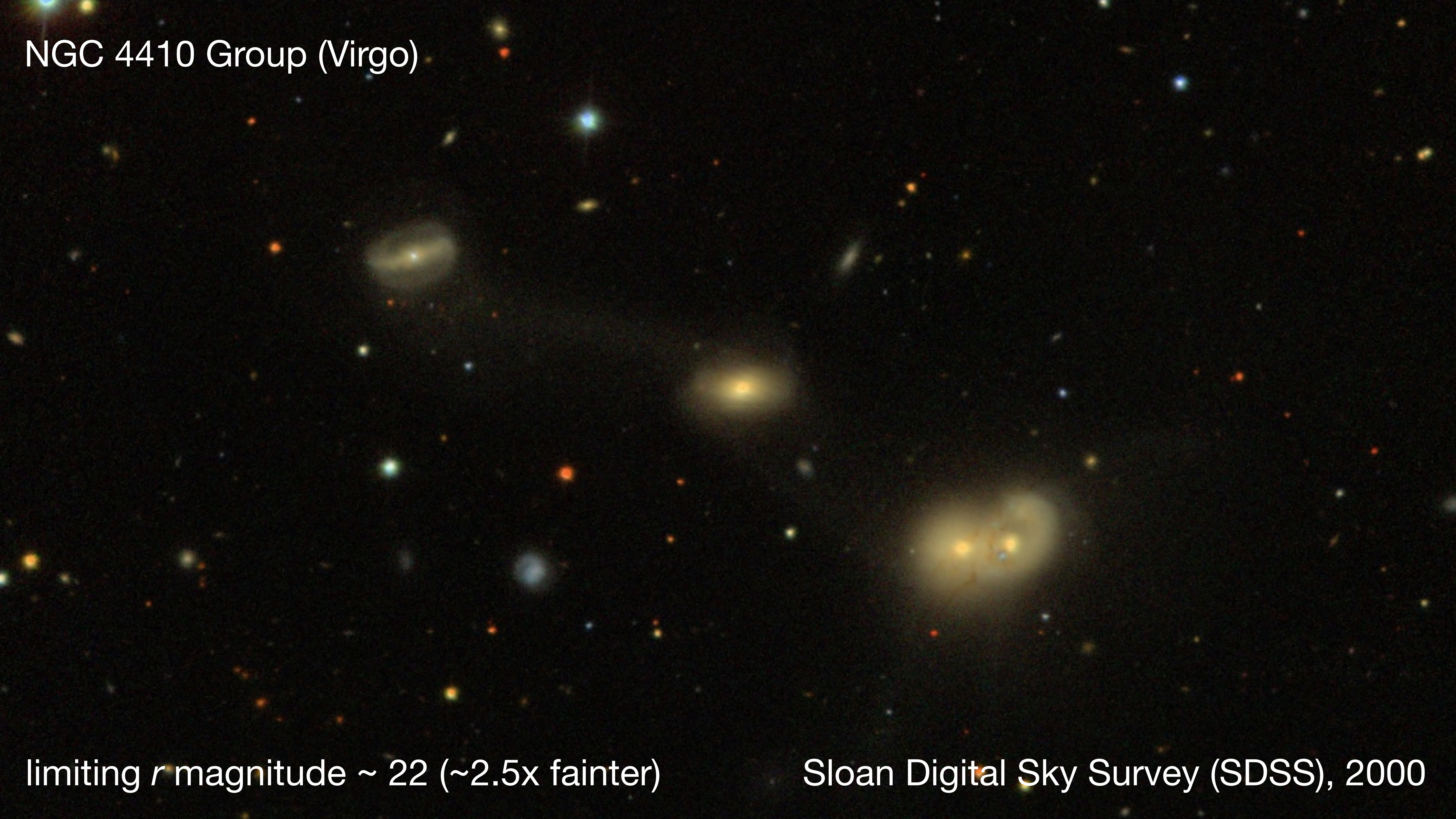
NGC 4410 Group (Virgo)

limiting *r* magnitude ~ 21 (~sky)

Palomar 48-inch Schmidt (POSS-II) / DSS2, 1994



# NGC 4410 Group (Virgo)

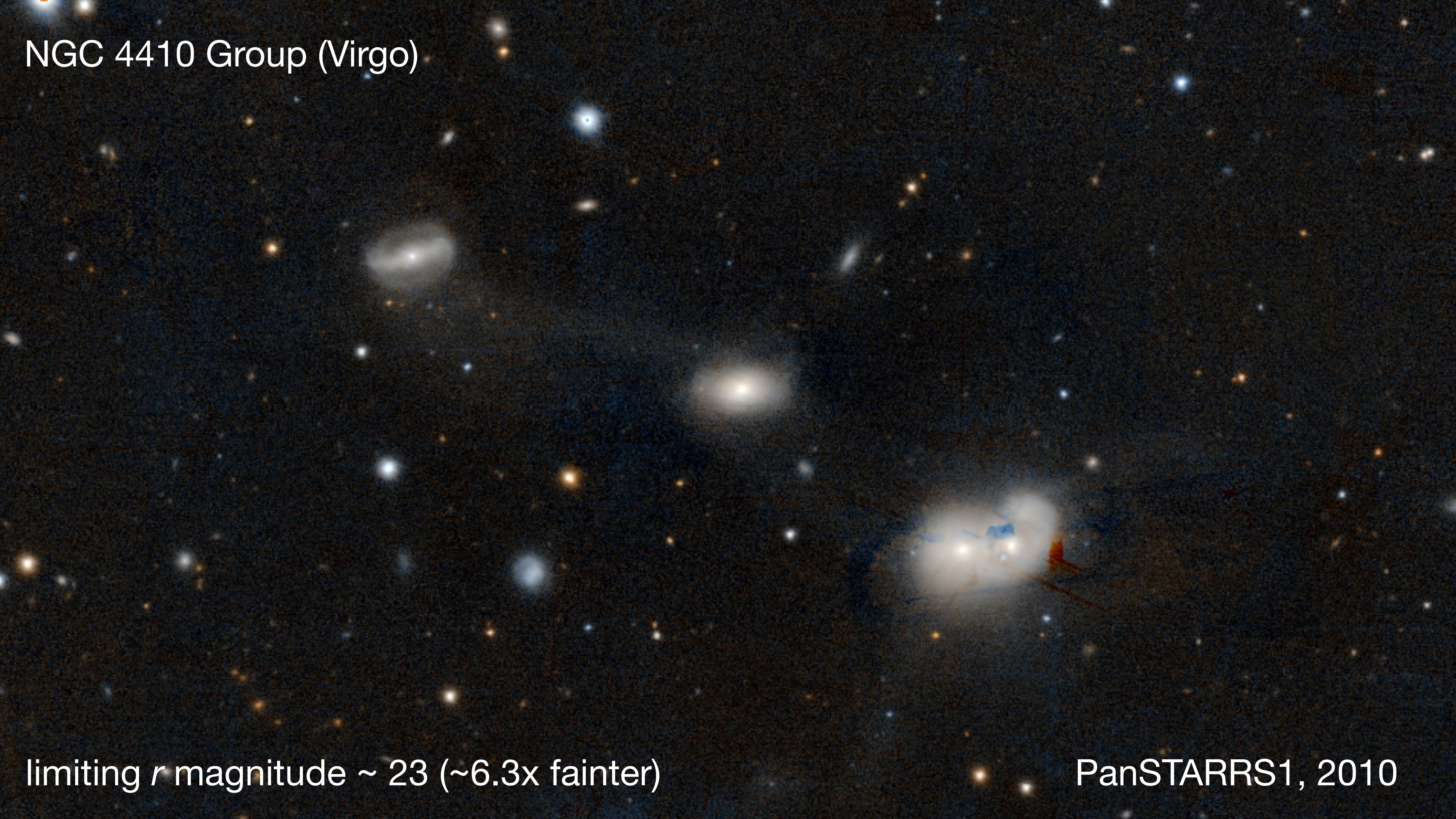


limiting  $r$  magnitude  $\sim 22$  ( $\sim 2.5\times$  fainter)

Sloan Digital Sky Survey (SDSS), 2000



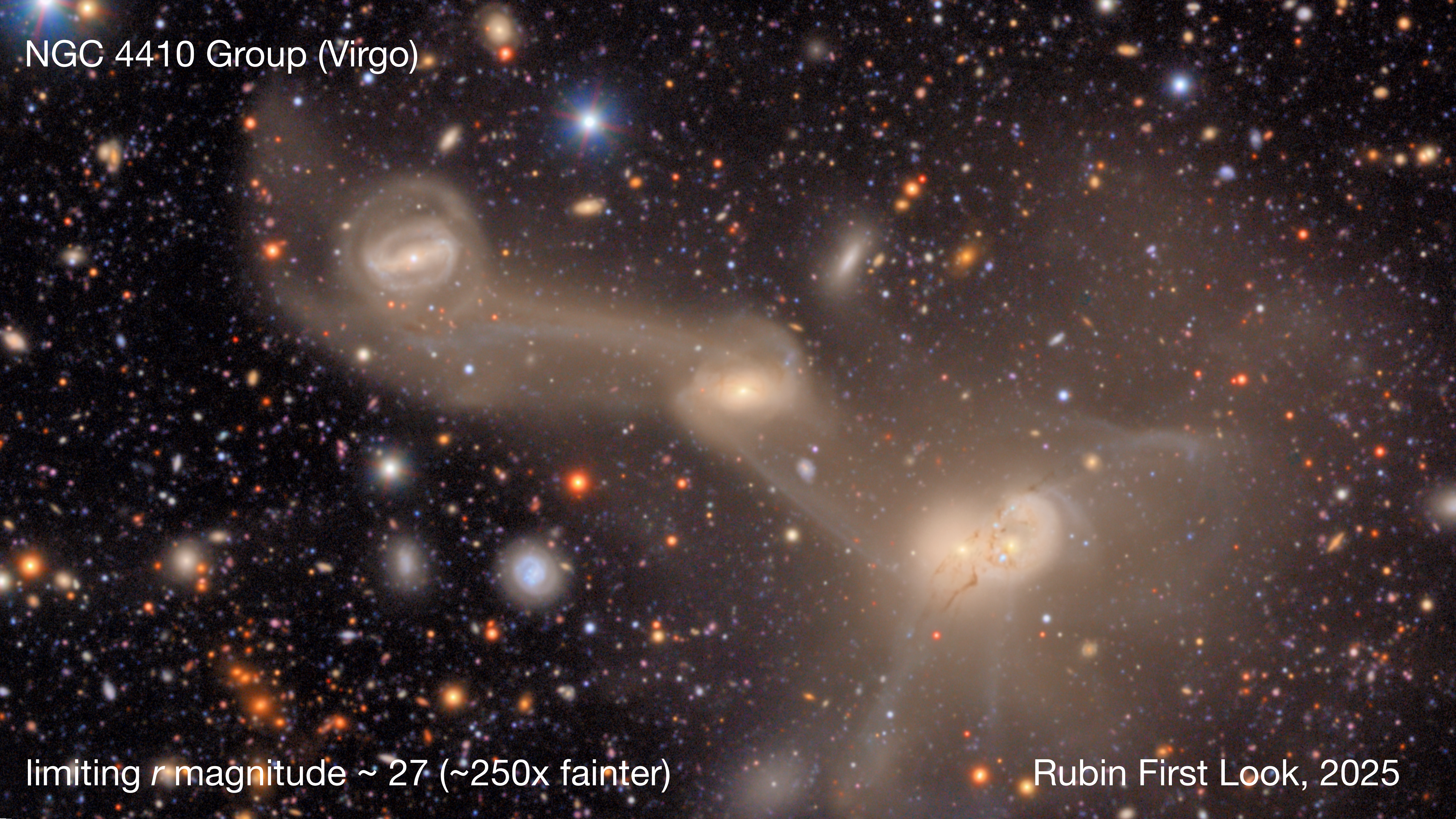
# NGC 4410 Group (Virgo)



limiting  $r$  magnitude  $\sim 23$  ( $\sim 6.3\times$  fainter)

PanSTARRS1, 2010





NGC 4410 Group (Virgo)

limiting  $r$  magnitude  $\sim 27$  ( $\sim 250\times$  fainter)

Rubin First Look, 2025



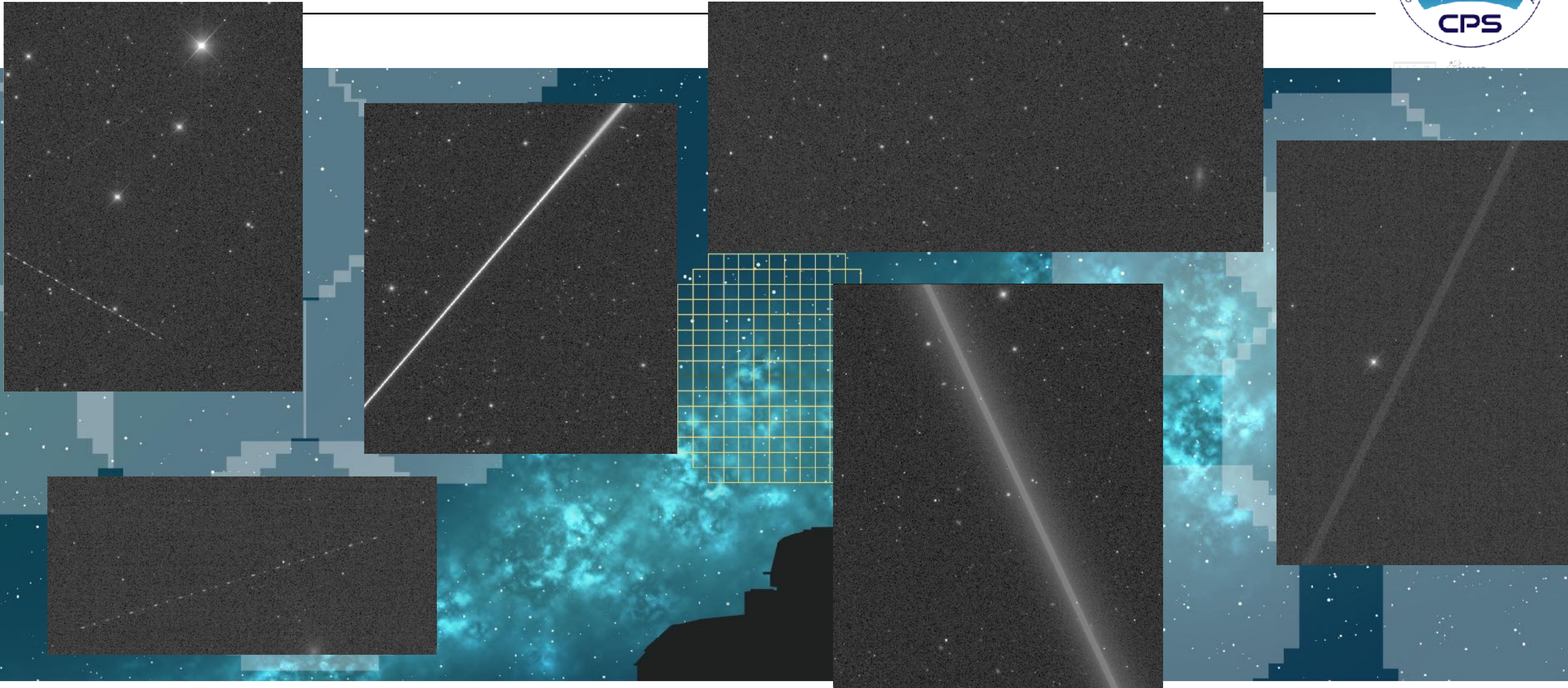


## The Changing Sky

- Rubin will detect up to **10 million changes per night** in the southern sky
- Rubin's nightly alerts will uncover **rare events** and enable **global follow-up**
- More than 800 images for each area of sky over 10 years will reveal **changes across all time scales**



# Also the Changing Sky


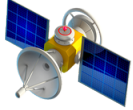






# Impacts to Rubin LSST science

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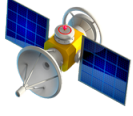




- Depends on **imperfectly known** quantities:
  - Number of satellites and their orbits 
  - The satellite brightness distribution 
  - How the LSST Camera responds 
  - Image processing algorithms and pipelines 



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  - The satellite brightness distribution 
  - How the LSST Camera responds 
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- **Rubin's potential for discovery is also its vulnerability to satellites**



- Huge field of view means low pixel loss fraction



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  - Number of satellites and their orbits 
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  - How the LSST Camera responds 
  - Image processing algorithms and pipelines 
- Rubin's potential for **discovery** is also its **vulnerability to satellites**
- Huge field of view means low pixel loss fraction

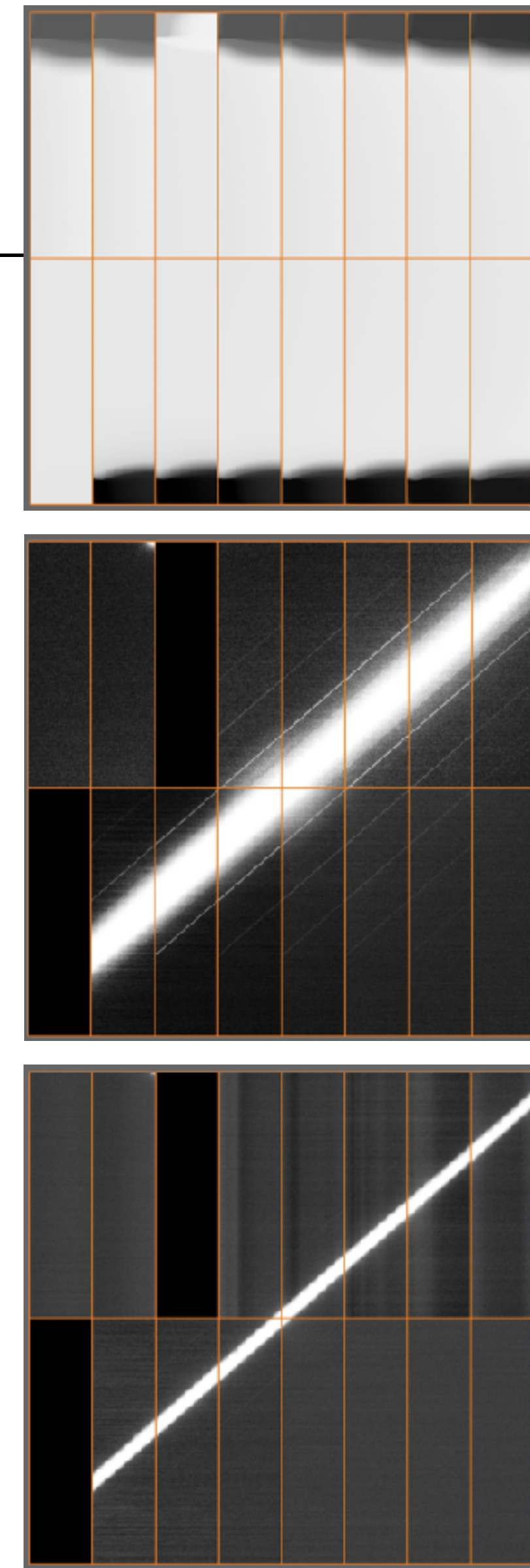


0–1 mag  
*BlueBird*

4–5 mag  
*Starlink DTC*  
*Guowang*

6–7 mag  
*Starlink*  
*Leo (Kuiper)*  
*Qianfan\**  
*OneWeb\**

*\*higher altitude ~1000 km*

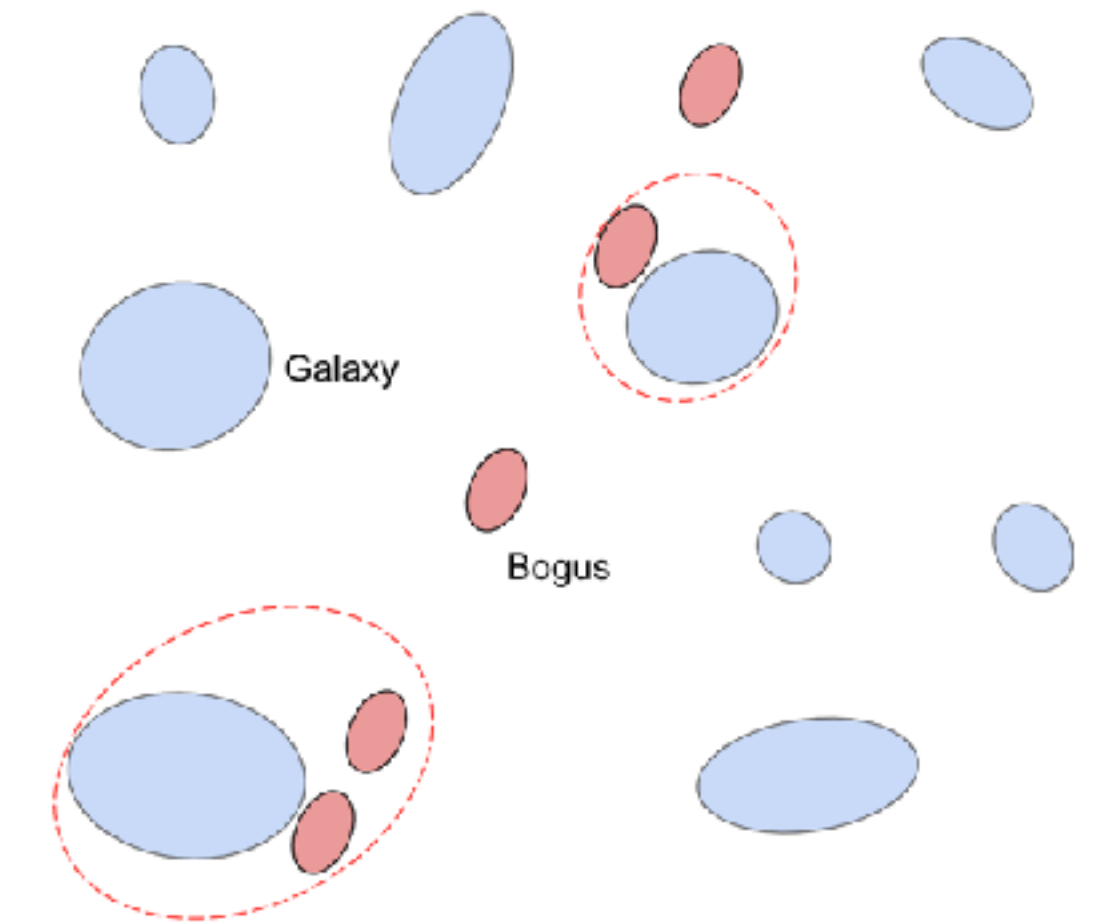
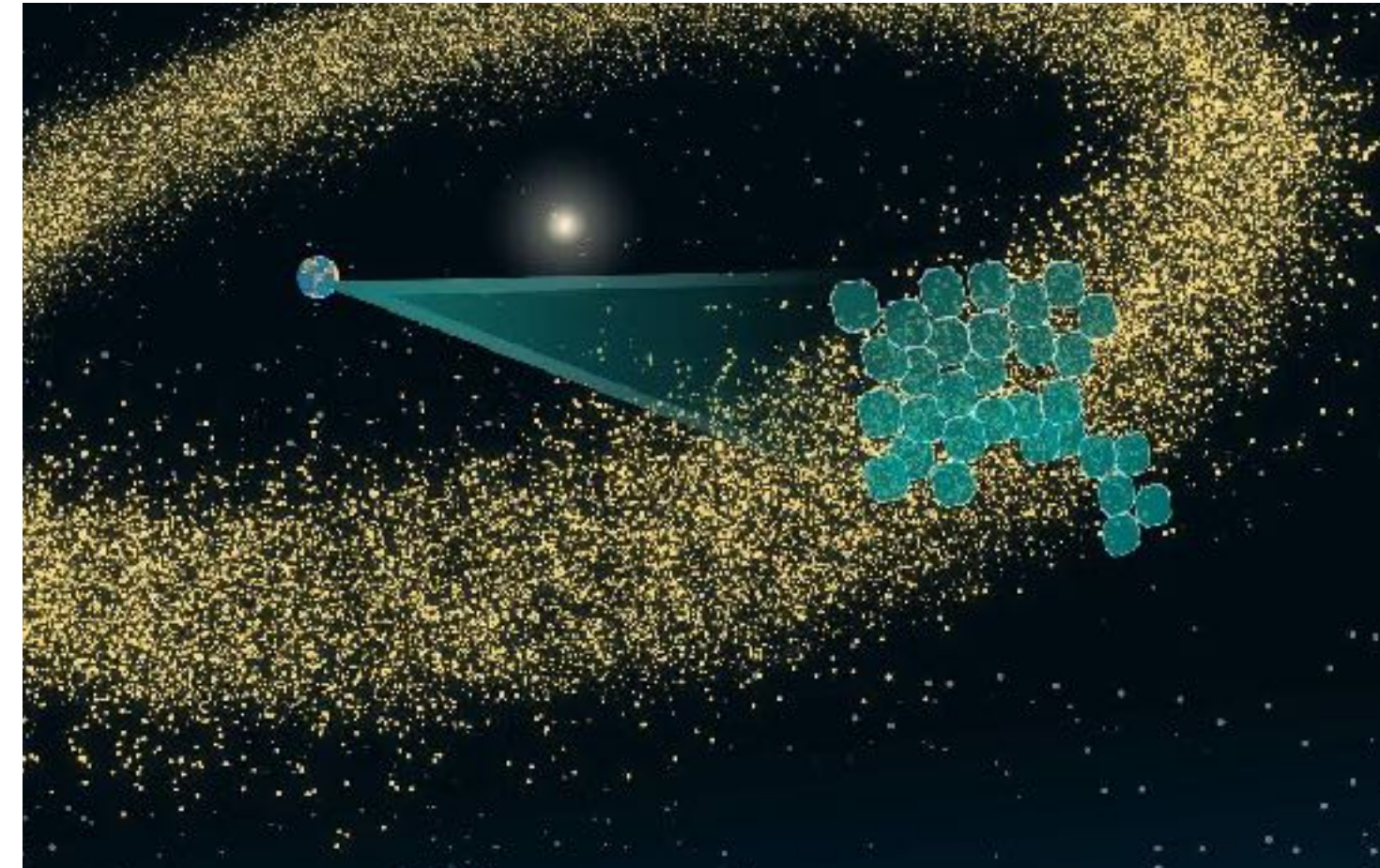




# Fewer discoveries, more systematic errors



- **Solar System**  
Fewer asteroids discovered
- **Cosmology**  
Errors from spurious signals

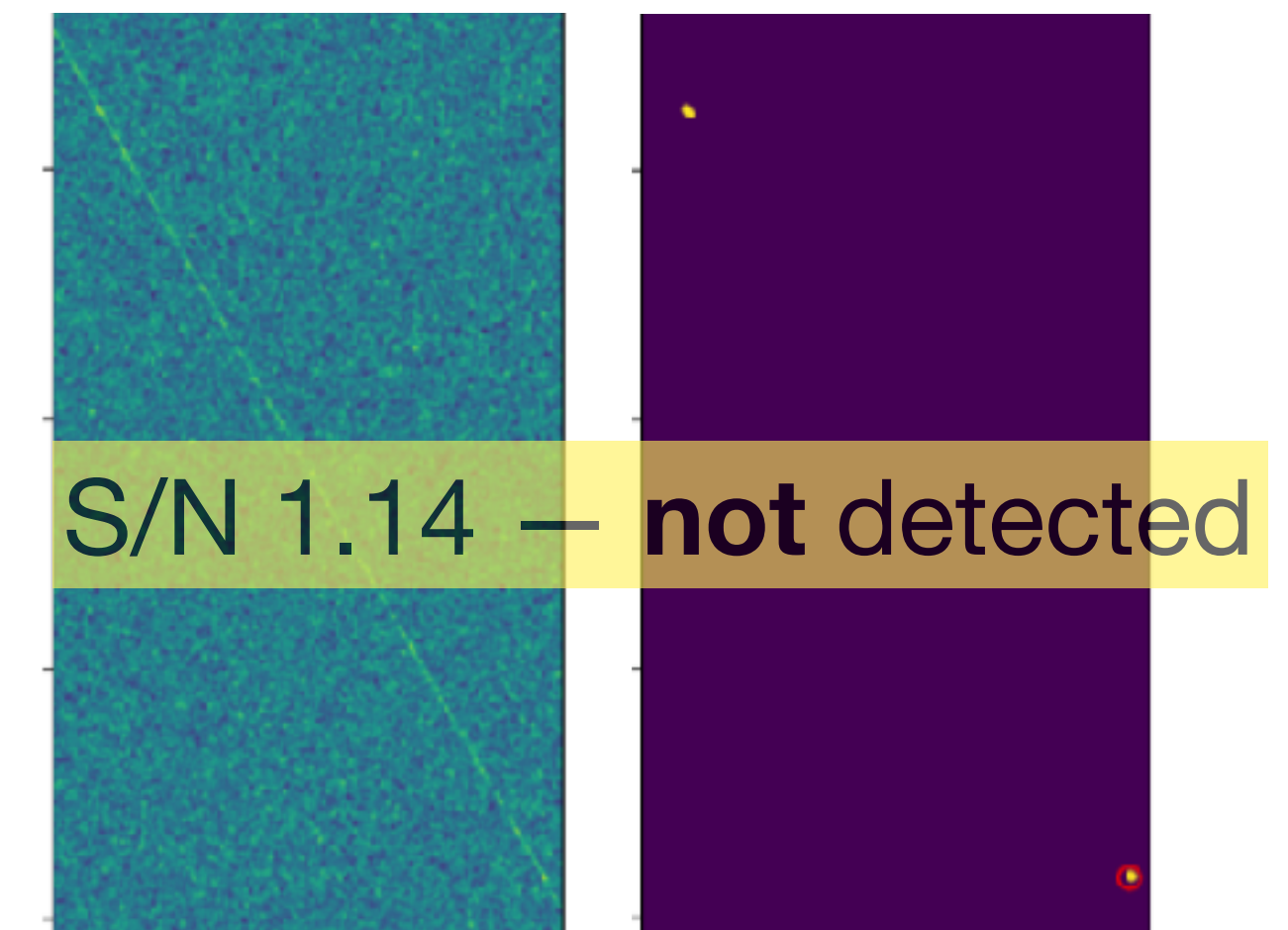
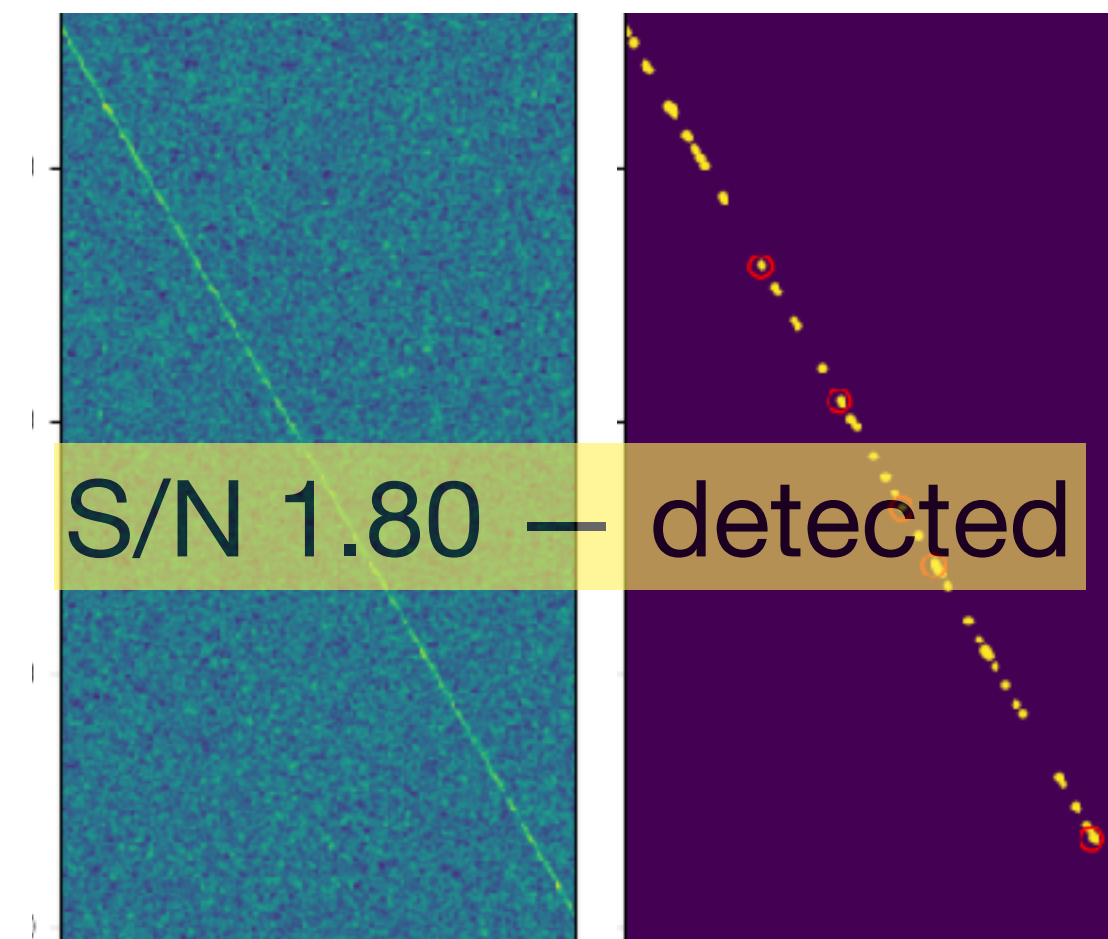
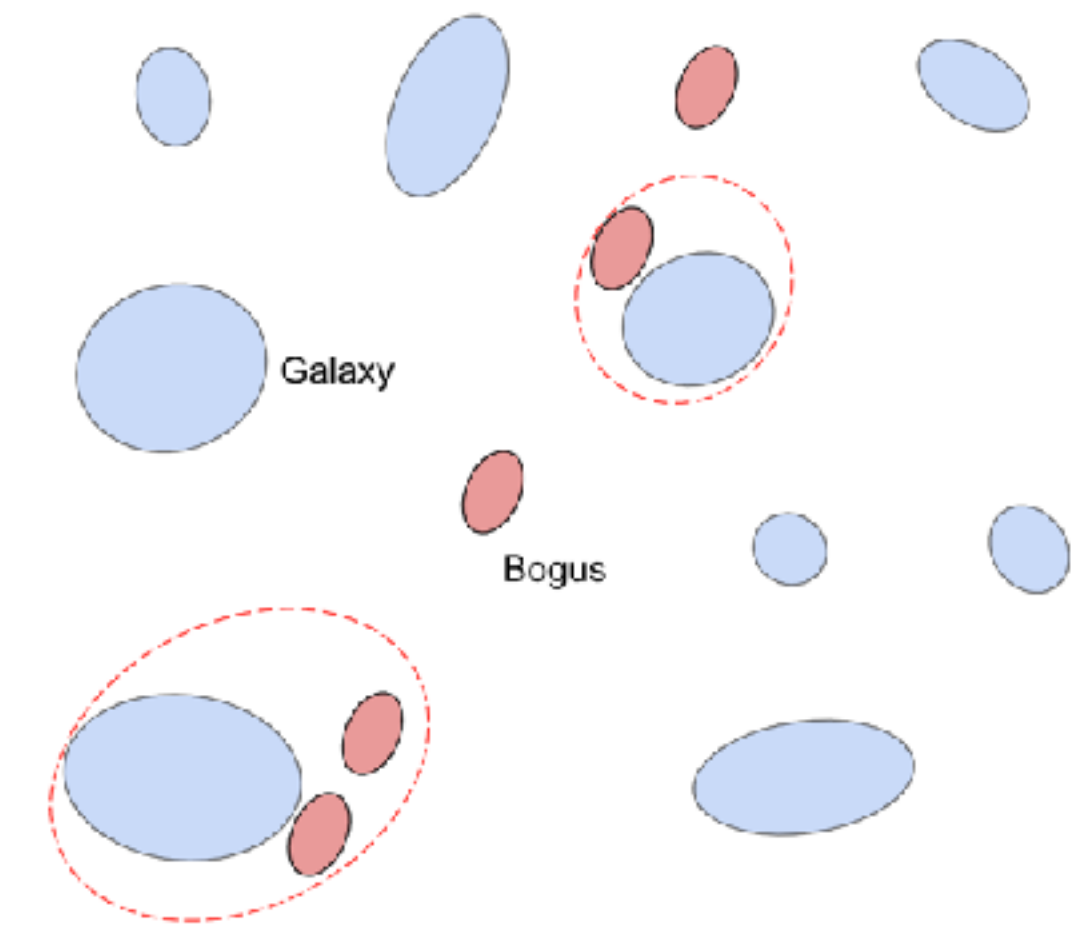
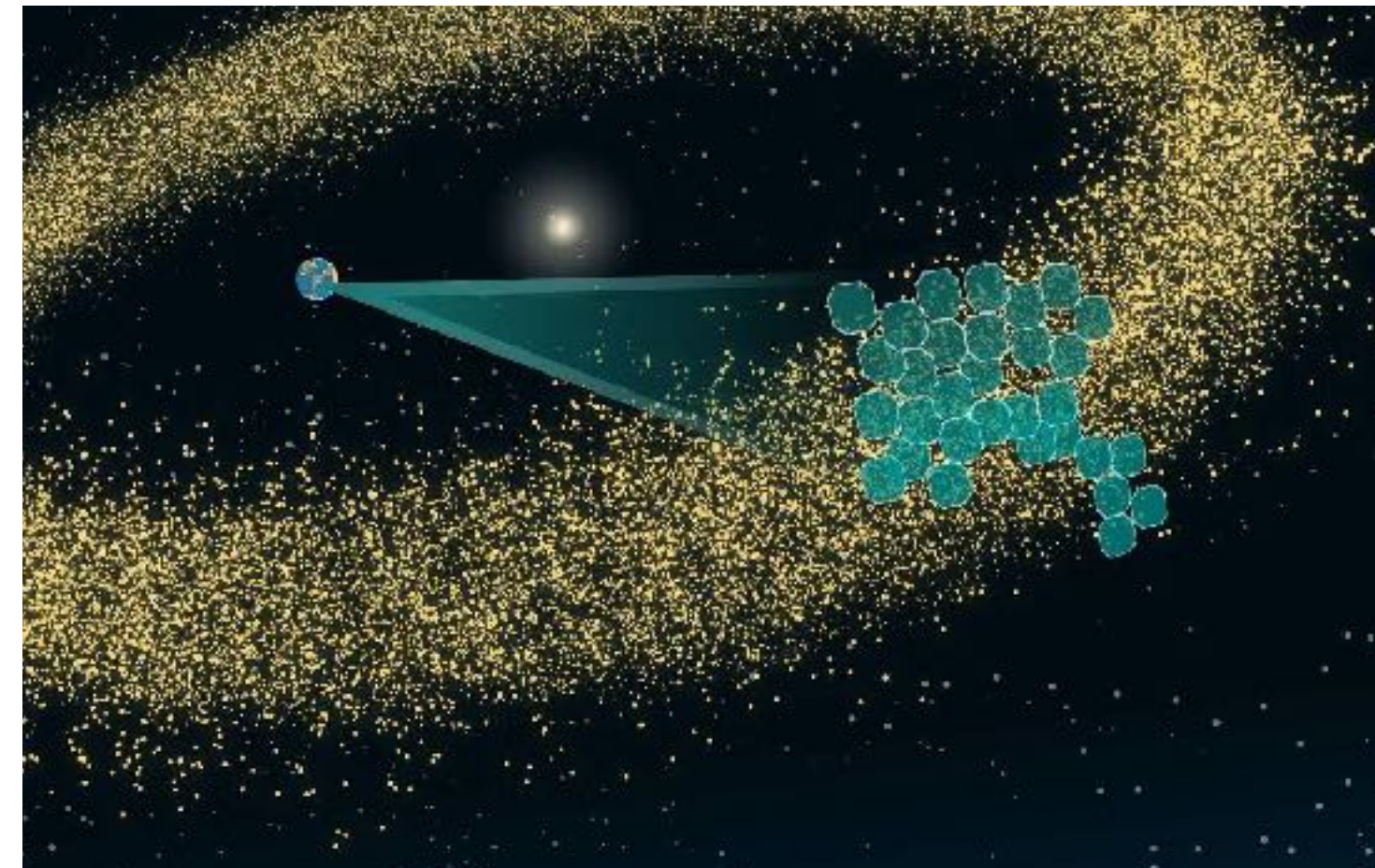




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- **Faint Universe**  
Susceptible to contamination

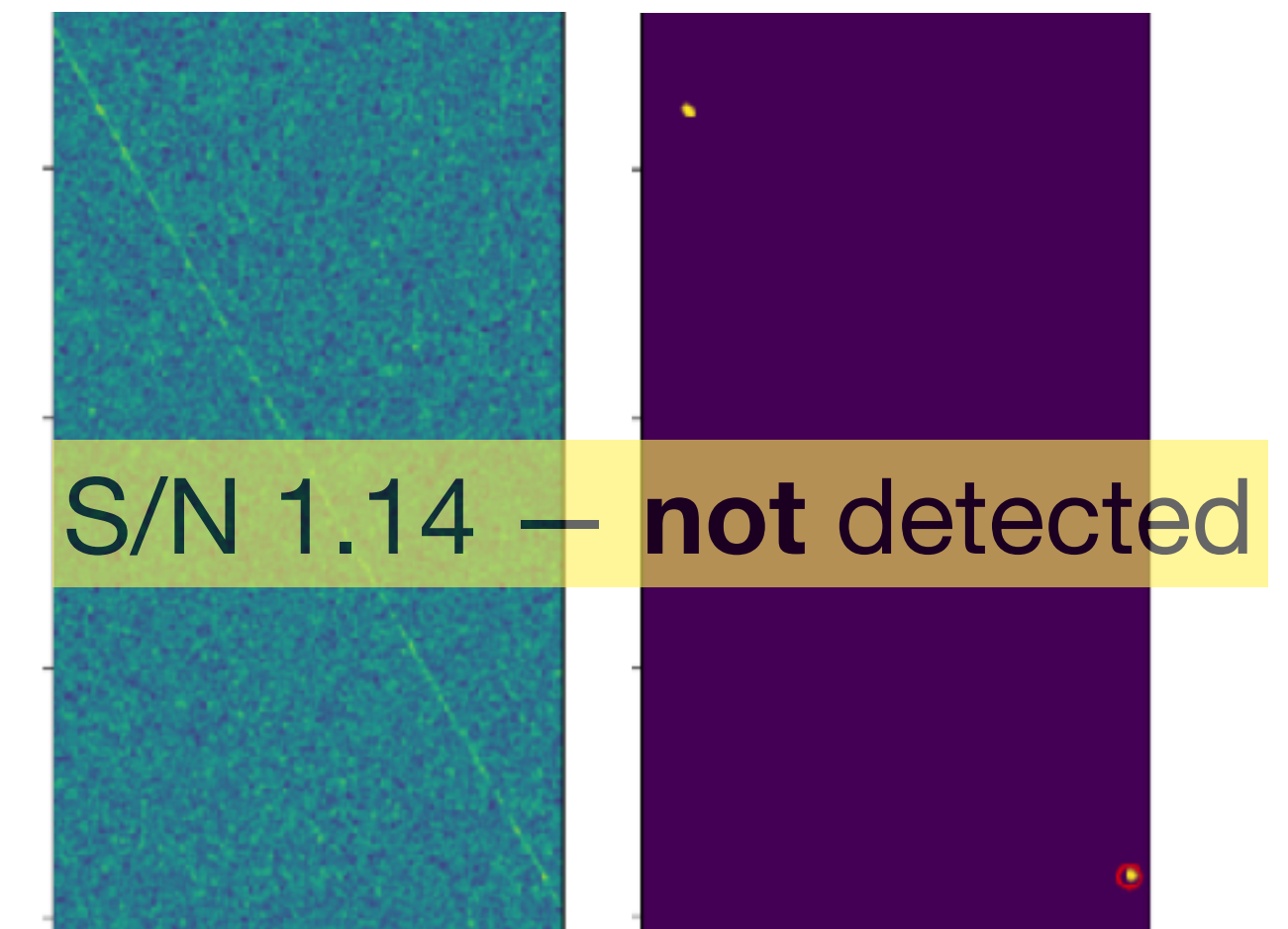
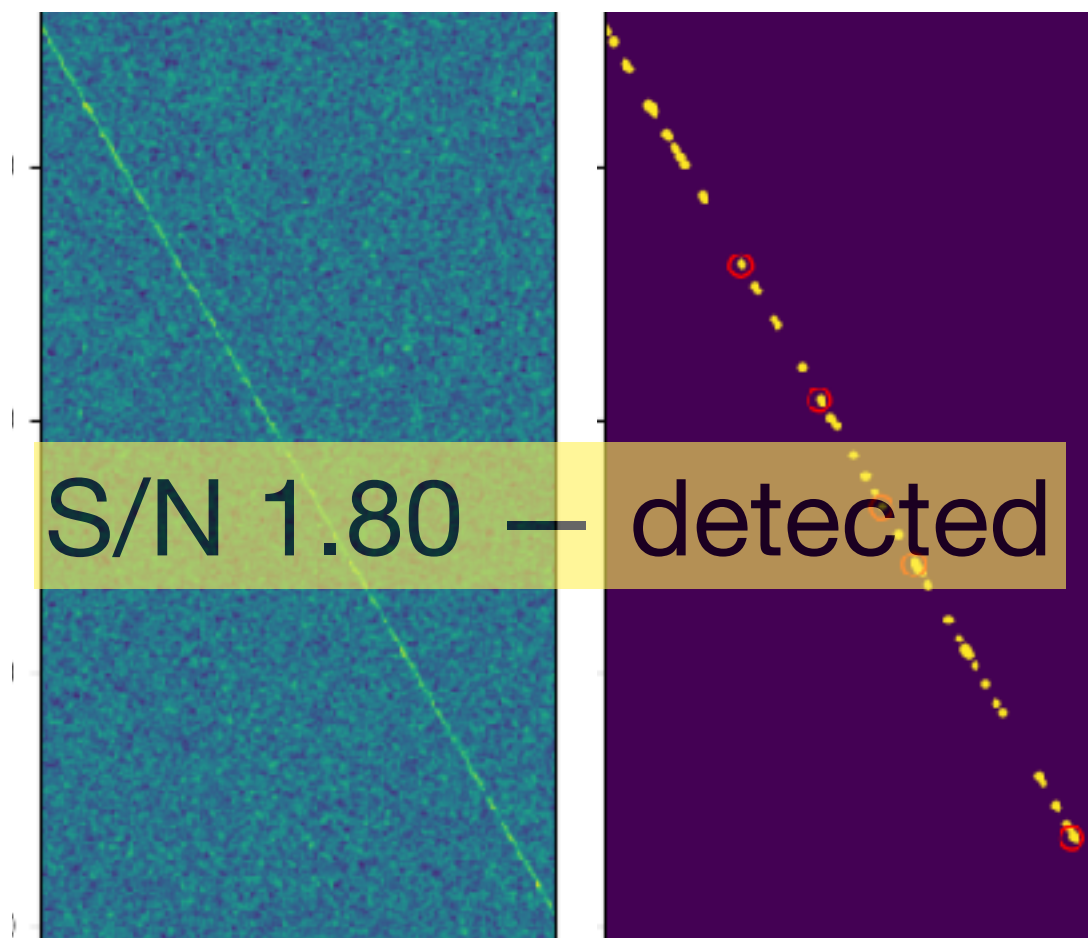
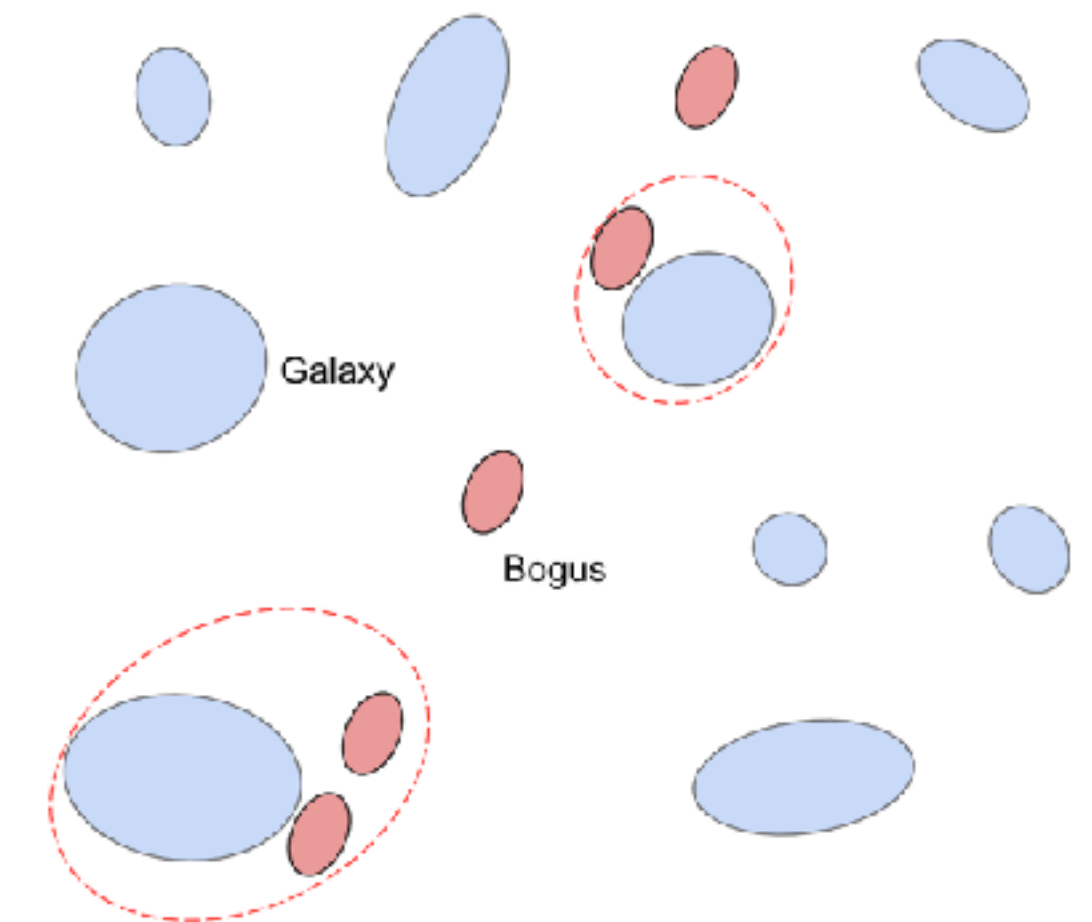
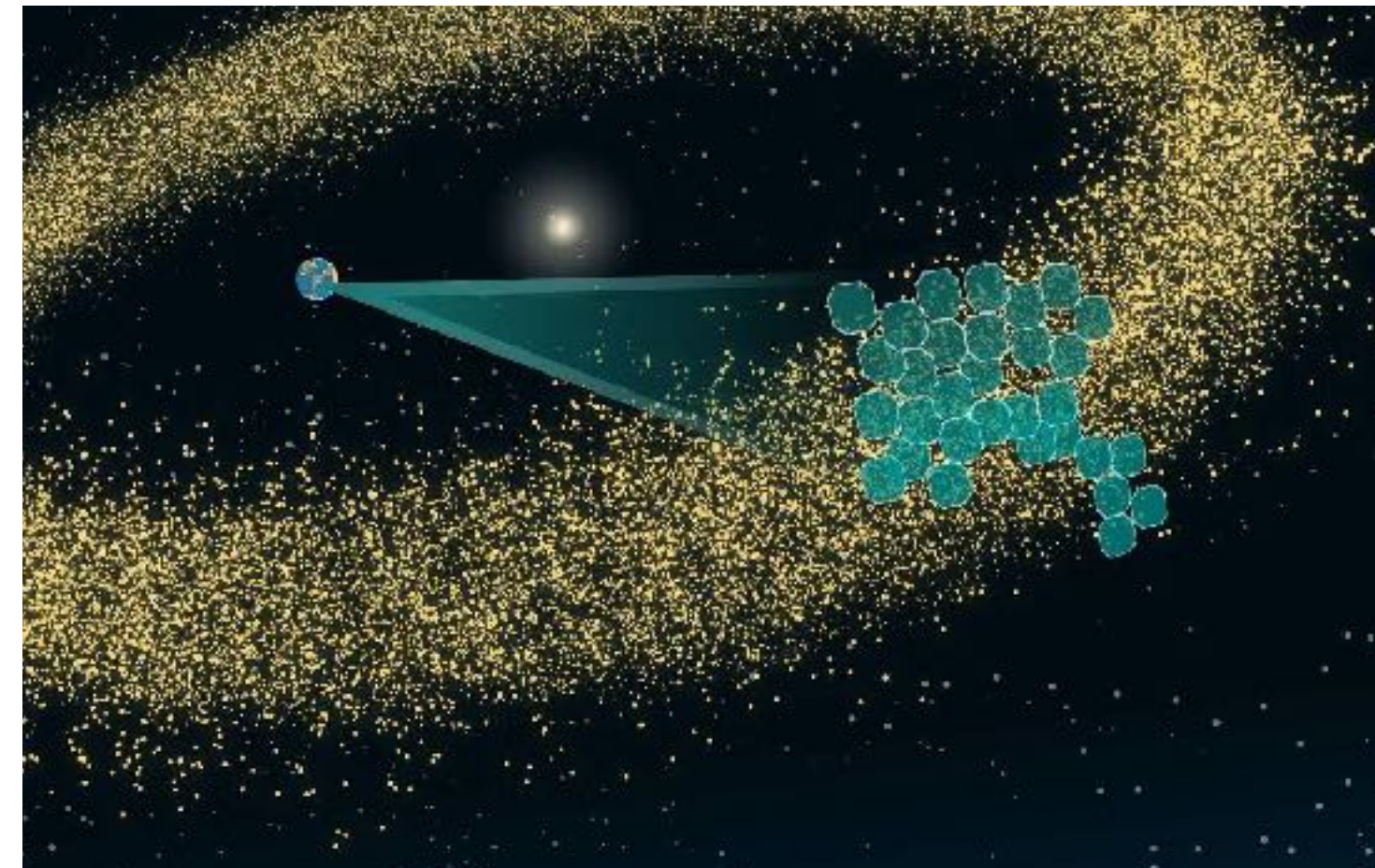




# Fewer discoveries, more systematic errors



- **Solar System**  
Fewer asteroids discovered
- **Cosmology**  
Errors from spurious signals
- **Faint Universe**  
Susceptible to contamination
- **All science**  
Need more time to complete LSST





# How to maximize Rubin science?



- *Already happening, still improving*
  - **Reject outliers** when building deep image stacks
  - Exclude known satellites and long trails from **Alerts**
  - Find and **label streaks and glints** in difference images

*Credit: Alysha Shugart*





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  - Find and **label streaks and glints** in difference images
- *Goals for Rubin's decade of operations*
  - Use the LSST scheduler to **avoid observing** satellites or trigger operator **attitude adjustments**
  - Quantify impacts from faint streaks or sky background brightness **below the detection threshold**
  - Build a database of **all** satellite signatures in LSST

Credit: Alysha Shugart



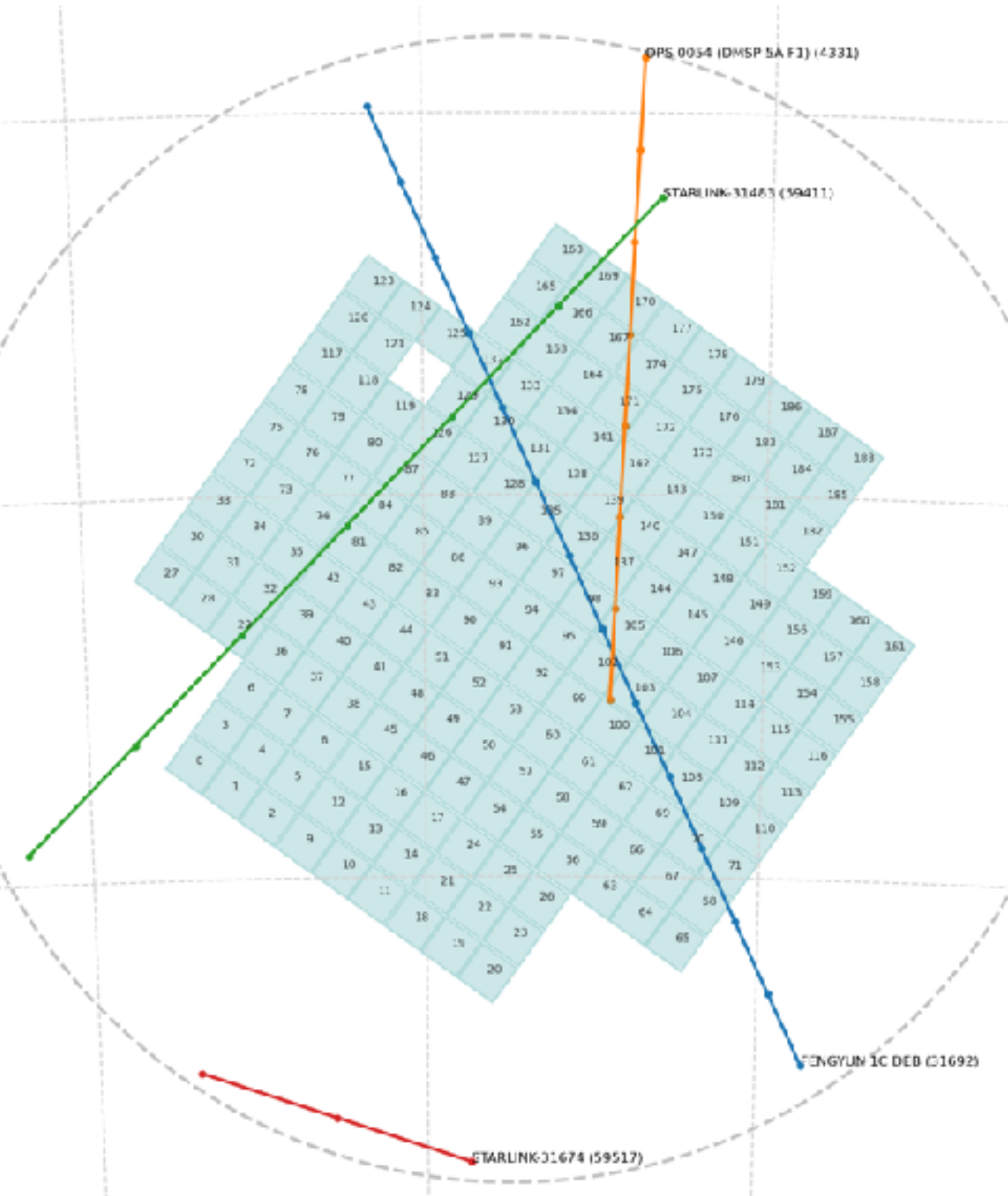
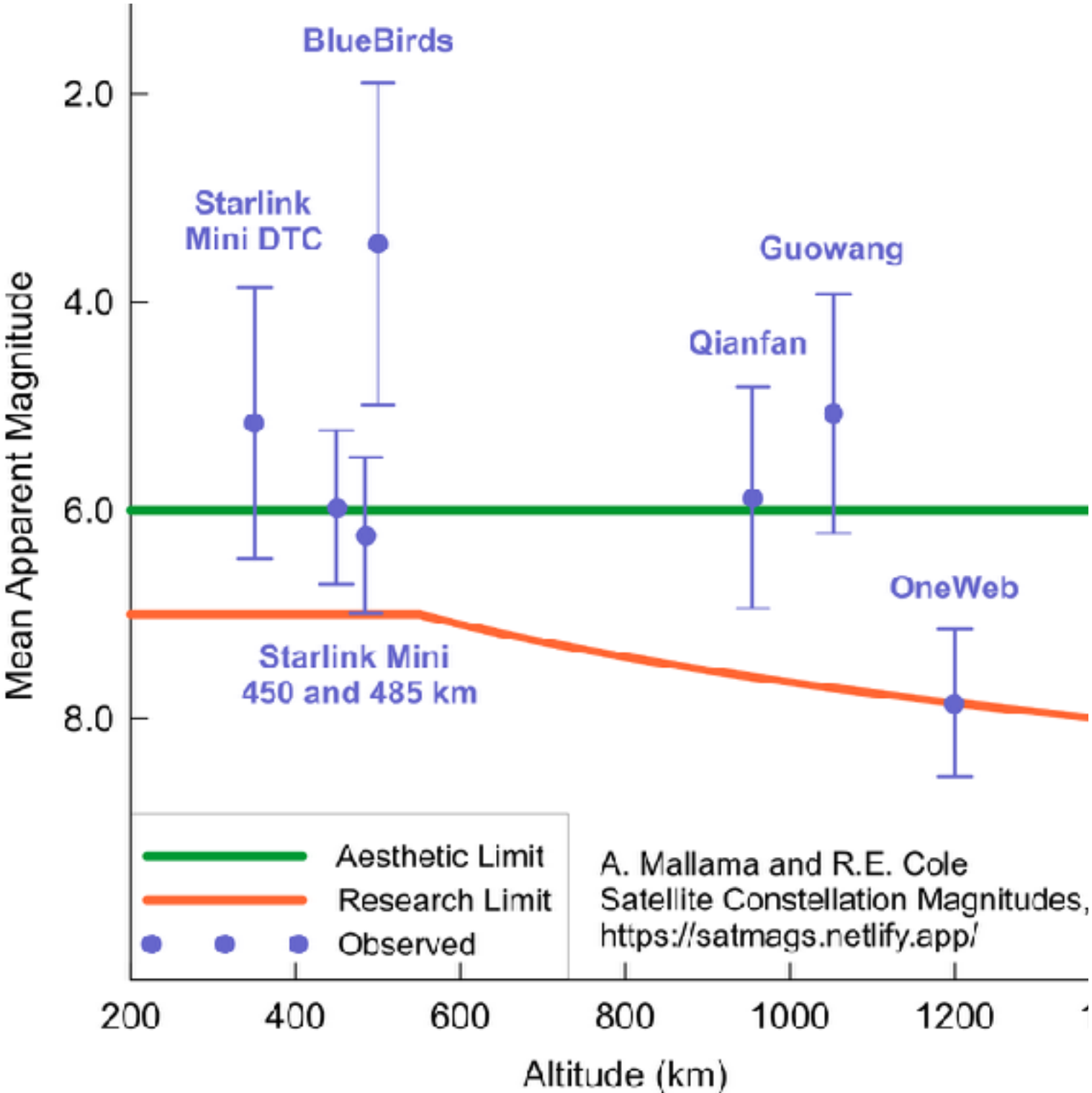
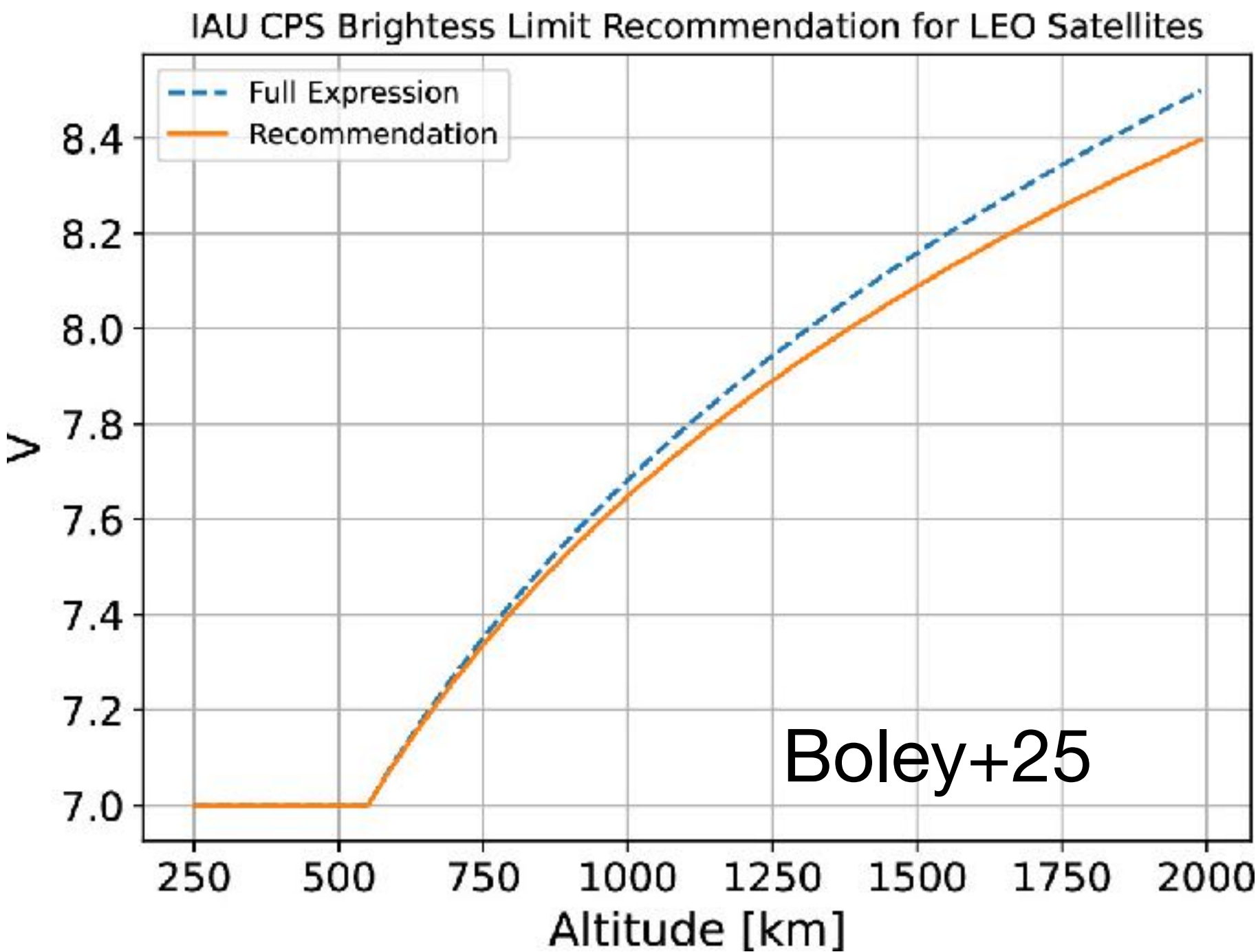
2025 Aug **Workshop Report**  
<https://arxiv.org/abs/2509.19758>



# No observatory can do this work alone



- **SatHub** at the IAU CPS coordinates satellite observations, provides feedback to operators, shares technical expertise, builds software, and develops recommendations



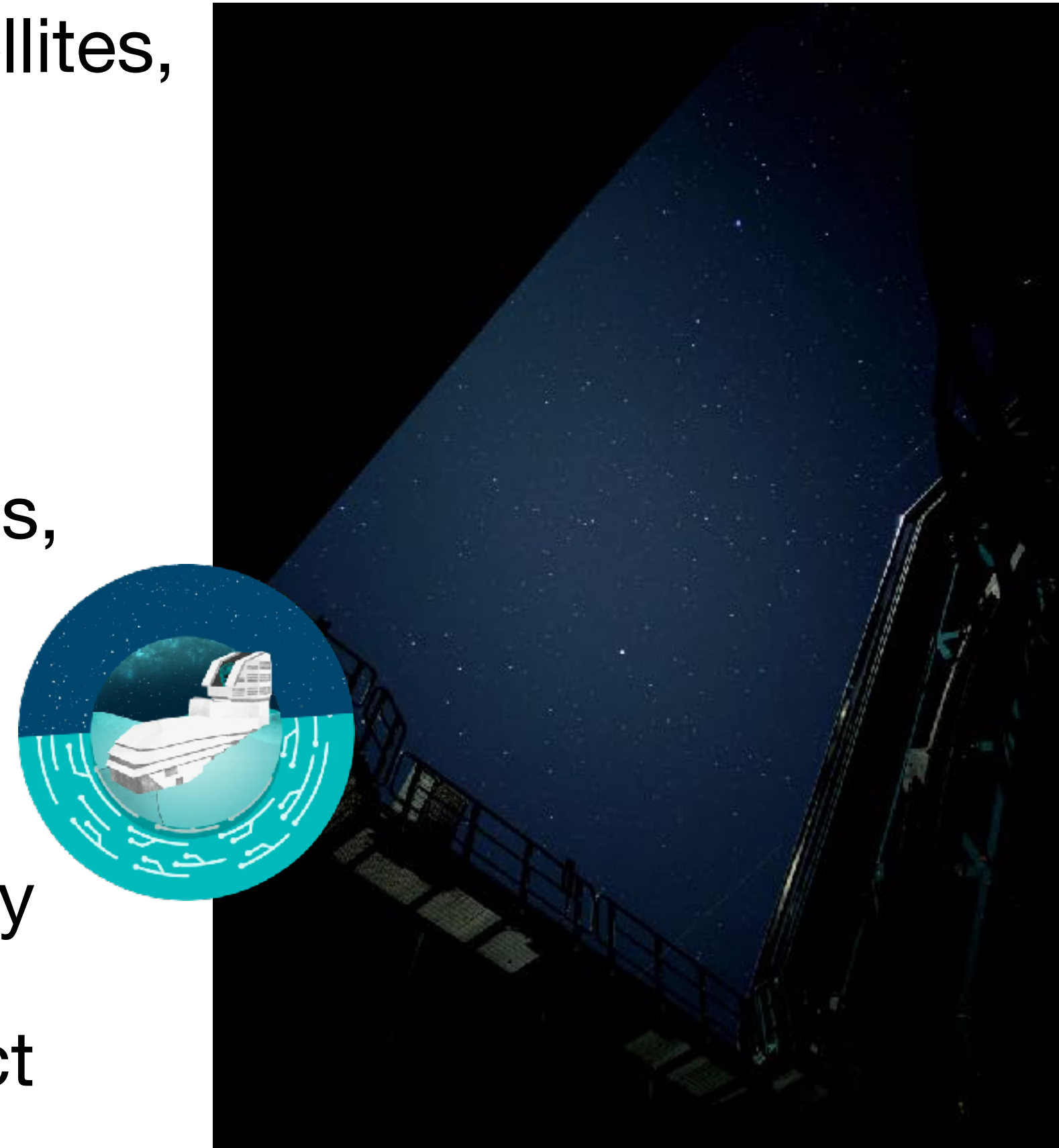
[satchecker.readthedocs.io](https://satchecker.readthedocs.io)



# Welcome to Rubin's Operations Era ✨



- Rubin's ten-year LSST will see lots of satellites, but will still do ground-breaking science
- To ensure this, **operators** can:
  - Share timely ephemerides, coordinate launches and de-orbits, darken satellites, explore hardware sharing, collaborate
- To ensure this, **astronomers** can:
  - Develop metrics and standards, quantify impacts to science, catalog observed satellite signatures to faint levels, expect pernicious contamination, collaborate
- The IAU CPS welcomes you: [cps.iau.org](https://cps.iau.org)



*Credit: Alexandre Boucaud*

## Astronomers' favorite satellites

Fewer  
Smaller  
Darker  
Lower\*  
Predictable  
Long-lived  
Shared

*\*it's complicated*