

Agenda

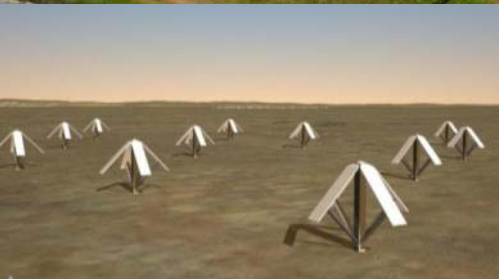
Session		Presenter	Starting Time	Time Allocated
Tuesday, 19 April 2011				
	Panel members arrive at Schiphol			
	Lunch		12:00 – 13:00	1 hour
1	Review Panel (closed meeting)	Reviewers	13:00 – 13:30	30 minutes
	Welcome and General Overview of SKA	RTS	13:30 – 13:50	
	General Overview of AA	AvA	13:50 -14:05	15 minutes
	Purpose and Context of the AA CoDR	KC	14:05 – 14:20	15 minutes
	AA Science Case	SR	14:20 – 14:50	30 minutes
	AA System Requirements	AG	14:50 – 15:20	30 minutes
	Coffee Break		15:20 – 15:40	30 minutes
2	High Level System Description	AF	15:40 – 16:25	45 minutes
	Engineering: Technical AA Concepts	JGbdV	16:25 – 17:15	50 minutes
	AA Risks and Risk Management	AvE	17:15 – 17:30	15 minutes
	Review Panel (closed meeting)	Reviewers	17:30 – 18:00	30 minutes
	Drinks and Dinner		19:00	
Wednesday, 20 April 2011				
3	Deployment and Operations	AF	09:00 – 09:45	45 minutes
	Strategy to proceed to Next Phase	AF/ AvE	09:45 – 10:15	30 minutes
	Coffee Break		10:15 – 10:45	30 minutes
	Summary of AA for SKA	AvA	10:45 – 11:15	30 minutes
	Review Panel (closed meeting)	Reviewers	11:15 – 12:30	75 minutes
	Lunch		12:30 – 13:30	60 minutes
4	Initial feedback to AA Team	Reviewers	13:30 – 14:30	60 minutes
	Summary and closing of the review	AvA/KC	14:30 – 14:45	15 minutes

General overview ; Organizing AA's for the SKA

Arnold van Ardenne

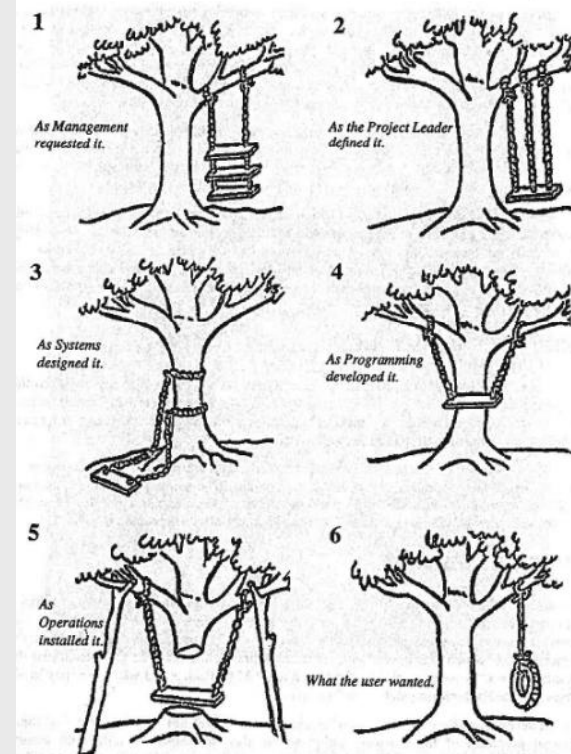
AAVP-Coordinator

ardenne@astron.nl



SKA AA System Design 2005-2012

- (Early R&D 1995- 2004)
- SKADS - EC funded SKA design Study 2005-2009
 - EC FP6 funding 2005-2009 (€ 10.0M) plus additional for Marie Curie
- AAVP 2010- 2012
 - nationally funded (Eur/Aus) supplemented by PrepSKA
 - Context provided by SKA/SPDO
 - Activities based on System design approach
- AA-Relevant Pathfinders & Demonstrators
 - LOFAR-ILT, MWA, LWA, EMBRACE , APERTIF
 - Others for e.g. SP/network purposes e.g. JIVE/UniBoard



Continuation:

AAVP after 2012 (“Pre-Construction incl. detailed design)
part of next SKA funding based on PEP

AA-low, following SKA1 BL specs

Table 3: Phase 1 Sparse Aperture Arrays

Table 3: Phase 1 Sparse Aperture Arrays		
Sparse Aperture Array		50, 180-m diameter stations
Lower Frequency	70 MHz	Dual polarization (2 orthogonal)
Upper Frequency	450 MHz	Single element covering full range
Number of antennas	11,200	Per station
Total physical aperture	$1.3 \times 10^6 \text{ m}^2$	50 Stations
Area per antenna	2.27 m^2	180 m diameter station
Dense/Sparse Transition λ^*	2.6 m (115 MHz)	A_e per element is equal to packing density.
Array Configuration	50 stations	
Core (radius <0.5 km)	~50% (25 stations)	Fractional total number of AA stations.
Inner (1< radius<2.5 km)	~20% (10 stations)	"
Mid (2.5<radius<180 km)	~30% (15 stations)	In clusters of 5 stations (Total of 15 clusters)
Core filling factor	0.81	$A_{\text{core}} = 7.9 \times 10^5 \text{ m}^2$; $A_{\text{ant}} = 6.4 \times 10^5 \text{ m}^2$
Number of beams	480	
Instantaneous bandwidth per beam	380 MHz	Assumes full bandwidth is available (70-450 MHz)
Digital Outputs		
Sample streams	960	Max - sub-bands
bits per sample	4	

Science reference : DRM version 1.3

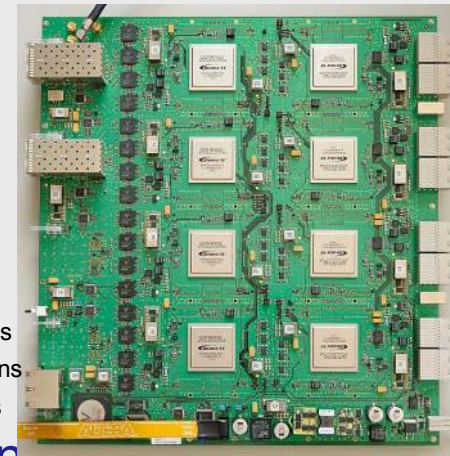
Next steps: detailed technical requirements/design explorations

AAVS1: AA-low

- Demonstrate electromagnetic and front end performance
 - Sufficient collecting area (250-500m²?), possibly two versions
 - Focus on **broadband** antenna and front-end design
 - Evaluation using existing processing
 - UNIBOARD, ROACH or LOFAR / MWA

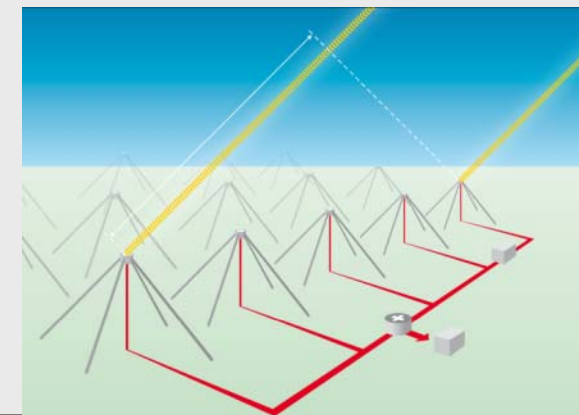


- 14 Layers
- 340x366mm
- 7304 components
- 25798 connections
- 271 meter traces



- Goal: establish antenna, tile and station configuration
 - Extensive station simulations required

- Concept Des. Review 19/20-04-2011
- Prel. Design Review 31-12-2011
- Commissioning 31-12-2012
- Reporting 30-6-2013



Int. Murchison Widefield Array

Key players

- AUS/ICRAR, CSIRO
- US/MIT-Hayst,, e.o.
- India/RRI



LOFAR Lessons learned Areas vs industrial participation

Courtesy: M. van Haarlem

- Procurement Rounds
- Supervising Mass Production
- Producing High Band Antenna
- Station Construction



and in parallel

The Advanced Instrumentation Program

- Further development of innovative wide-field “*radio camera*” technologies with the potential to enhance Phase 1 and be a major part of Phase 2:
 - dense aperture array (FoV $\sim 200 \text{ deg}^2$)
 - phased array feeds (PAFs) on the dishes (FoV $\sim 30 \text{ deg}^2$)
- Development of ultrawideband feeds
- Enhancing SKA₁ and being a major part of SKA₂ ; tbd 2014

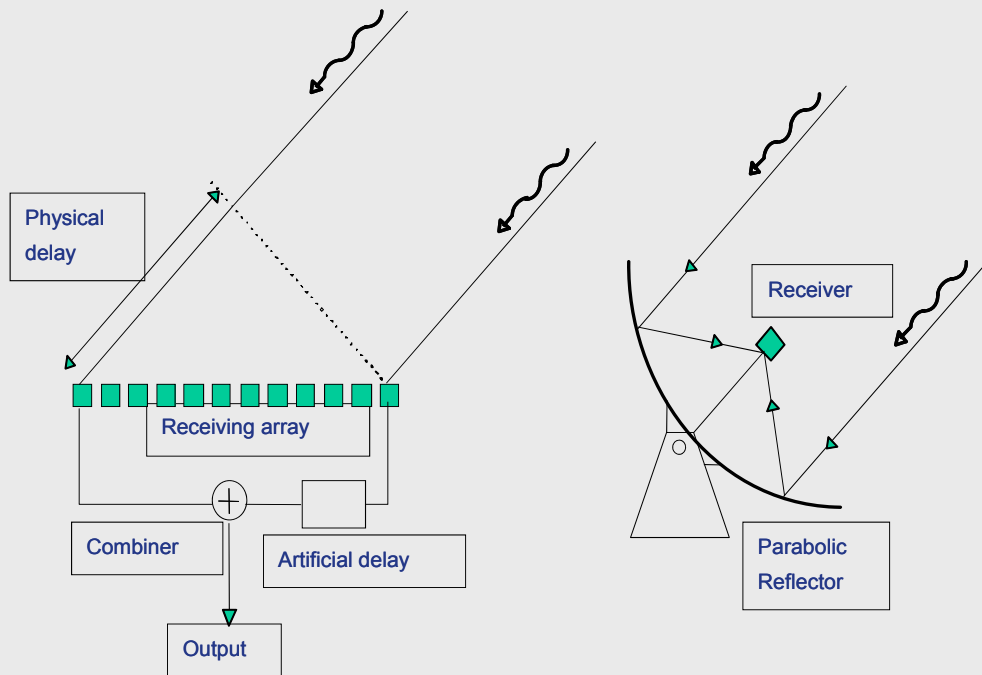
Decision in 2016

Draft specs for AA_MID/AIP

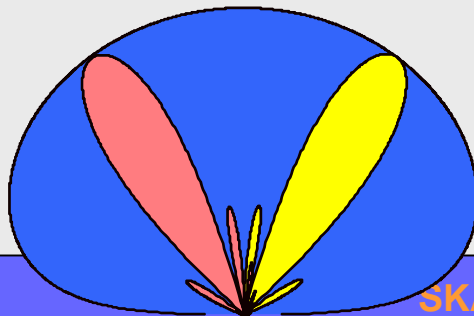
AAVS			Remark
	A_{eff}	~2000 m ²	Effective area
Front-end	$A_{\text{eff}} / T_{\text{sys}}$	40 m ² /K	0.5 – 1.2 GHz, at broadside
	Array distribution	14 stations	2D configuration, baseline < 1km Indicative, uv coverage calculation will determine number of stations
		0.4 – 1.4 GHz	Smooth degradation for 1.2 – 1.4 GHz $A_{\text{eff},1.45\text{GHz}} > 0.5 A_{\text{eff},1.2\text{GHz}}$ No Grating lobes < 1,45GHz (< 45° scan)
	Inst. Bandwidth	350MHz	
	Array type	Dense	Close packed, Nyquist sample incoming wave front, approx. constant A_{eff} , best dynamic range
	System Temp	<50K	@ 1GHz
	Scan angle	±45°	
	Beamformer	Phase and time delay ≤ 45° steps	Analogue beamformer at tile level
	Polarisations	2 linear	
	Sidelobes	<12 dBc	With uniform taper, other tapers possible with implications for A_{eff}
Processing	<i>Substation beamforming</i>		
	Inputs	128 x 4	Assumes 14 substations of 128m ² , each square meter produces 1 RF output, Dual pol, 2 FOVs
	No. beams	32	Generated within each FoV
	Freq. channels	16.000	
	Correlator		For 14 substations, full Stokes, 2 pol
	Imager		
System	Phase/amplitude stability		
	Sensitivity		

For PDR: detailed technical requirements/design explorations

Principles widefield Radiotelescopes for SKA



LOFAR (dipool-antennes)



Exloo (NL)

Onsala (Sw)

Potsdam (D)

Tautenburg (D)

Effelsberg (D)

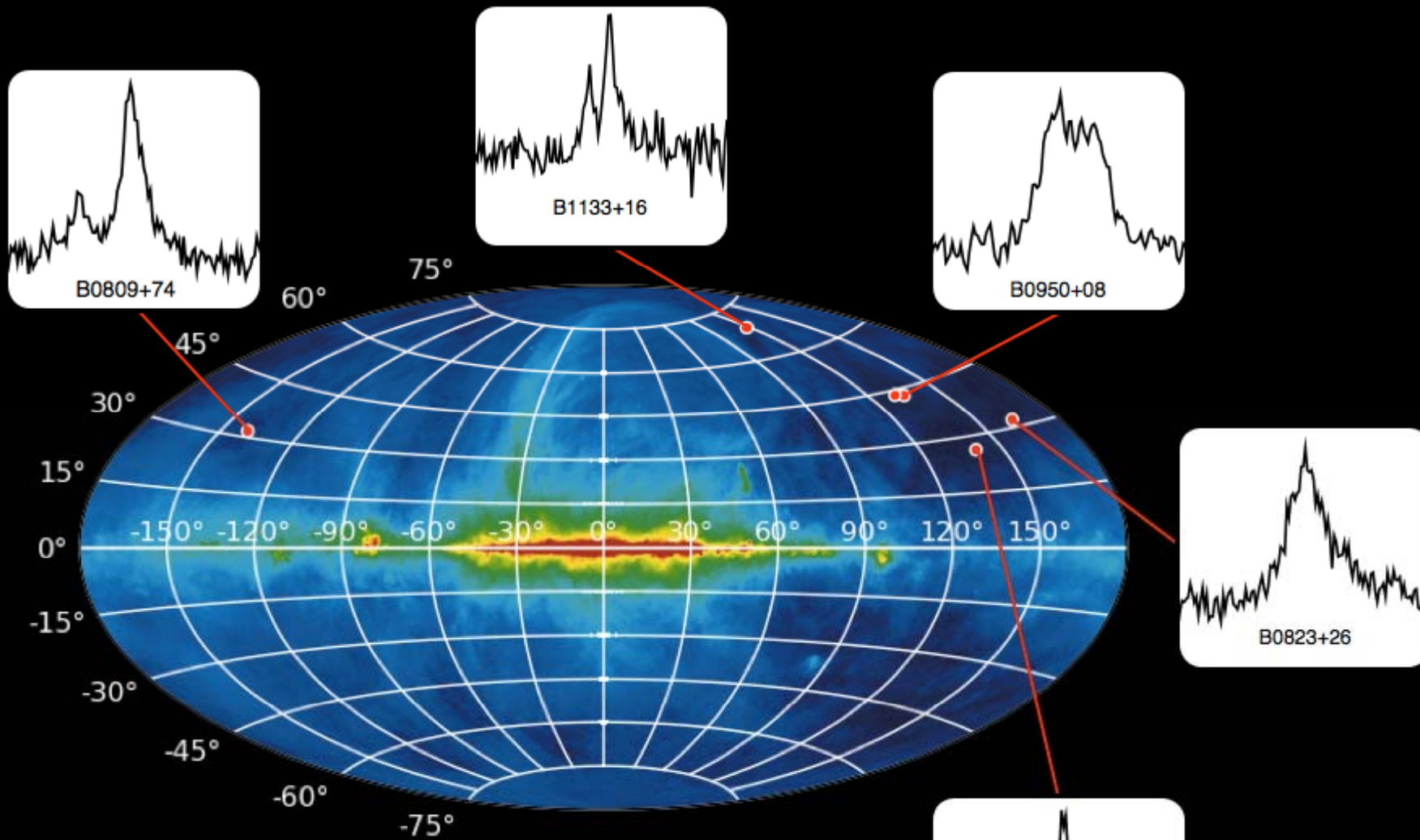
Unterweilenbach (D)

Chilbolton (UK)

Nançay (F)

**LOFAR Stations Across Europe;
70% naar NI industrie**





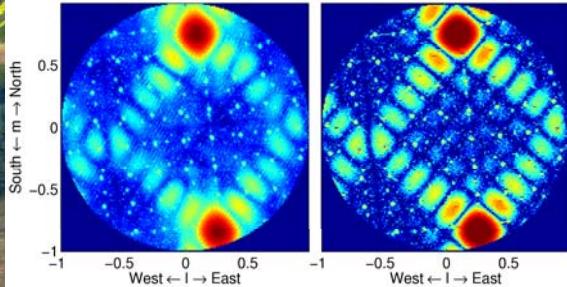
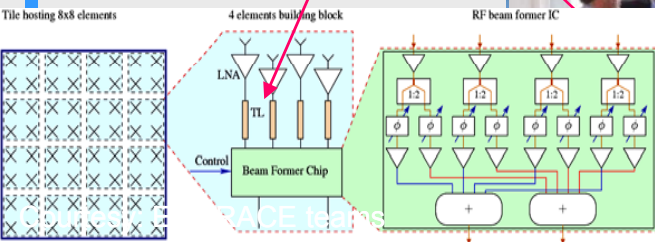
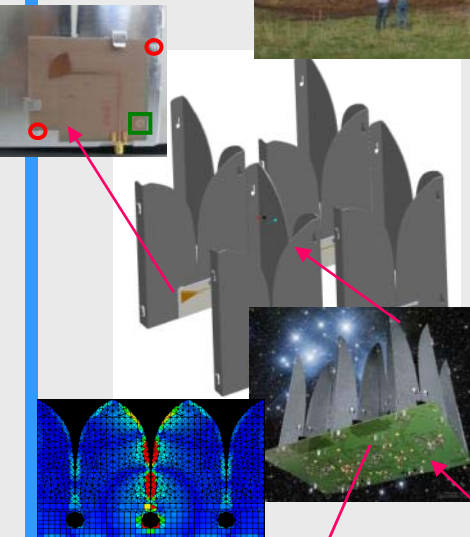
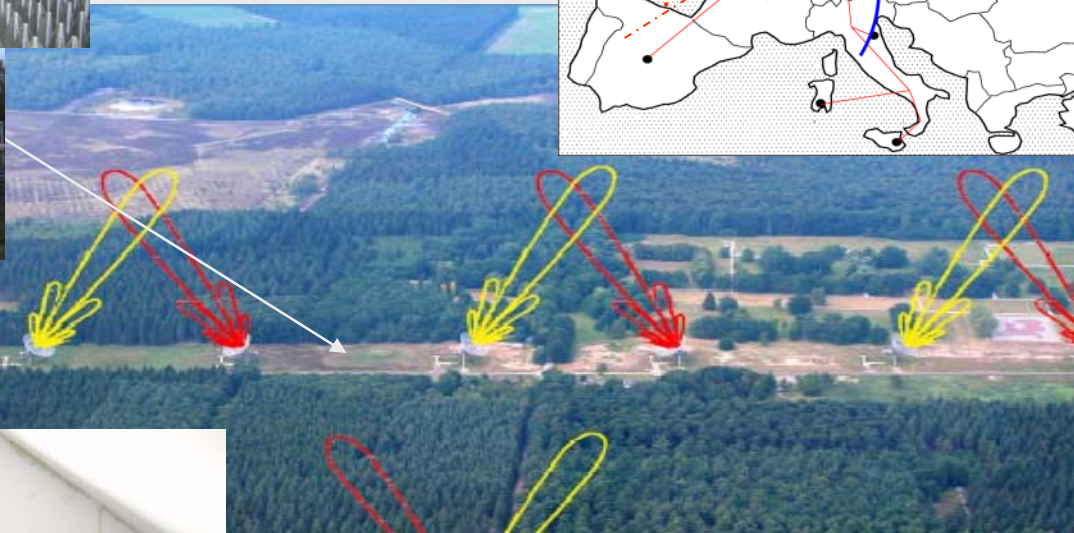
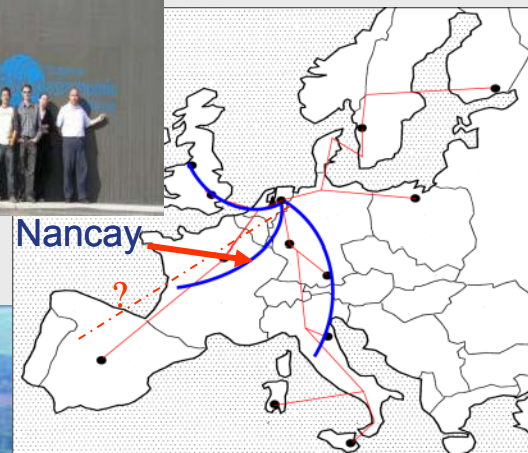
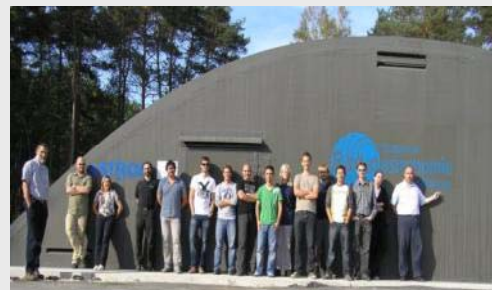
**simultaneous multi-beam
observations in the LOFAR low
band**

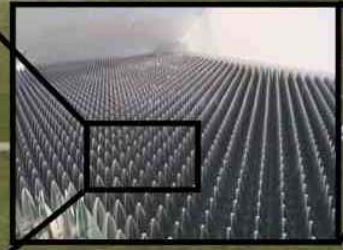
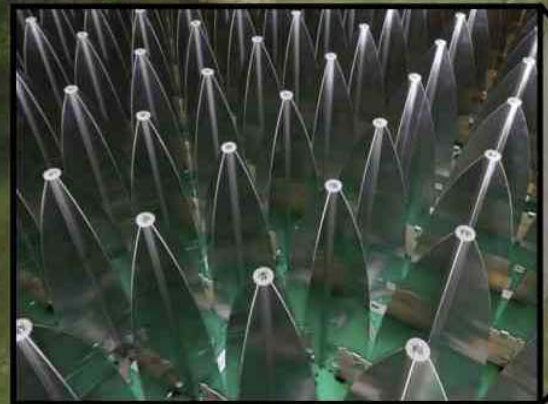
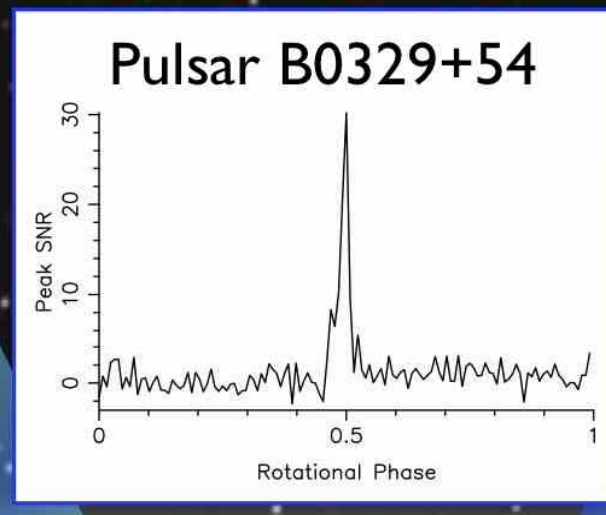
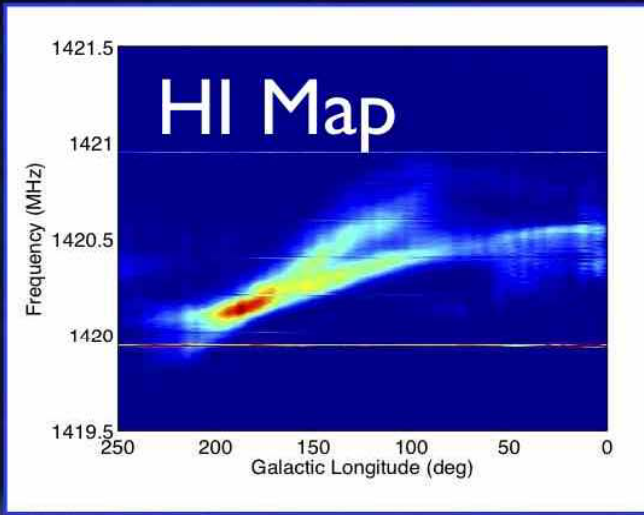
EMBRACE;

a world-first 0.5-1.5GHz Dense SKA-AA Demonstrator

FOV: ~ 140 sq. deg. / tile beam @ 1200 MHz

Courtesy: ASTRON, OPAR, INAF, MPIfR





EMBRACE Dual Beam

Principal Deliverables in time

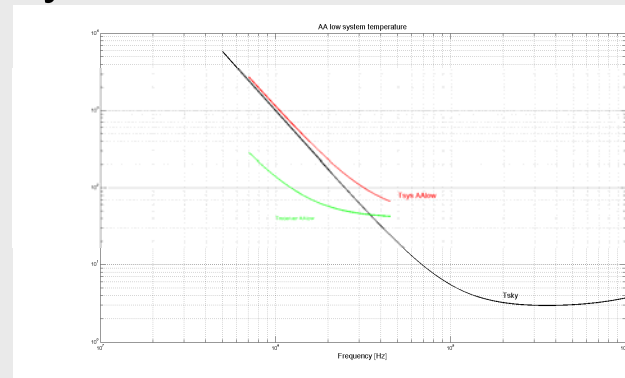
AA-low	Technol. demonstrator & Costed sub-system design	AAVS1	2013
	SKA ₁ full pre-production prototype array	AAVS2	2015
	SKA ₁ manufacturing data package		2015
AA-mid	single array performance demonstrator	AAVS1	2013
	multi-array interferometer performance demonstrator	AAVS2	2014
	SKA ₁ capable costed design		2015
AA system	AA-lo and AA-mid CoDR		2011
	SKA ₂ technology roadmap & costed system design		2015

Programmatics 2011

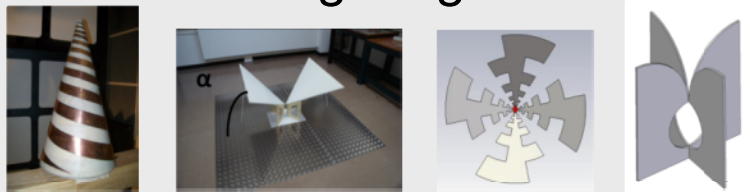
- 1Q2011: Updated AAVP project- and activity plan Updated org-diagram, Collaboration agreement being finalized
 - AAVP reports to “Board” i.e. AAVP Steering Group (met once);
 - Stakeholders responsible “per national activity”
- Appointed Project Manager and Verification Scientist
 - Derek McKay-Bukowski (PM)& Ilse van Bommel (VS)
 - Funded through PrepSKA-SPDO
- SPDO AADomain specialist 50%
 - Andre Gunst
- AA-CoDR 19-20 April follows SKA deltaSystem CoDR
- AA-low workshop held in Perth 6-9th September (ICRAR is hosting)
 - Main purpose: Preparing for Preliminary Design AA-Low review in December
- Prelim Des. Review in connection to AAVP Conference Dwingeloo
 - 12-16th December
 - PDR dates TBC
- **Note: Workshop ICT in Aveiro(Pt) 24-25 May**

AA- design issues

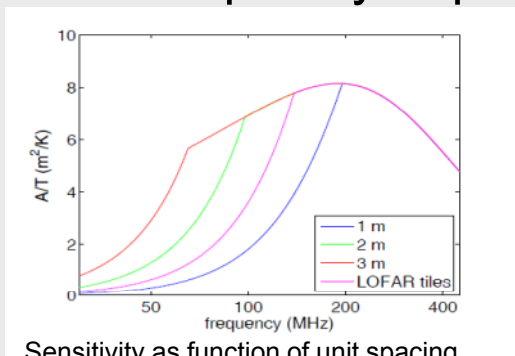
- Dealing with strong frequency dependent skynoise vs receiver noise



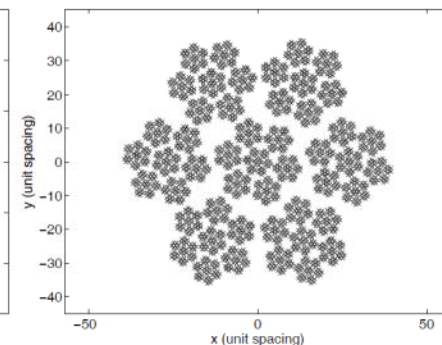
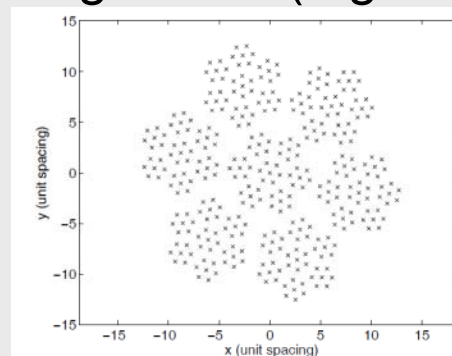
- Antenna design e.g dual vs single band



- Optimal frequency dependent Configuration (e.g. “snowflake”)



Sensitivity as function of unit spacing



- Dense vs sparse vs science requirements

Bregman Wijnholds 2011, ASTRON-RP-468
Grating lobes -20dB

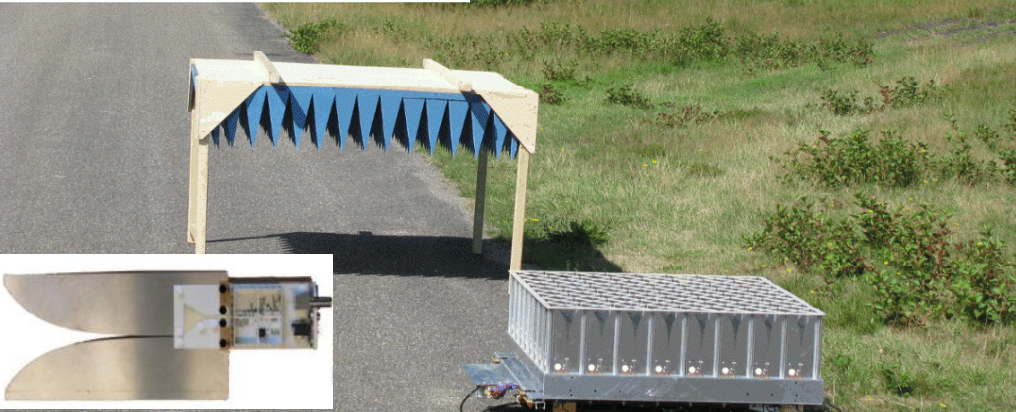
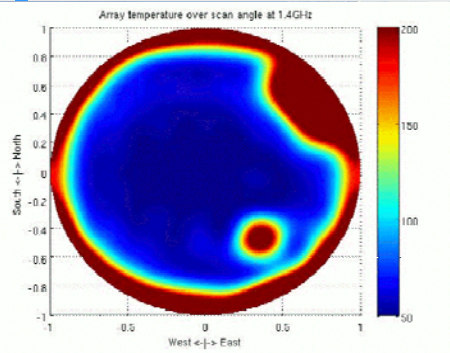
- Connected to configuration, filling factor, calibratability, e.m. design, etc. See see above.

AA's for SKA-Mid; outlook on system noise performance; <50 K T_{sys}

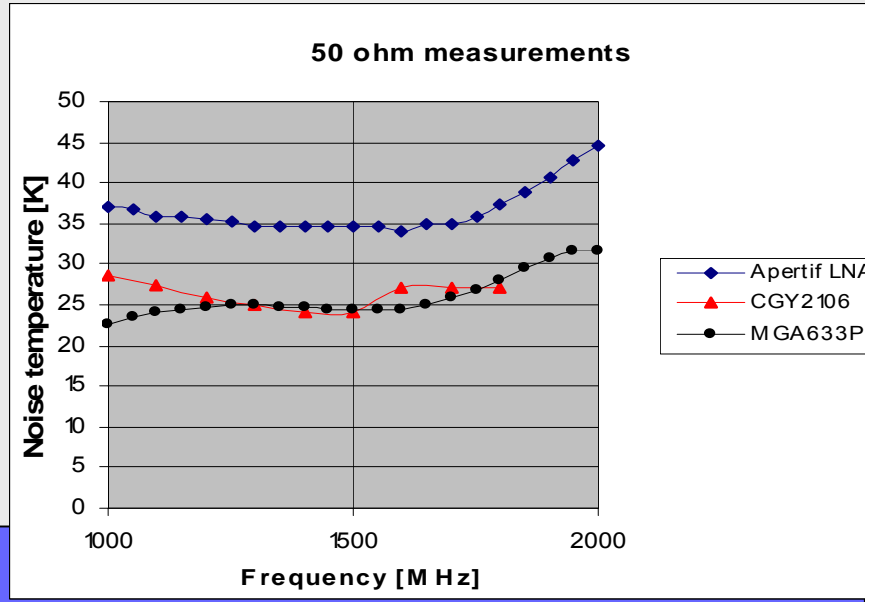
- setup for Y-factor method and zenith results
- system temperature distribution over the sky

T_{sys} noise budget @ 1400 MHz for Apertif AA-tile (49-element array, broadside)

Antenna loss	4 K	
LNA (incl. 2-nd stage)	35 K	< 25 K
Noise coupling	7 K	
Spill-over	0 K	
Sky noise	3 K	
Total	49 K	<39 K



Courtesy: Bert Woestenburg, Laurens Bakker, Wim van Cappellen



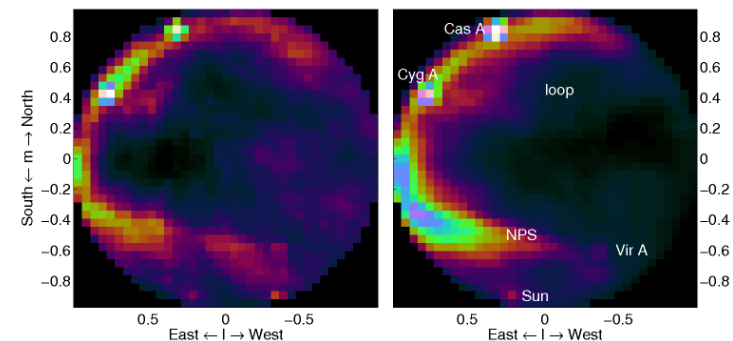
Calibration and Imaging

- Strong Calibration , Imaging and Commissioning team
 - Strong diffusion from LOFAR
 - Expanded to allow fresh thinking
 - Task lead by Keith Grainge (UCam). Core: Keith, Stefan Wijnholds (SP, algorithms, station calib), Oleg Smirnov (ME, HDR imaging), Ronald Nijboer (algorithms, math.), Tobia Carozzi (widefield polarimetric Imaging), Fred Dullich (beam sim.)
 - Associated: Jan Noordam, Parisa Noorishad, Paul Alexander, Rosie Bolton
- Complex multi-issue task
 - Dealing with frequency dependencies, configuration, computational efficiencies e.g. modelbased station calibration (N cubed vs NxNlogN) vs other approaches
 - (effects of) Multiple sources (changing over frequency), correlated noise, low SNR, propagation effects, computing resources (large data rates (Tbts, Ebps), real-time calibration)
- Interacts with SKA-Calim and co-workers
 - US, Aus, Europe
- Now drafting issue-based workplan
 - Dealing with TIDs, strong sources, mutual coupling, beamshape stability, calib. accuracy?

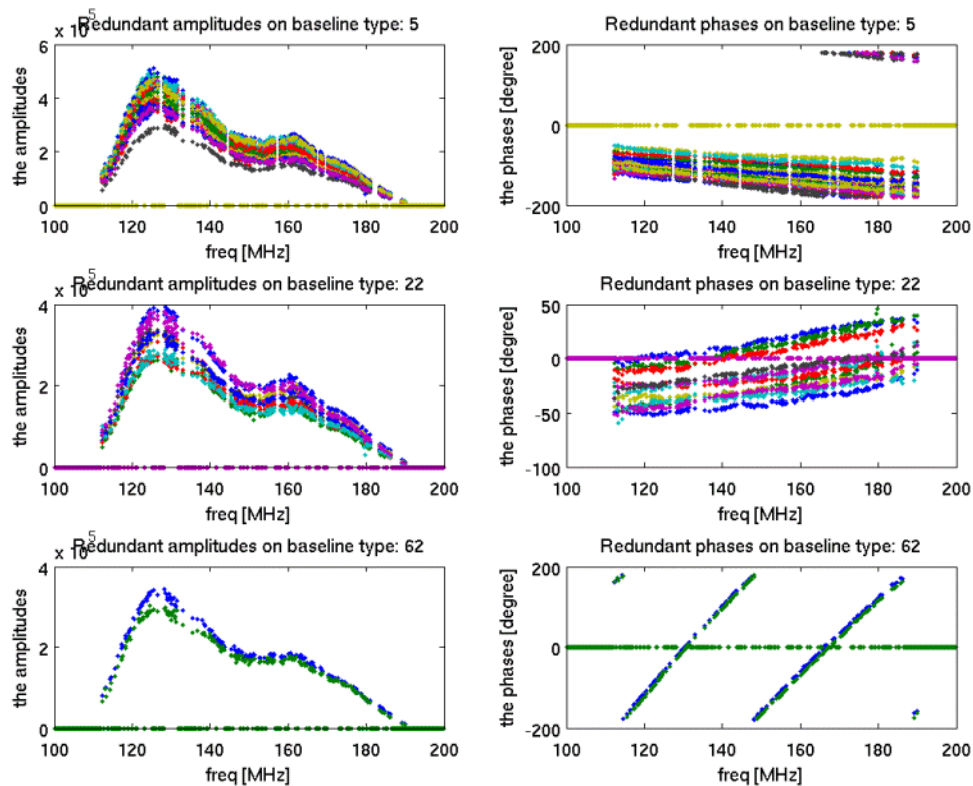
DFT vs. LS imaging (1)

S.J. Wijnholds, URSI Benelux Forum, Jun. 2009

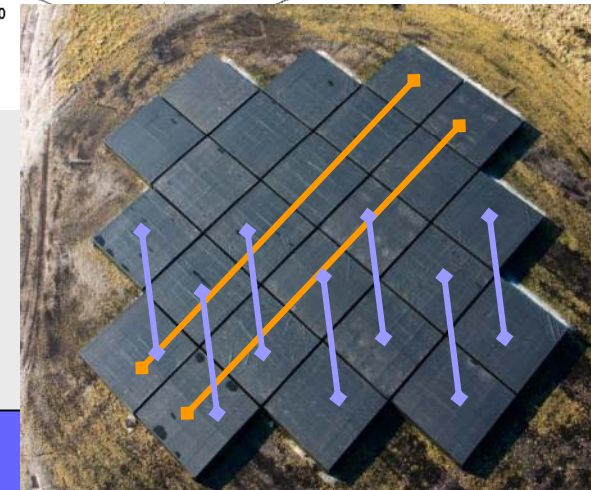
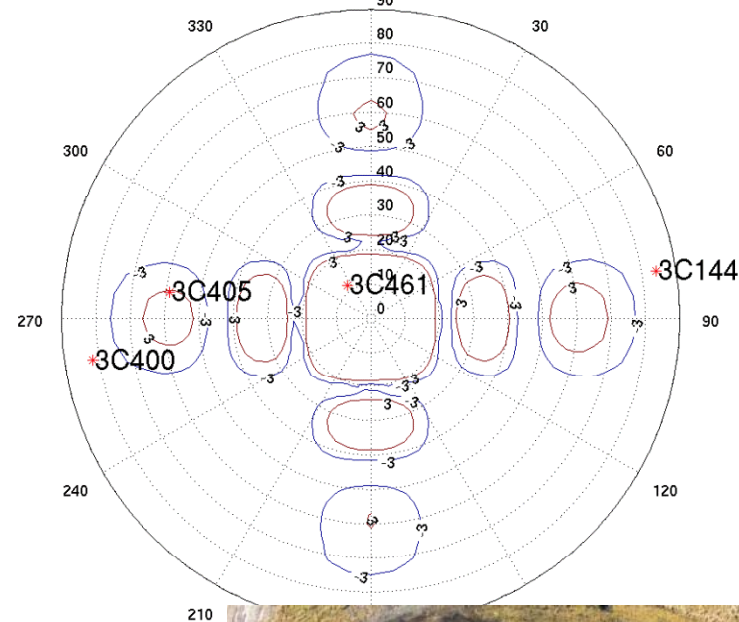
S.J. Wijnholds, Ph.D. thesis, Mar. 2010



Demo redundancy monitoring: 24h observation at CS302, LOFAR HBA tile beam is formed at the zenith



the local sky in view of CS302, at: 05-Sep-2009 00:32:36

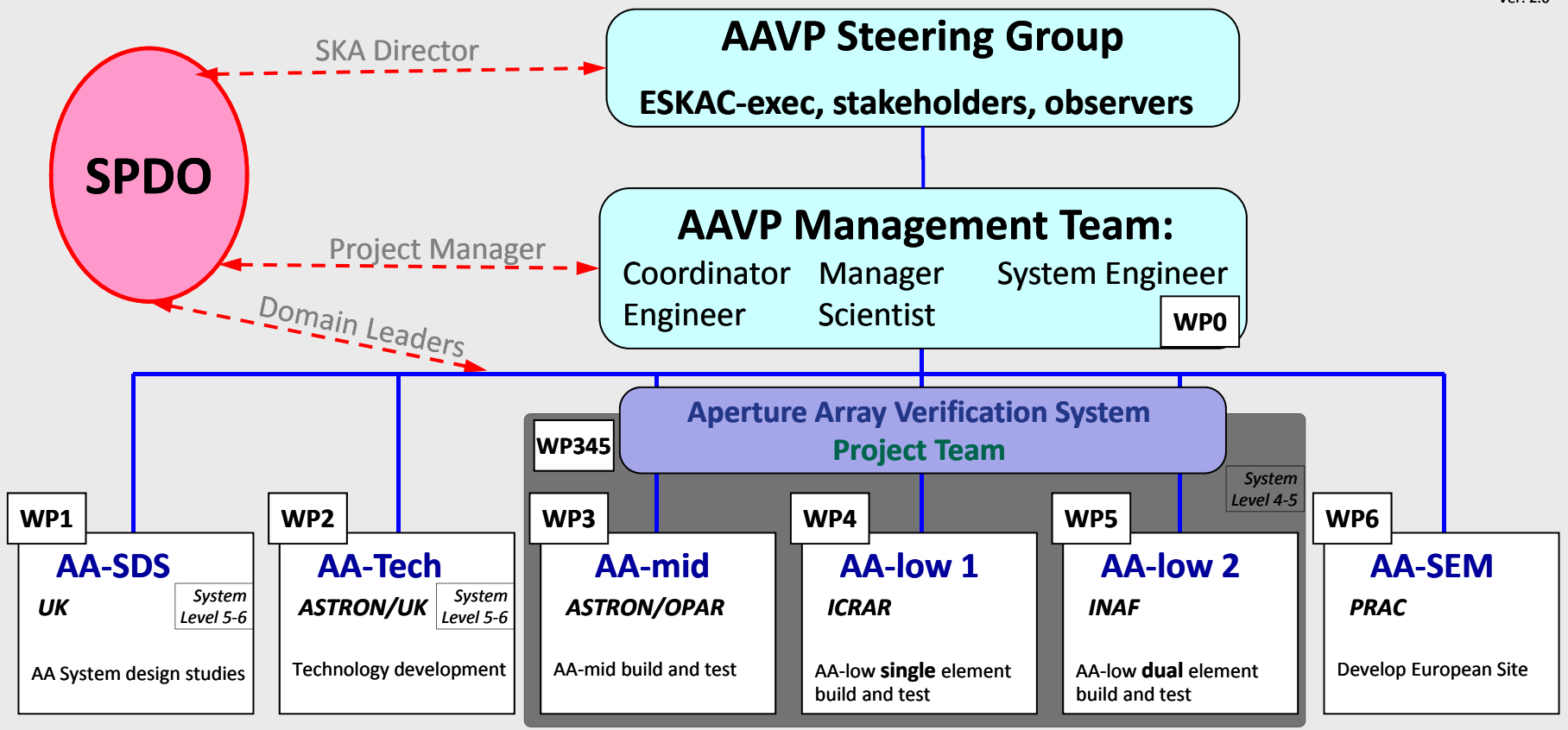


When redundancy assumption holds for HBAs, its calibration algorithm works well

Courtesy: Parisa Noorishad, Stefan Wijnholts

Who are involved?

Consortium/Institute	Country	Main focus
UK-PrepSKA (UCam, UOxford, UMan)	UK	Mgt & Various Techn.
ASTRON/Kapteyn Inst._RUG/Leiden Obs._Univ.Leiden/Radb.Univ./U.A'dam	NI	Mgt & Various Techn.
MPIfR	Germ.	Dig. Sign. Proc.
Obs. De Paris/Obs. Bord./Ud'Orléans	Fr	IC; BF & A/D
INAF_IRA	It	AA-Low
OAN	Sp	Observer
OSO_ChTU	Sw	Calibr.
<i>Torun Centre of Astronomy_UMK</i>	<i>Pol</i>	<i>Tbd</i>
JIVE	Europe	Dig. Sign. Proc.
PRAC	Port	Site dev.
ICRAR	Aus	AA-Low
NRF & ... Observer status	R.S.A. & US	RSA:ESD/EMC?



List of WPx - Leaders and Task Names

AvA 201103v03

AAVP Leading participant	WP Tasks	Leading Participant	WP-Task	WP & WP-T Leaders
WP0 - Coordination and Management	Coordinator	ASTRON	Coordination and Management	Arnold van Ardenne
ASTRON			Program Office (ASTRON)	Truus van den Brink
www.ska-aavp.eu			Program Manager (ASTRON)	Andre van Es
secretary@ska-aavp.eu			Program Scientist (Oxford)	Steve Rawlings
			System Engineer (Cambridge)	Andrew Faulkner
			Program Engineer (ASTRON)	Jan Geralt bij de Vaate
WP1 - AA-System Design Studies	WP Leader	Cambridge	WP1 Manager	Andrew Faulkner
Cambridge	WP1 -T1.6		Taskleaders, SPDO_SE	
WP2 - AA-Technologies	WP Leader	ASTRON	WP2 Manager	Jan Geralt bij de Vaate
ASTRON/UK	WP2 -T1.8	UMAN	Taskleaders, SPDO_PE	Tony Brown
WP3a - AAVS Project Team and System Design	WP3a Leader	Cambridge	Project Manager	Derek McKay-Bukowski
Cambridge/ASTRON	Callm leader	ASTRON	Verification Scientist	Ilse van Bommel
	AA-DS	Cambridge	Calibration & Imaging	Keith Grange
	WP3-5		AA-Domain specialist	Andre Gunst
			WP Project managers & System Designers	
WP3 - AA-Mid	WP Leader	ASTRON	Project Manager	Johan Pragt
ASTRON/Obs.Nancay	WP3-T1.5	Obs.Nancay	System Design	Philippe Picard
		ASTRON	System Design	Dion Kant
			Taskleaders	
WP4 - AA-Low1	WP Leader	ICRAR	Project Manager	Jan Geralt bd Vaate (a.i.)
ICRAR	WP4-T1.4		System Design	tbd
			Taskleaders	
WP5 - AA-Low2	WP Leader	INAF	Project Manager	Jader Monari
INAF	WP5-T1.4		System Design	Federico Perini
			Taskleaders	
WP6 - AA-Site Emulation	WP Leader	IT	WP6 Manager	Domingos Barbosa
PRAC	WP6-T1.4	Logica	Site development	
			Taskleaders	

Industrial involvement

“SKA as green ICT-intensive Research Infrastructure”

- Increased need for Industrial involvement
 - Moving from Technology demo’s toward Industrial pototypes
 - Increase Political support; country dependent
 - Stipulate European role in SKA; also: green design as attractor
- Preparations for ISKAF2011:
 - Stimulate Industry Involvement
- Example: SKA-Eur. & Industry:
 - Neth.: Industry Consortium (Broad)
 - UK: Meetings, Discussions
 - Sweden: Discussions
 - Italy: Large Industries
 - Portugal: Large Industries (Networks, Energy/Solar)
 - Germany: Energy/Solar, other?
 - France: MicroElectr., other?
 - Aus.: Industry Consortium (Broad)

AA-CoDR 19-20 April;

- Docs for your perusal over 21
- Contributions from whole team assembled through AAVP –MT plus AA-domain specialist and SPDO contribution
- Full Charge agreed and drafted
- Panel members:
 - Peter Dewdney, SPDO
 - Raf Roovers NXP Central R&D
 - Richard Williams Cambridge Consultants
 - Colin Lonsdale, MIT/Haystack
- Assessment reporting
 - First impression end meeting (closed session)
 - Written report in a few weeks