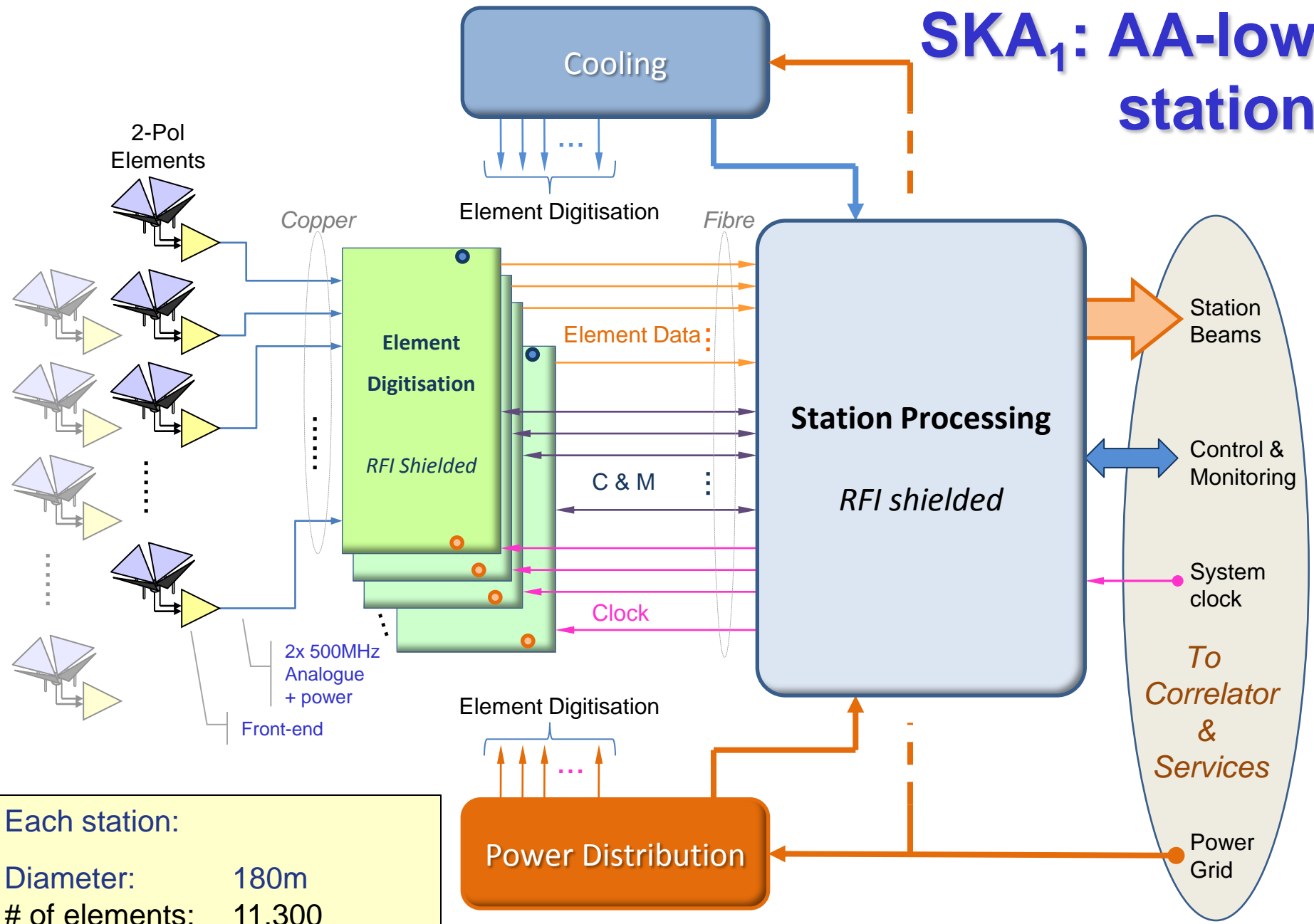


# AA-low to SKA<sub>1</sub>, spec, demonstrators and purpose

Andrew Faulkner / Jan Geralt Bij de Vaate

# SKA<sub>1</sub>: AA-low station



Each station:

Diameter:	180m
# of elements:	11,300
Element links:	Copper

- Limited specification available, memo125:
  - 70 – 450MHz
  - 2000 m<sup>2</sup> / K (z=10)

	AA-mid	AA-low
Low frequency	400 MHz	70 MHz
High Frequency	1400 MHz	450 MHz
AA Type	Dense	Sparse
Element	ORA or Vivaldi	Single (or dual)

Taking the knowledge from  
LOFAR, MWA , Paper etc.

And use latest technology.....

# Proposed specification

- From draft SKA phase 1 system description, June 2010



**Table 4: Representative System Performance Factors for Sparse AAs**

Antenna Efficiency	>90%	Beam former efficiency
Projection loss	0.5	Max at 30 degrees elevation
Minimum Elevation Angle	30 deg	Set by ant. <u>design</u> ; determines max. core FF
$T_{inst}$	150	
$T_{sys}$ at 130MHz	550	400 K average sky noise @ 130MHz
System Equiv. Flux Density (each Sparse AA station)		70MHz 450MHz
System Equiv. Flux Density (250 Sparse AA stations)		70MHz 450MHz
$\Delta S (1\sigma)$ @ 70MHz		Min. detectable cont. flux (Stokes I).
$\Delta S (1\sigma)$ @ 450MHz		Min. detectable cont. flux (Stokes I).
Field-of-View @ 2.3 m wavelength	700 deg	Scales as $\lambda^2$ <u>Beamforming on 5 x 5m tiles</u>
$A_e/T_{sys}$	565 m <sup>2</sup> /K	130MHz, z=10, bore sight, $A_e$ per antenna 1.8m <sup>2</sup>
All Antennas		> 90 MHz & < 420 MHz
<u>SSFoM</u> @ $\lambda = 4.3$ m, $T_{sys} = 2450$ K		Dual <u>polarisation ant</u> grouped into 50 stations.

# Possible AA-low implementation

- From draft SKA phase 1 system description, June 2010

## 7 Sparse Aperture Array

**Table 3: Representative Implementation Based on Sparse Aperture Arrays**

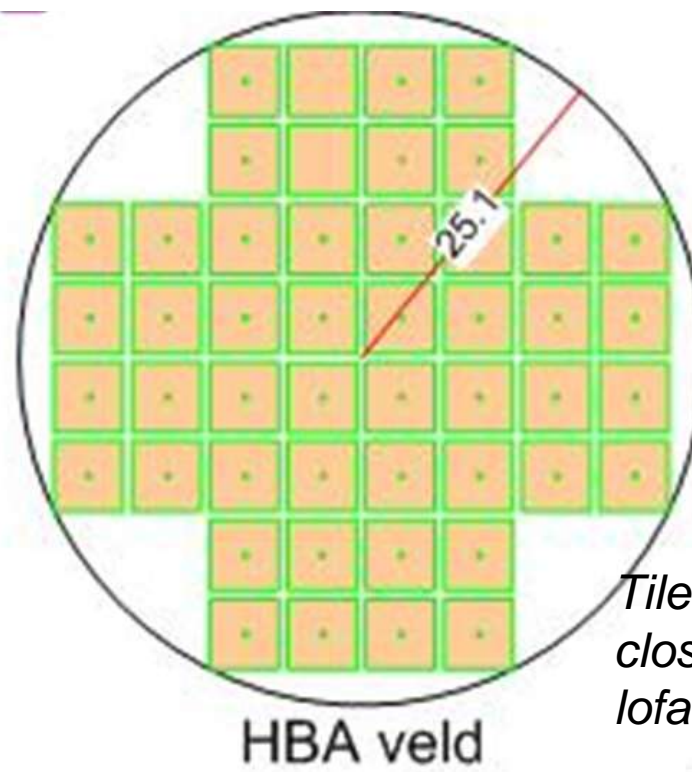
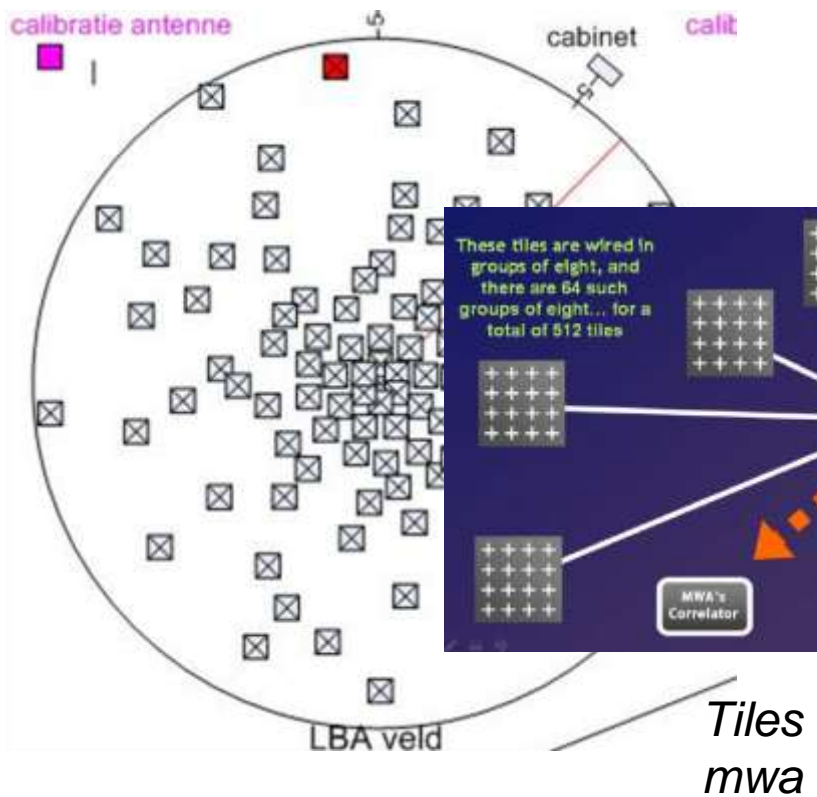
<i>Sparse Aperture Array</i>		50, 180-m diameter stations
Lower Frequency	70 MHz	Dual pol'n (2 orthogonal)
Upper Frequency	450 MHz	
Number of antennas	11.300	Per station
<i>Array Configuration</i>	50 stations	
Core (radius <0.5 km)	~50% (25 stat.)	Fractional total number of AA stations.
Inner (1 < radius < 2.5 km)	~20% (10 stat.)	"
Mid (2.5 < radius < 180 km)	~30% (15 stat.)	In clusters of 5 stations (Total of 15 clusters)
<i>Digital Outputs</i>		
Sample streams	480	Max sub-bands
bits per sample	4	bandwidth dependent
<i>Signal Transport System</i>		Optical fibre to central data processor
Radius < 100 km	4.38Tb/s *	Data rate per station
		Optimized layout, buried fibre, highly multiplexed.
<i>Central Data Processing System</i>		
<u>Correlator</u>		
Max Beams per station per coarse frequency channel		Nyquist sampled on a square grid

# Options: tailoring gain

- Scan range
  - higher gain antennas possible when limited to 45 degrees
  - improves  $A_{\text{eff}} \sim 3\text{dB}$
  - Reduces  $T_{\text{sys}}$
  - 60 degrees scan will be degraded

# Station configuration

- Station dynamic range: configuration



Tiles on regular grid closed packed lofar high band

# Trade offs

	Deployment cost	Freq bandwidth	$A_{\text{eff}}/T_{\text{sys}}$	Dynamic range
Random elements	High	Good	Optimal	Optimal
Random tiles	Moderate	Good	Good	Good
Regular grid tiles	Low	Optimal	Good	(strong gratings)

LOFAR, MWA, AAVP will need to confirm/complete the trade-offs

Table should be expanded: front-end / system study in AAVP



# Key specifications for the AAVS1/2 AA-low designs

- $A_{\text{eff}}/T_{\text{sys}}$  over the 70-450 MHz freq band:  
science priorities required

50MHz??

- FoV requirement:
  - sub station beamforming/data reduction.....
  - Output of a station (180m $\emptyset$ ) will be between 150 and 1500 10Gb/s data channels
  - Uniboard2 type beamforming development

**Table 2. Peak sensitivity at Zenith**

Freq MHz	$T_{\text{sys}}$ K	$A_{e8}/T_s$ $\text{m}^2 \text{K}^{-1}$	$A_{e16}/T_s$ $\text{m}^2 \text{K}^{-1}$	$A_{e32}/T_s$ $\text{m}^2 \text{K}^{-1}$	$A_{e64}/T_s$ $\text{m}^2 \text{K}^{-1}$
69	2650	566	452	302	226
85	1595	940	752	502	376
98	1125	1334	1066	712	534
120	695	1446	1726	1150	864
138	500	1500	2400	1600	1200
170	322	1538	2460	2484	1864
196	238		2520	3360	2520
240	163		1840	3238	3680
276	120			3334	5000
392	75			2666	4000
538	50			2000	3000

- Bregman, June 2010
- Element packing density increases to the right columns: dense for the lower frequencies
- Resulting in
  - large bandwidth
  - Stable impedance for noise match
  - Lower  $A_{\text{eff}}$  at lower freq

Can optimise low freq for EoR or higher freq for Pulsars etc...

Date	Milestone/event	Description
2010Q3	SKA1 definition published	AA-lo highlighted selected for Phase 1
2011Q1	<b>CoDR</b> : Aperture Array	Design review covering AA-low and AA-mid
2011Q1	AAVS test requirements	Agree tests required of the AAVS
2011Q2	AA element review	Review status of the alternative AA elements
2011Q4	AA-mid front-end design	Incl. element, LNA and analogue beamforming
2011Q4	AA-lo element performance review	Select elements that will be tested on AAVS <sub>1</sub>
2011Q4	AAVS <sub>1</sub> digitisation and beamformer design complete	Digital tile beamformer for AA-low and AA-mid
2011Q4	<b>CDR</b> : AAVS <sub>1</sub>	Approve AAVS <sub>1</sub>
2012Q2	<b>PR</b> : AAVS <sub>1</sub>	Approval for AAVS <sub>1</sub> production
2012Q3	AAVS <sub>1</sub> sub-system production	Industry builds AAVS <sub>1</sub> single array subsystems.
2012Q3	Construct AAVS <sub>1</sub>	Build AAVS <sub>1</sub> central processing,
		Install AA-low antennas
		Install AA-mid tiles
2012Q4	Commission AAVS <sub>1</sub>	AAVS <sub>1</sub> proving and calibration
2013Q2	<b>PDR</b> : AAVS <sub>2</sub>	Approve design of AAVS <sub>2</sub> with knowledge from AAVS <sub>1</sub>
2013Q2	AAVS <sub>1</sub> reports (+all verifications...)	Evaluation of AAVS <sub>1</sub> –lo and –mid
2013Q3	<b>CDR</b> : AAVS <sub>2</sub>	Design completed for AAVS <sub>2</sub> -lo and –mid
2013Q4	<b>PR</b> : AAVS <sub>2</sub>	Approval for AAVS <sub>2</sub> production
2014Q1	Construct AAVS <sub>2</sub>	
2014Q3	Commission AAVS <sub>2</sub>	<i>Need to consider other site prep going on</i>
2014Q4	AAVS <sub>2</sub> –low	Evaluation of early AAVS <sub>2</sub> –low using tooling
2015Q1	SKA1 pre-production systems	Complete the manufacture of SKA <sub>1</sub> systems using production tooling (AA-low & Dishes)
2015Q2	<b>PDRs</b> : AA-mid; SP-WBF, PAF	PDR for advanced technology selection, to inform SKA <sub>2</sub> decision
2015Q4	<b>PR</b> : SKA1	Final production reviews of all aspects SKA1
2016Q1	Start SKA1 production	

# AAVS1: AA-low

- Demonstrate electromagnetic and front end performance
  - Built on low RFI site, ideally chosen SKA site
  - Sufficient collecting area, possibly two versions
  - Focus on **broadband** antenna and front-end design
  - Evaluation using existing early processing - UNIBOARD
  - Establish antenna, tile and station configuration
    - Extensive station simulations required and checked
  - Critical Design Review      31-12-2011
  - Commissioning                31-12-2012
  - Reporting                        30-6-2013

- Pre production phase:
  - Preparing for mass-production
- AAVS2 prototype
  - Test the tooled, production components
  - SKA<sub>1</sub> performance compliance
  - Bandwidth, T<sub>sys</sub>, stability, signal processing, calibratability etc..
  - Critical Design Review      30-9-2013
  - Construction                      2014
  - Commissioning                    2015
  - Reporting for SKA1 PR      31-12-2015

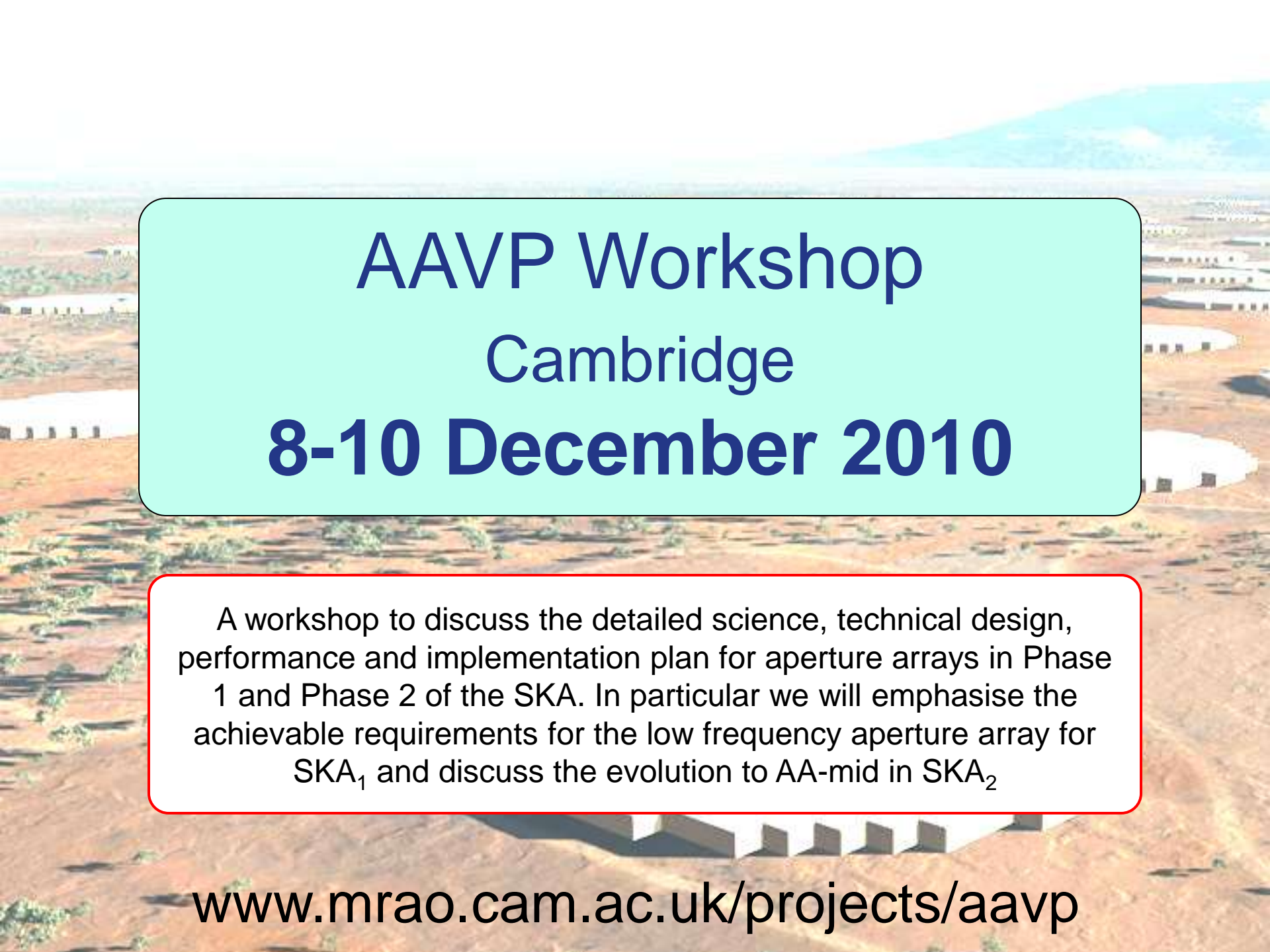
# AAVS2 requirements

- In order to ramp-up production in 2016, AAVS2 should be as large as possible: e.g. 1-2% SKA<sub>1</sub>
- Built with SKA<sub>1</sub> tooling
- Approaching 180 meter diameter station including digital station beamforming
- On the SKA site
- Acceptance tests of this station should be considered in absence of correlator

Schedule is tight: custom MMICs, photonic links, signal processing boards

## Message is:

- This is a *very* tight timeline
- Must have specification stability ASAP
- The window for changes is passing....



# AAVP Workshop

## Cambridge

### 8-10 December 2010

A workshop to discuss the detailed science, technical design, performance and implementation plan for aperture arrays in Phase 1 and Phase 2 of the SKA. In particular we will emphasise the achievable requirements for the low frequency aperture array for SKA<sub>1</sub> and discuss the evolution to AA-mid in SKA<sub>2</sub>

[www.mrao.cam.ac.uk/projects/aavp](http://www.mrao.cam.ac.uk/projects/aavp)