

LO and Timing Requirements

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- Timing Requirements (pulsars)
- Clocks
- Coherence/Stability requirements
- Commercial freq standards
- Calibration issues
- General implementation issues
- Phase transfer

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- Time & Frequency references are fundamental requirements for timing observations & interferometry
 - Connected arrays already achieve phase transfer over 10-200km sufficient to maintain coherence for 20 GHz and above
 - VLBI arrays use independent H-masers for observations >20 GHz; phase stable on timescales of minutes to hours at ~5 GHz
 - Issues?
 - check requirements
 - review frequency standards, distribution techniques
 - phase calibration
 - manageable & cost-effective implementations for >1000 elements

Timing Requirements

Pulsar Timing

few x 10ns now for best MSPs

goal ~ 10ns SKA₁,
~few ns SKA₂ ?

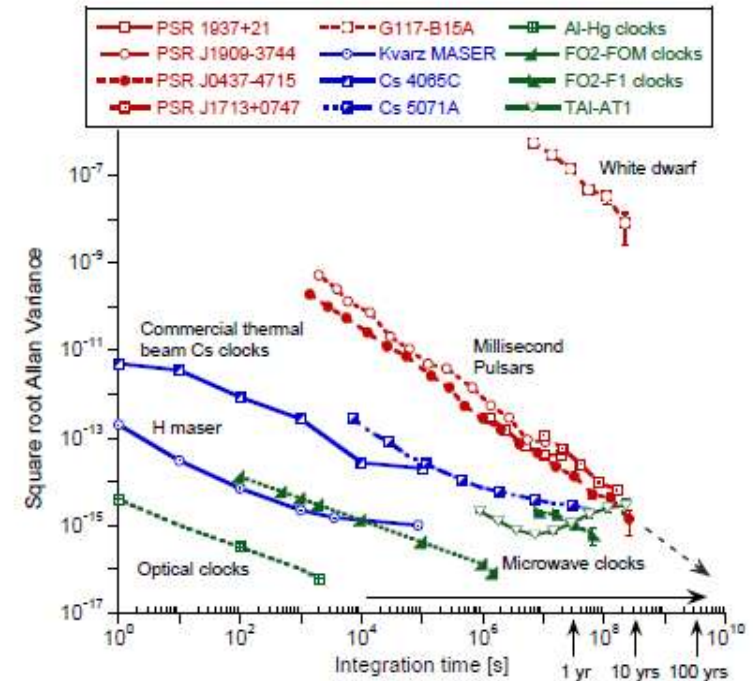
Time standard derived from local
H-maser

referenced through GPS
ultimately to TT(BIPM)

long term stability

Pulsar timescale?

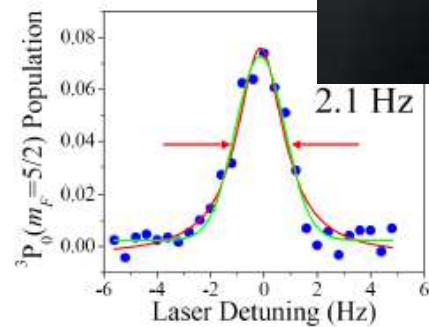
Requirements for timing fast
transient events?



Hartnett & Luiten 2010

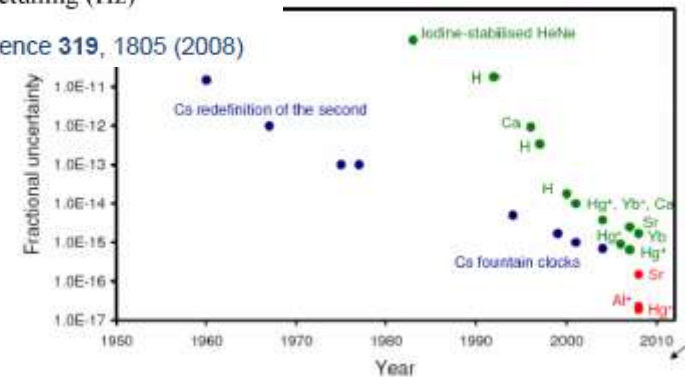
Optical clocks

- Trapped ion clocks now at 10^{-17}
- ESA have considered replacing H-masers with optical clocks at major ground stations (ESA/NPL report)
- ~ €1M cost + development
- Significant environmental issues (B-field, acceleration, acoustics)
- Should radio astronomers have similar ambition
- SKA *could* be location for significant frequency/time standard
 - links to other time standard labs
 - links to astro timescales
 - links to s/c navigation, gravity experiments



2S_1

A. Ludlow et al. Science **319**, 1805 (2008)



Coherence (Stability) Requirements

- Coherence function $C(T)$

$$C(T) = \frac{1}{T} \int_0^T e^{j\phi(t)} dt$$

Loss $\langle C^2(T) \rangle$

For Gaussian ϕ

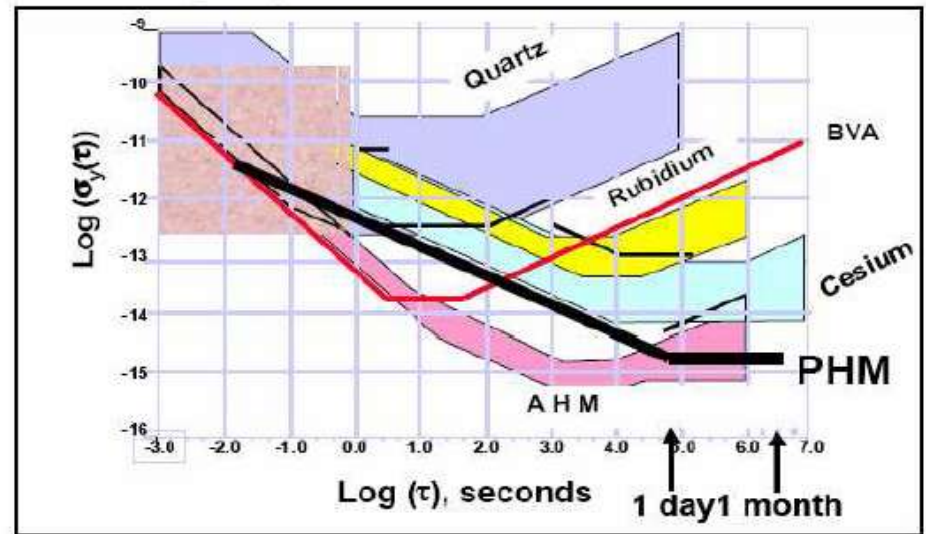
8 dg rms \rightarrow 1% loss

2 ps at 10 GHz

- Clearly must achieve this at integration (~ 1 s) level
- Longer timescale requirements depend on calibration approach

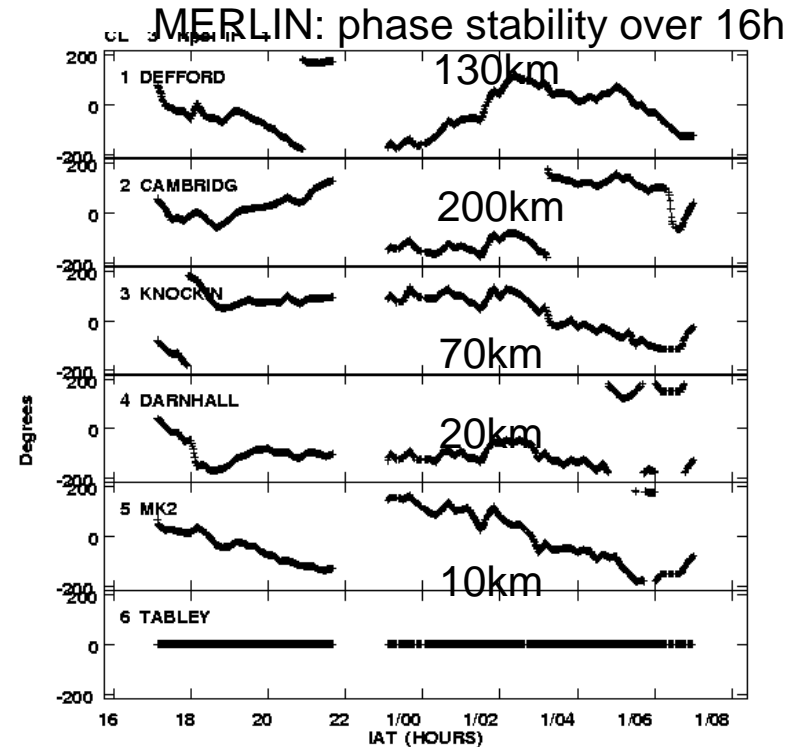
(Commercial) Oscillator performance

- At 1 sec best quartz oscillators (BVA) are very good!
- Active H-maser better > 10s
- At 1000s:
Passive H-maser : 10^{-14}
Active H-maser 10^{-15}



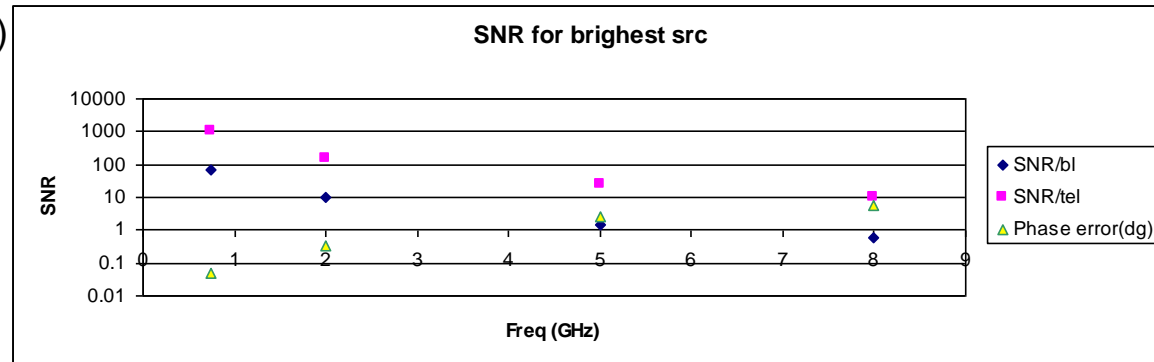
Stability requirements and calibration techniques

- Current cm-wavelength interferometers use phase referencing:
 - switch to (compact, bright) ref source
 - minimize angle (atmosphere, geometry); <5dg
 - minimize time (atmosphere); < 10 min
 - require linear phase drift over switching cycle
 - accustomed to stability
- Dynamic range $\sim n_a n_t^{1/2} / (\Delta\phi)$
 - VLA 1hr; MERLIN 12hr
 - $\Delta\phi$ 10dg \rightarrow DR \sim 400:1
 - SKA₁ 1500:1



SKA₁ Calibration

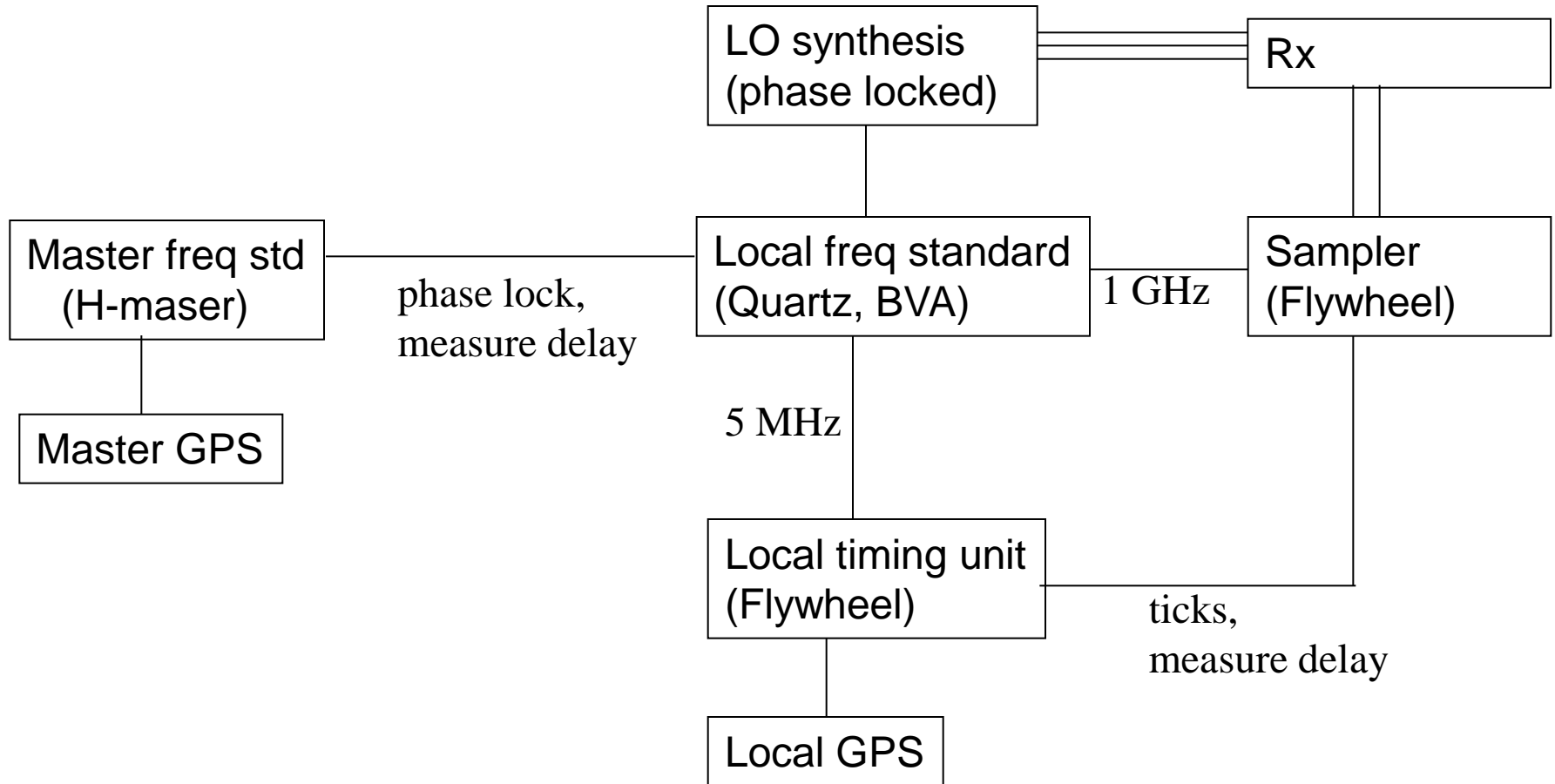
- 0.5 - 10 GHz (b/w 0.5..1..2...4 GHz)
- $T_{\text{sys}} = 30$;
- 250 x 15m dishes
- $t_{\text{int}}=100\text{s}$
- SNR/bl on expected brightest source in FOV
- Per tel Phase error
~8 dg at 8 GHz
- Based on 1.4 GHz source counts
- For 200km baselines; ~30% sources unresolved (Garrington & Garrett, Gurvits et al); stations helps
- But $t_{\text{int}}=10\text{s}$ difficult
- Require 100-1000s stability for >5 GHz
(Could manage with 10-100s < 5 GHz)



Requirements and general implementation issues

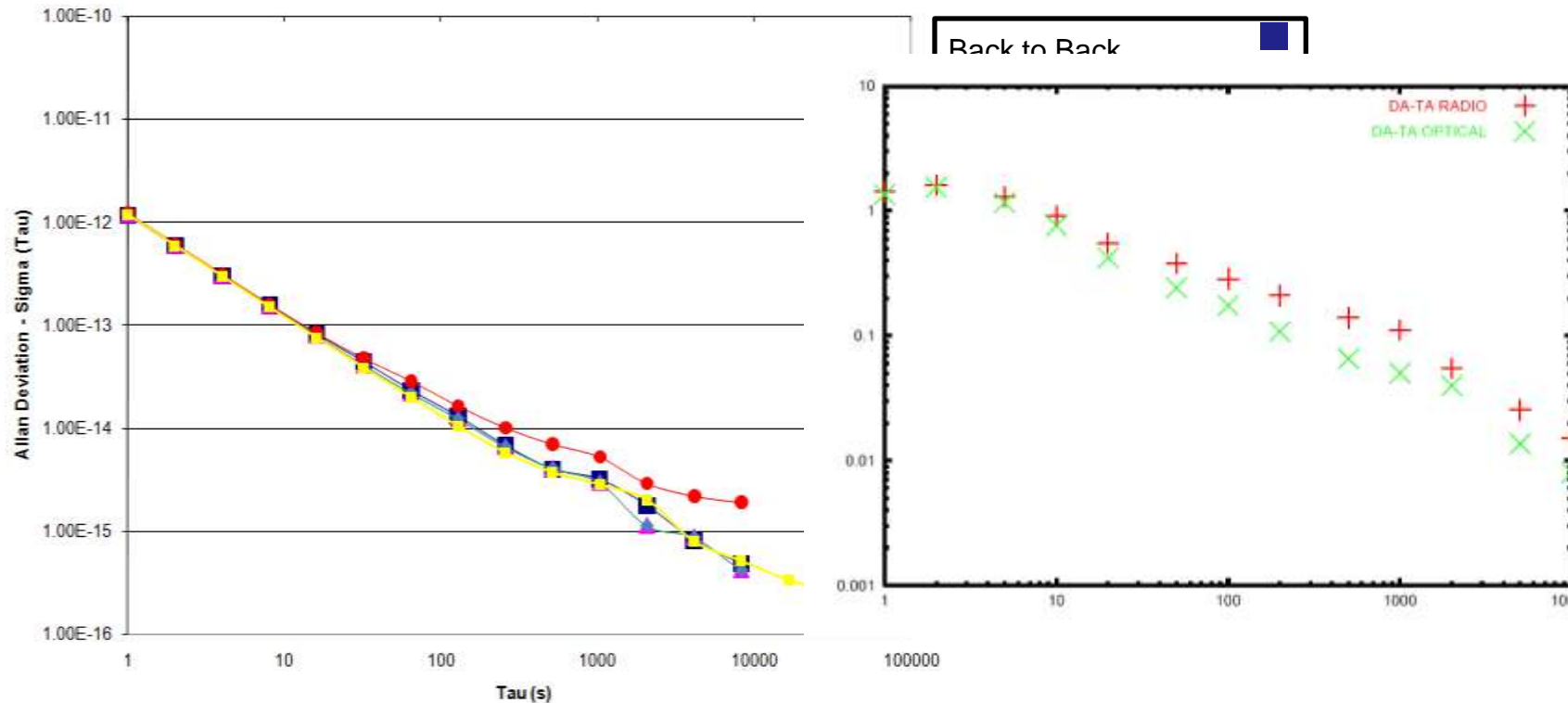
- Rb, Quartz (BVA), PHM probably not good enough
 - insufficient stability on 100-1000s
- Need H-maser(s) + distribution system
 - take advantage of short term stability of BVA at remote site
- Dishes/stations need to derive LOs, sampler clocks, 1pps from local station oscillator (BVA)
 - trace 1pps to UTC via GPS
 - phase-locked multipliers for LOs

Time & Freq system



Phase transfer systems

- $\sim 1\text{-}2\text{ps}$ over 1s-10min demonstrated in SKADS

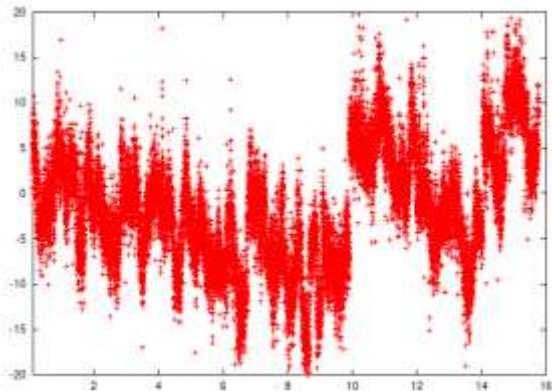


- Most distant stations could use independent clocks

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- Requirements
 - reference 1 pps ticks coherent (jitter < 0.1 ns) with central clock LO maser.
 - measure 1pps against local GPS ticks ~10ps
 - provide optical timecode as required (cf WIDAR correlator TC)
 - measure the round trip delay to sampler
 - provide tick-timing offsets to the correlator/data acquisition system for corrections in software

Time & Freq Processing at dish/station

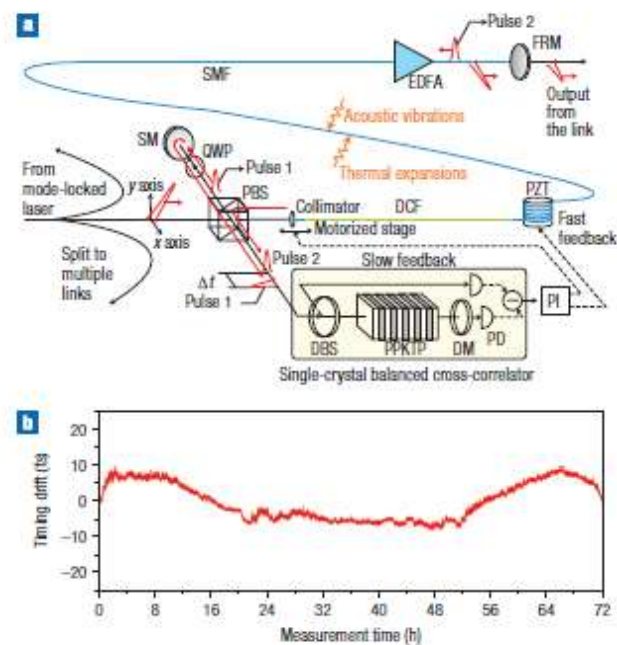
- e-MERLIN Implementation (B Anderson et al)
 - low cost iLotus (ex Motorola) GPS: 2-3 ns stability
 - 5/10 MHz input (maser/crystal)
 - Clock synthesis chip: 128 & 512 MHz
 - FPGA: Flywheel divider chain by 2048, 625, 5 and 20 generates the 1 second ticks from the 128 MHz
 - Divide by 625, 5 and 20 are cleared by the 1 second ticks from the GPS receivers
 - The GPS ticks and the incoming TCs sample the divide by 2048 to measure coarse delay (8ns)
 - Fine delays (0.5 ns) are obtained by processing multiple copies of offset ticks.



s/w correction can improve to <5ns

Femtosecond time distribution

- X-ray Free Electron lasers use fs pulses; require fs synchronisation of major, distributed subsystems
- Use mode-locked lasers & active compensation on fibre links; 6 fs rms over 72 hrs/300m
- Techniques for locking 10 GHz reference with stability 2×10^{-19}
- Will be developed for facility use, across range of applications



Kim et al 2008

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- Requirements
 - improved calibration analysis using sky and system simulation
 - other timing requirements (transients)
 - low frequency requirements (LOFAR)
 - formalise for system design
 - Review
 - facility-class optical clocks; time transfer techniques
 - Inputs from Pathfinders/Precursors
 - e-MERLIN, EVLA: Phase transfer on fibre
 - e-MERLIN: local GPS clocks
 - LOFAR: Rb
 - MeerKAT: fibre phase studies
 - Phase transfer
 - design for low-cost, scalable, reliable pulse distribution system
 - performance evaluation

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- Timing requirements for pulsars: 10ns
 - Case for next-generation (optical) clock
 - Interferometry requirements
 - 8dg @ 10 GHz (2 ps)
 - stability over 100-1000s
 - used to stability over many hours
 - active H-maser + distribution system
 - Architecture
 - central freq std & clock → phase transfer → local standard referred to local GPS
 - Phase transfer
 - ps stability over 1-1000s demonstrated using pulse transmission on fibre over 100km, multiple hops; working on scalability





