

What have we learnt from SCORPIO?

State-of-the-art and current issues

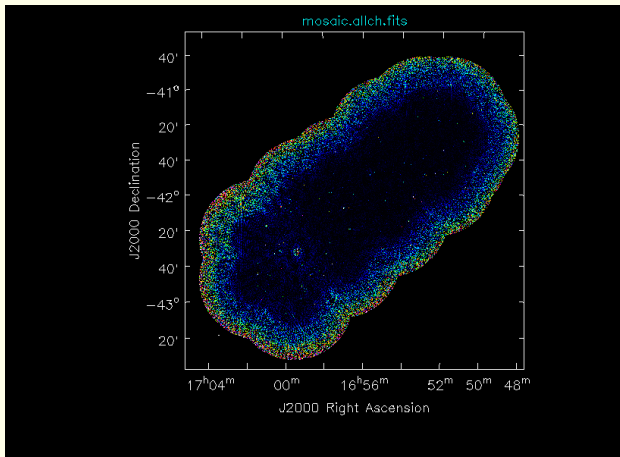
Adriano Ingallinera

INAF - Osservatorio Astrofisico di Catania

What have we done so far?

Observations

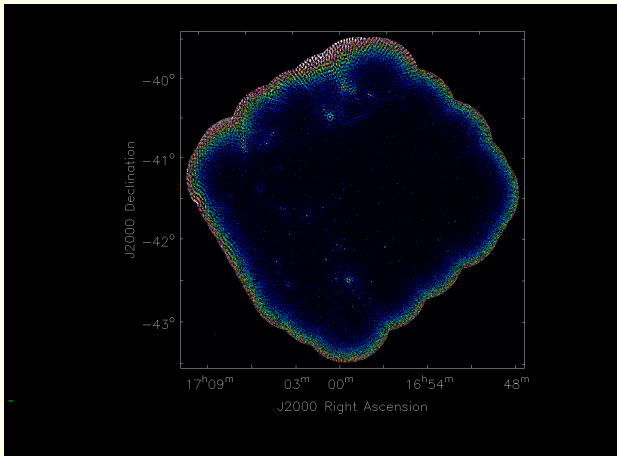
Pilot region with ATCA extended configurations (6x)



What have we done so far?

Observations

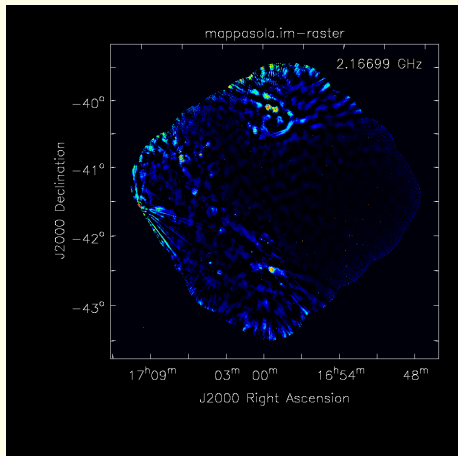
Entire region with ATCA extended configurations (6x)



What have we done so far?

Observations

Entire region with ATCA compact configurations (EW3xx)



What have we done so far?

Observations

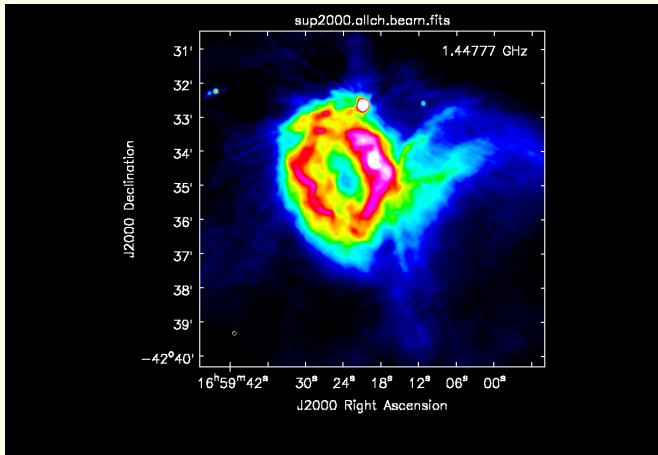
Entire region with Parkes (Apr 2016, 1.2 - 1.8 GHz)



What have we done so far?

Data reduction

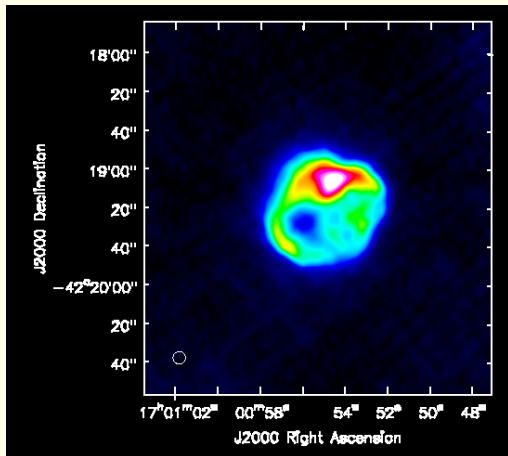
Stokes-*IQUV* maps divided in 7 sub-bands from ATCA data



What have we done so far?

Data reduction

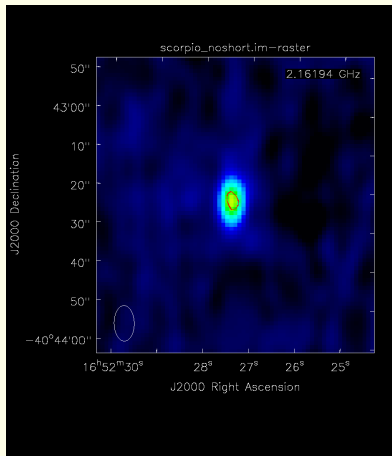
Stokes-*IQUV* maps divided in 7 sub-bands from ATCA data



What have we done so far?

Data reduction

Stokes-*IQUV* maps divided in 7 sub-bands from ATCA data



What have we done so far?

Data analysis

Point source extraction and catalogue (Umana et al. 2015)

ID	l (deg)	b (deg)	RA (J2000)	Dec (J2000)	S (mJy)	ΔS (mJy)	Area (beam)	Matching catalogues
SCORPIO1_001	343.0025	1.7604	16:50:12.21	-41:48:56.2	33.33	1.03	1.0	-
SCORPIO1_002	343.0051	0.2234	16:56:39.13	-42:47:08.2	0.49	0.09	1.0	-
SCORPIO1_003	343.0134	0.6086	16:55:02.85	-42:32:15.5	7.74	0.29	1.4	NMGm
SCORPIO1_004	343.0138	1.7208	16:50:24.33	-41:49:56.4	1.60	0.17	1.0	N
SCORPIO1_005a	343.0139	1.1508	16:52:46.37	-42:11:42.1	12.80	0.50	2.0	N
SCORPIO1_006	343.0152	0.1166	16:57:08.49	-42:50:39.6	1.27	0.16	1.7	NMGW
SCORPIO1_007	343.0157	-0.1830	16:58:25.48	-43:01:49.4	1.51	0.12	1.2	NMG
SCORPIO1_005b	343.0186	1.1577	16:52:45.61	-42:11:13.0	106.56	3.20	1.4	-
SCORPIO1_009	343.0201	1.5248	16:51:14.28	-41:57:09.1	1.34	0.11	1.0	-
SCORPIO1_010	343.0203	0.3109	16:56:19.93	-42:43:08.3	20.97	0.64	1.2	Gmw
SCORPIO1_011	343.0216	0.7407	16:54:31.13	-42:26:53.1	1.12	0.10	1.3	G
SCORPIO1_012	343.0277	0.8702	16:53:59.71	-42:21:42.3	1.17	0.09	1.2	NMXAH
SCORPIO1_013	343.0317	0.8579	16:54:03.63	-42:21:59.3	0.44	0.08	1.0	NMH
SCORPIO1_014	343.0322	0.9655	16:53:36.65	-42:17:53.0	1.21	0.12	1.3	NH
SCORPIO1_015	343.0356	0.2182	16:56:46.68	-42:45:53.7	8.06	0.25	1.0	N
SCORPIO1_016	343.0394	1.5983	16:50:59.99	-41:53:26.8	2.49	0.13	1.0	I
SCORPIO1_017	343.0415	1.2316	16:52:31.80	-42:07:20.7	1.46	0.13	1.0	NM
SCORPIO1_018	343.0428	0.0266	16:57:37.15	-42:52:43.9	5.29	0.22	1.0	-
SCORPIO1_019	343.0457	0.8848	16:53:59.73	-42:20:18.9	71.55	2.15	1.0	Hm
SCORPIO1_020	343.0461	1.2752	16:52:21.85	-42:05:28.5	4.71	0.25	2.2	I
SCORPIO1_021	343.0468	0.3010	16:56:27.87	-42:42:16.1	0.47	0.06	1.0	NMGW
SCORPIO1_022a	343.0489	0.9666	16:53:39.77	-42:17:04.1	109.20	3.28	1.6	Gm
SCORPIO1_023	343.0501	1.6133	16:50:58.45	-41:52:22.7	1.75	0.17	2.1	NMWAI
SCORPIO1_024	343.0502	0.4249	16:55:57.01	-42:37:27.4	1.04	0.09	1.0	NMGH
SCORPIO1_022b	343.0514	0.9634	16:53:41.09	-42:17:04.3	23.35	0.72	1.6	m

Problems in Stokes-*I* maps

Imaging artefacts caused by:

- a** wide band ($\Delta\nu/\nu \sim 1$)
- b** diffuse Galactic emission
- c** extended Galactic sources
- d** bright barely resolved sources (mainly extragalactic)

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Solutions:

- > **a** mitigated by Taylor expansion and sub-band division
- > **b** and **c** mitigated by short-baselines
- > **d** unresolved, other CLEAN algorithm?

Problems with Galactic extended sources

Poorly imaged even with short baselines (LAS $\sim 4'$ at 3 GHz):

- > no accurate flux density measurement
- > no spectral index maps
- > morphology misinterpretation

Problems with Galactic extended sources

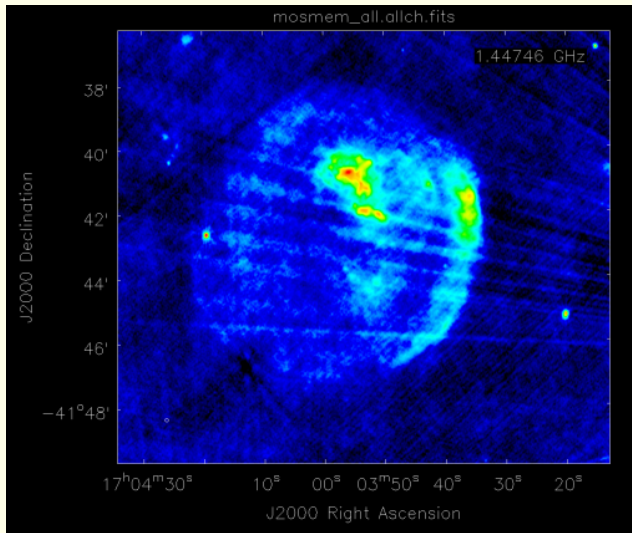
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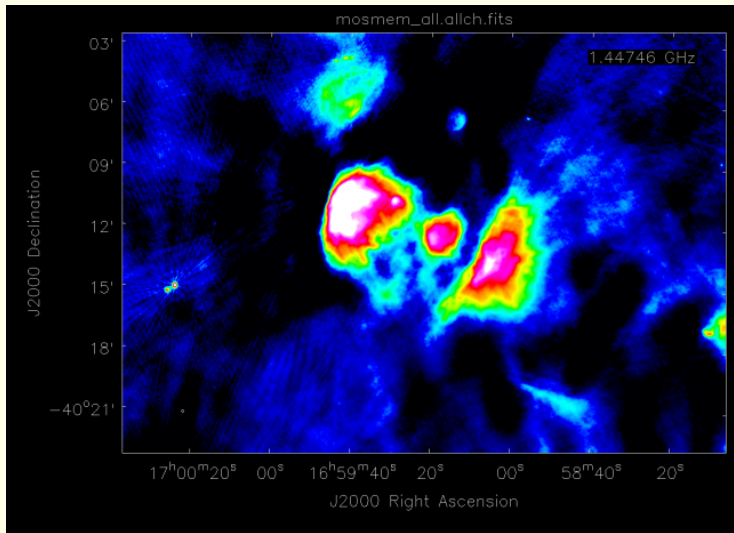
Solutions:

- > single-dish data from Parkes (under reduction)

Examples of poorly imaged sources



Examples of poorly imaged sources



Problems in Stokes- Q and - U maps

Artefacts similar to Stokes- I maps:

- > polarized bright resolved extragalactic sources?
- > polarized diffuse emission?

Problems in Stokes- Q and - U maps

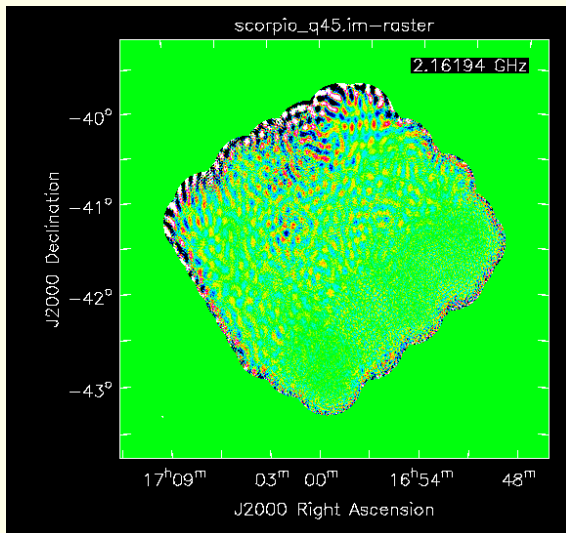
Artefacts similar to Stokes- I maps:

- > polarized bright resolved extragalactic sources?
- > polarized diffuse emission?

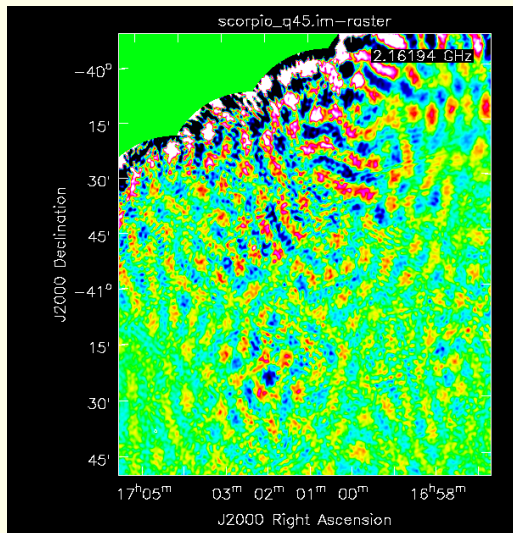
Solutions:

- > none, please help!

Problems in Stokes- Q and - U maps



Problems in Stokes-Q and -U maps



Despite difficulties, three papers already produced:

- > Umana et al. 2015 (pilot region)
- > Riggi et al. 2016 (extended source extraction)
- > Cavallaro et al. *submitted* (spectral indices)