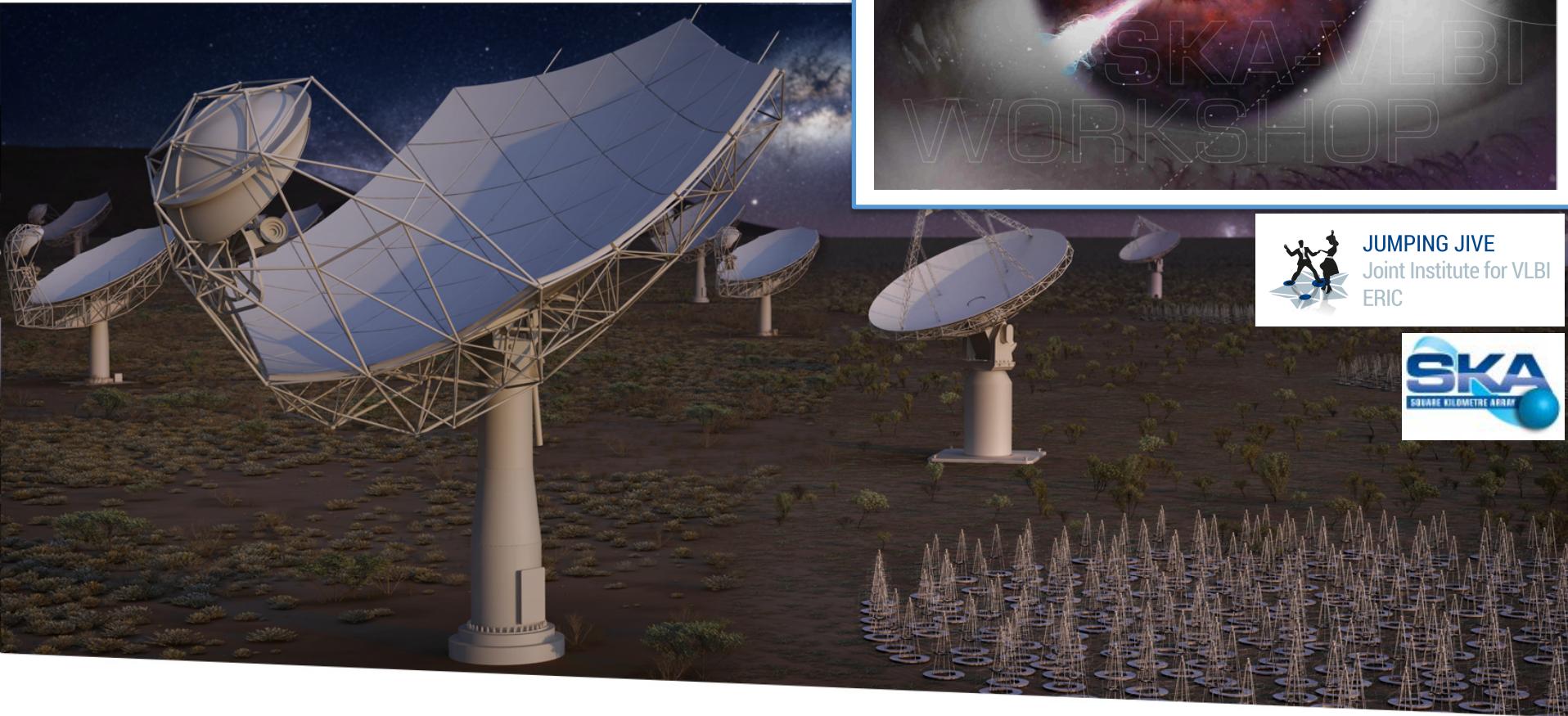


SKA-VLBI capacity and technique



SQUARE KILOMETRE ARRAY

Exploring the Universe with the world's largest radio telescope

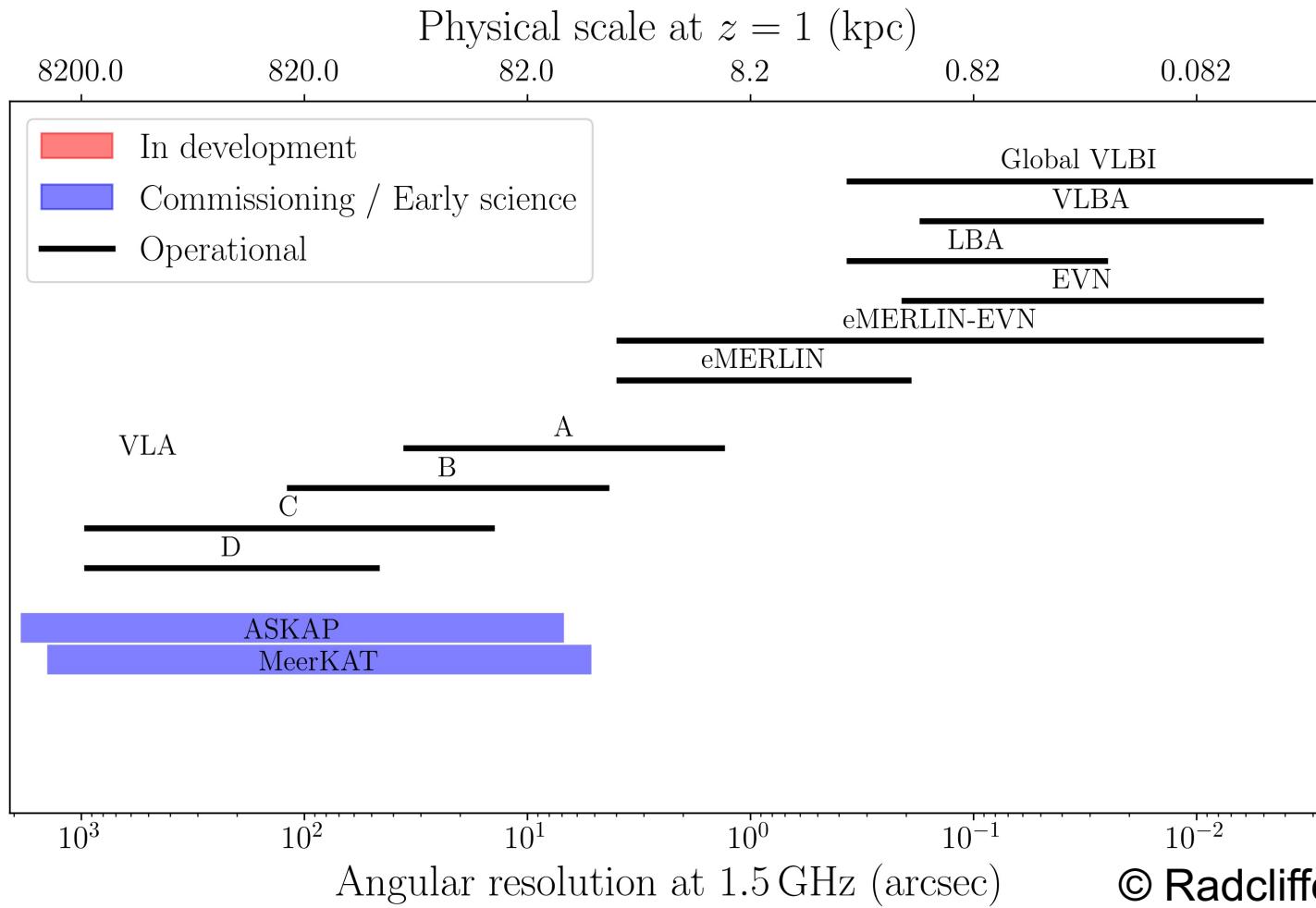


Cristina Garcia Miro / Tao An

JUMPING JIVE Project / SKA VLBI SW
miro@jive.eu JIVE / SHAO

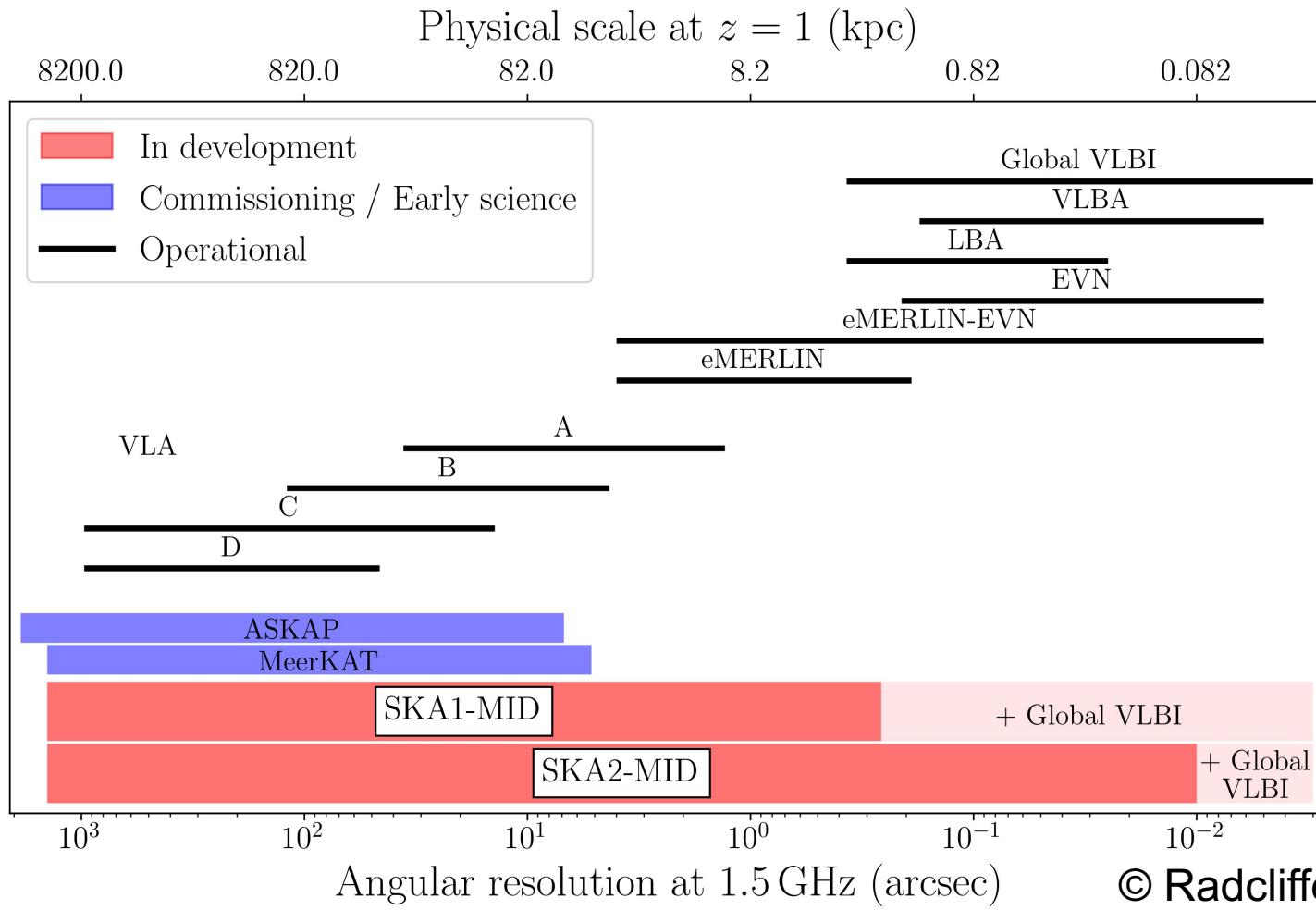


SKA-VLBI: broadest angular resolution



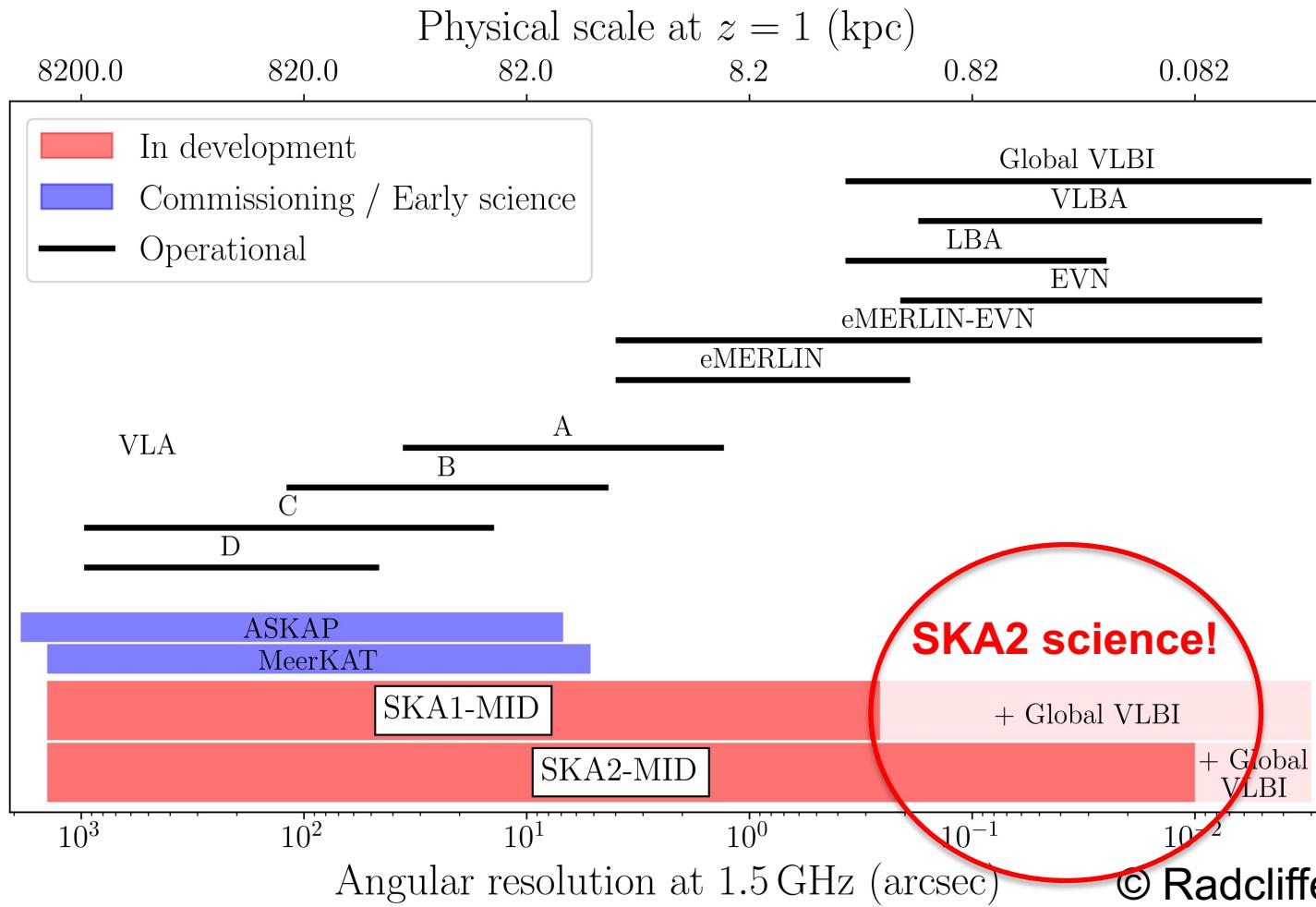


SKA-VLBI: broadest angular resolution





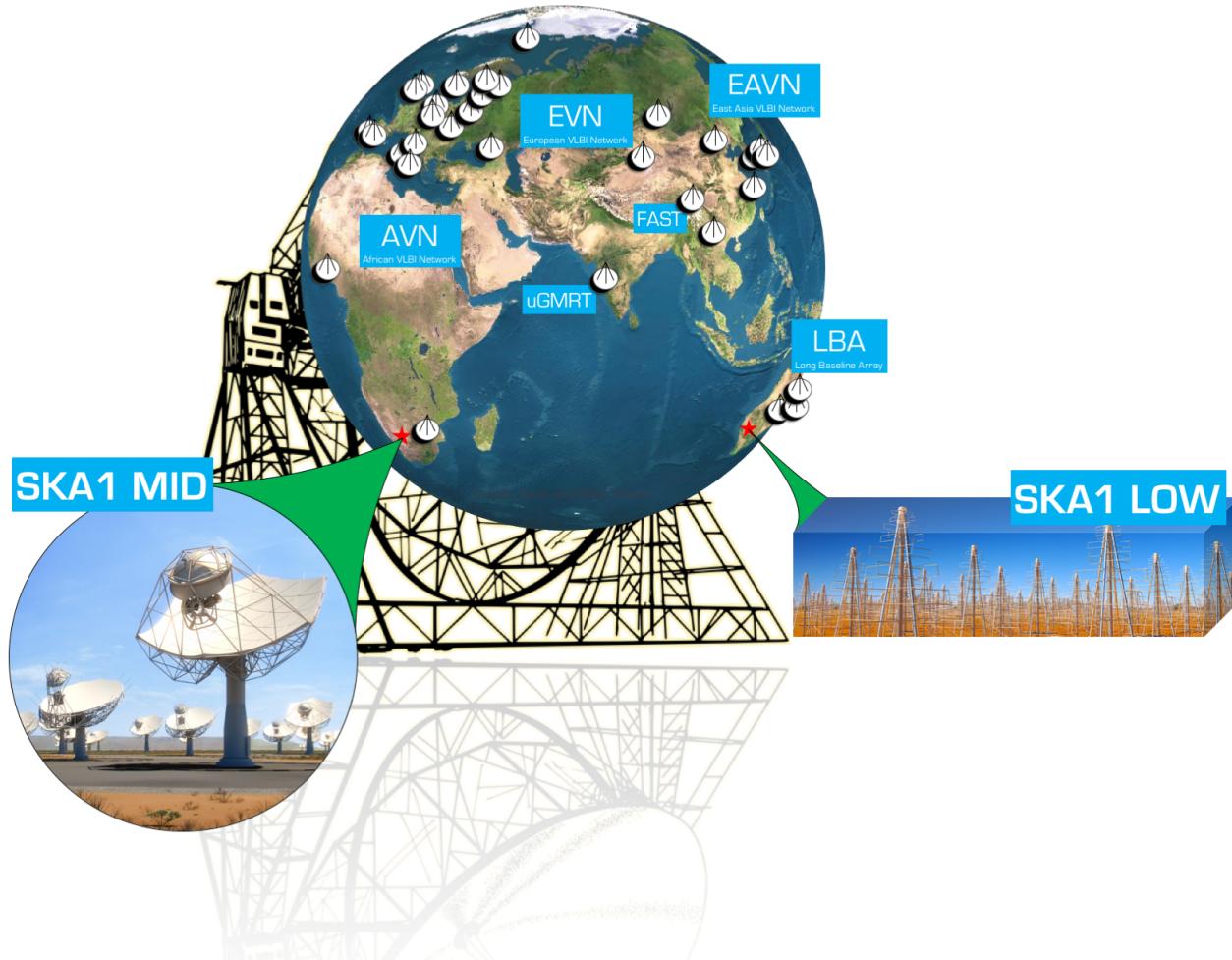
SKA-VLBI: broadest angular resolution



SKA-VLBI

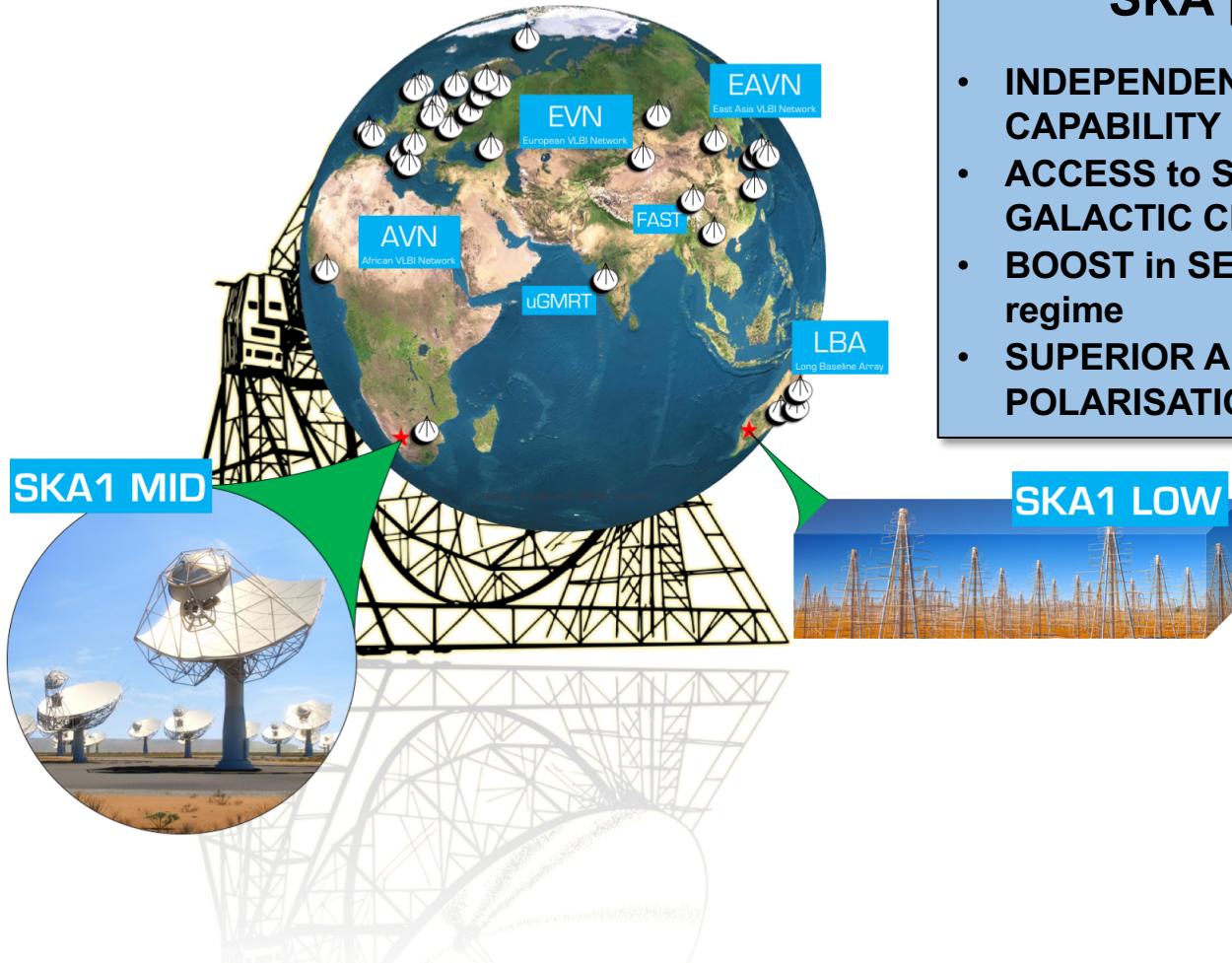


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SKA-VLBI

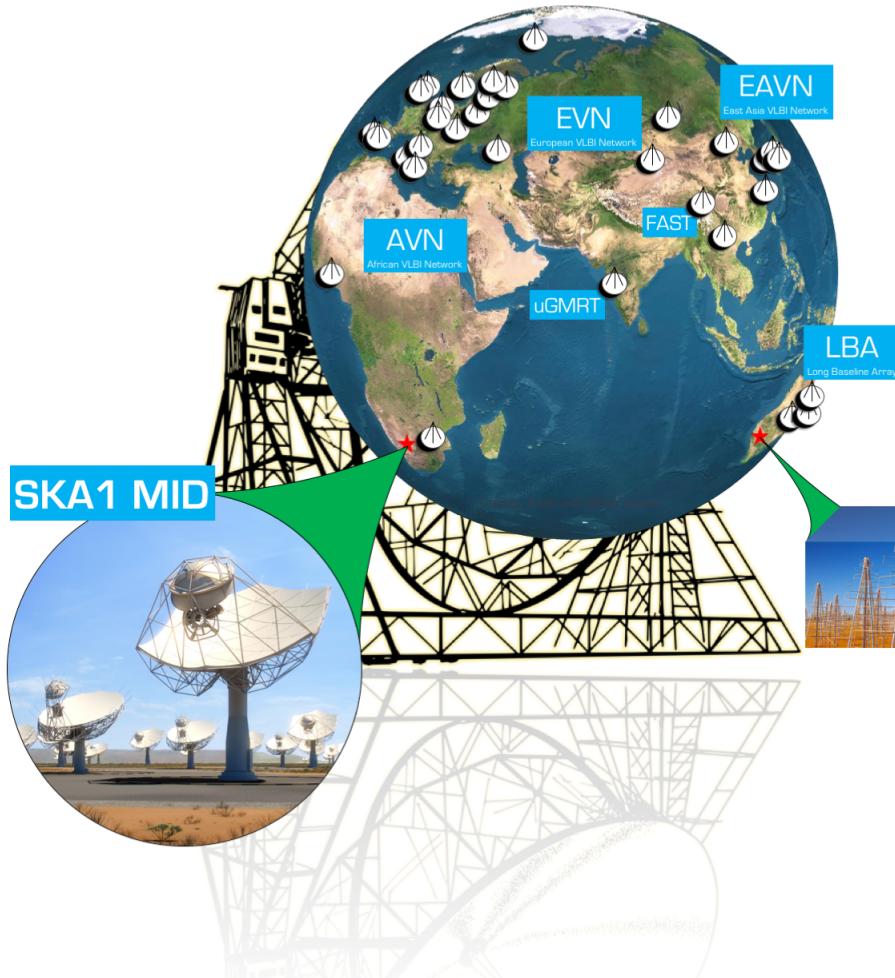


SKA provides ...

- INDEPENDENT MULTI-BEAM CAPABILITY
- ACCESS to SOUTHERN SKIES and GALACTIC CENTER
- BOOST in SENSITIVITY to μJy regime
- SUPERIOR AMPLITUDE and POLARISATION CALIBRATION



SKA-VLBI



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- INDEPENDENT MULTI-BEAM CAPABILITY
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- BOOST in SENSITIVITY to μJy regime
- SUPERIOR AMPLITUDE and POLARISATION CALIBRATION

SKA1 LOW

VLBI provides ...

- IMAGES of SKY at BROAD RANGE of ANGULAR SCALES RESOLUTIONS
- HIGH RESOLUTION for SKA HPSOs
- INDEPENDENT COMMISSIONING TOOL and EARLY PUBLIC RELATIONS OPPORTUNITIES
- AN ENTHUSIASTIC USER COMMUNITY



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SKA-VLBI Science



SKA-VLBI Science

- ✓ JUMPING JIVE 2nd deliverable: Portfolio of Science Cases
- ✓ SKA VLBI Science Working group support
- ✓ Outcomes inform the SKA-VLBI operational model and KSPs

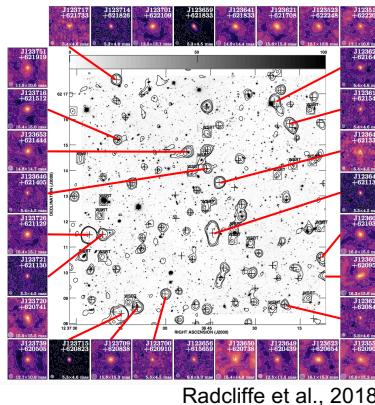


VLBI with SKA1-MID:

6 science cases updated (from SKA1 scientific use cases doc.)

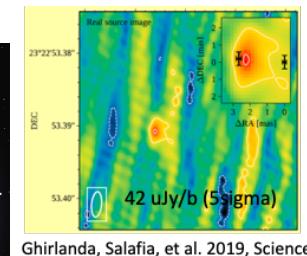
SKA continuum surveys

Adding high angular resolution to SKA surveys:
Giroletti et al.



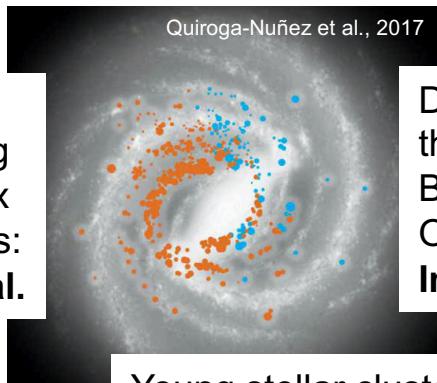
Transients

Resolving (ultra)-relativistic outflows with SKA-VLBI:
Paragi et al.



Our Galaxy, Astrometry and Col

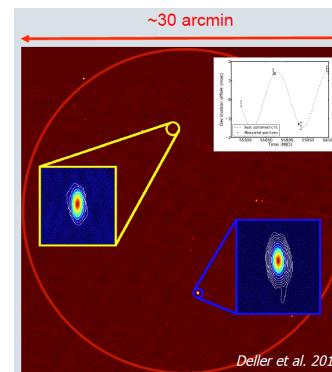
Galactic structure using maser parallax measurements:
Ellingsen et al.



Dynamics of the Galactic Bulge using OH masers:
Imai et al.

Young stellar cluster deep field: **Hoare et al.**

Pulsars, Astrometry



Parallax measurements of SH pulsars:
Deller et al.

VLBI with SKA1-MID:

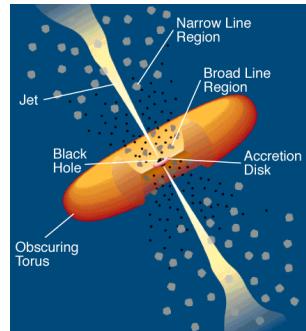
16 new science cases!

AGNs (6)

Polarimetric survey of a big AGN sample:
Agudo et al.

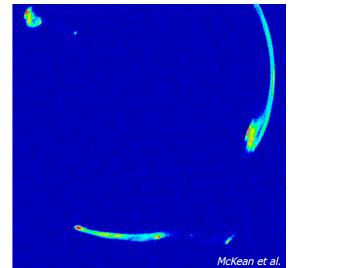
Intermediate mass black holes:
Mezcua et al.

Chasing merged and merging SMBH:
Anton et al.



HI absorption in high-z radio AGN:
Morganti et al.

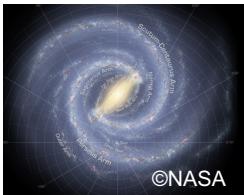
Extremely high-z AGNs: **Perger et al.**



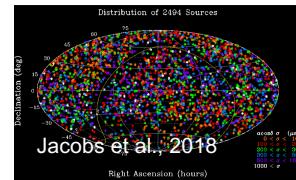
Astrometry (4)

Continuum sources in star forming regions:
Dzib et al.

Ultra-precise astrometry to the MCs: **Rioja et al.**



Gaia counterparts.
Charlot et al.

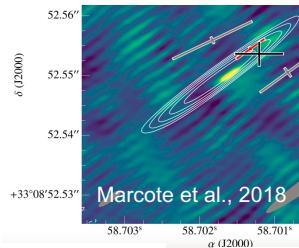


Radio and Gaia tie with radio stars: **Zhang et al.**

Transients (4)

FRBs and their hosts:
Paragi et al.

ULXs in the local Universe:
Middleton et al.

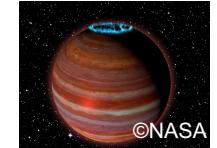


Superflares on low-mass stars:
Villadsen et al.

Inhomogeneous SNe at low freqs:
Chandra et al.

Stars, planets and ISM (2)

Pulsar scintillometry with SKA1-MID:
Kirsten et al.



Radio emission from massive exoplanets:
Gawronski et al.

VLBI with SKA1-LOW:

7 new science cases!

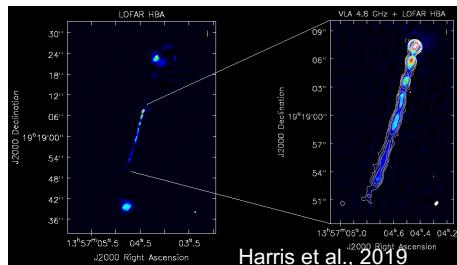


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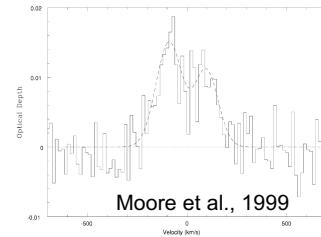


Galaxies and AGN

AGN physics at very low freqs: **Morabito et al.**

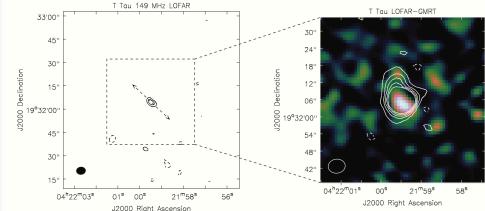


HI absorption at high z:
Gupta et al.



Transients

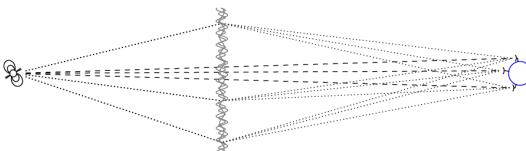
Jets from low mass YSO at very low frequencies: **Ainsworth et al.**



Coughlan et al., 2017

Pulsars and ISM

Pulsar scintillometry at very low freqs: **Kirsten et al.**



Stars, Planets, Astrometry

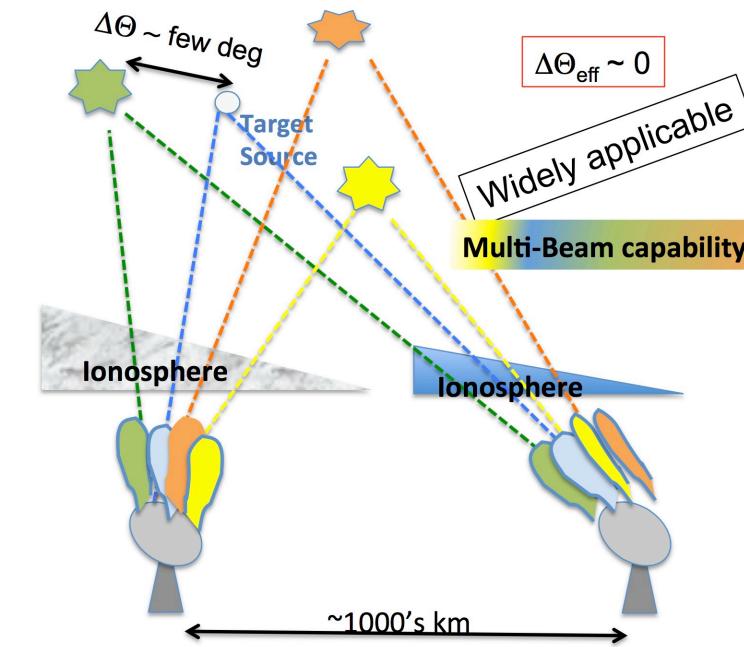
Precise astrometry for exoplanets detection:
Guirado et al.





VLBI with SKA1-LOW and SKA1-MID: 1 new science case

Multi-view astrometry with SKA-VLBI:
Rioja et al.



1 microarcsec relative
astrometric precision!

Rioja et al., 2018





VLBI with SKA: key operational concepts

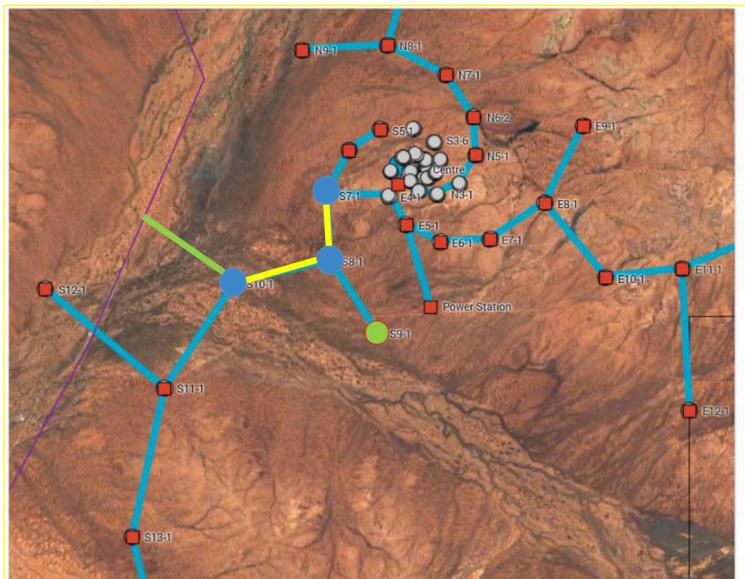
Multiple VLBI beams produced from a subarray of antennas/stations
typically the core



VLBI with SKA: key operational concepts

Multiple VLBI beams produced from a subarray of antennas/stations
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Independent subarrays:
different purposes, up to 16

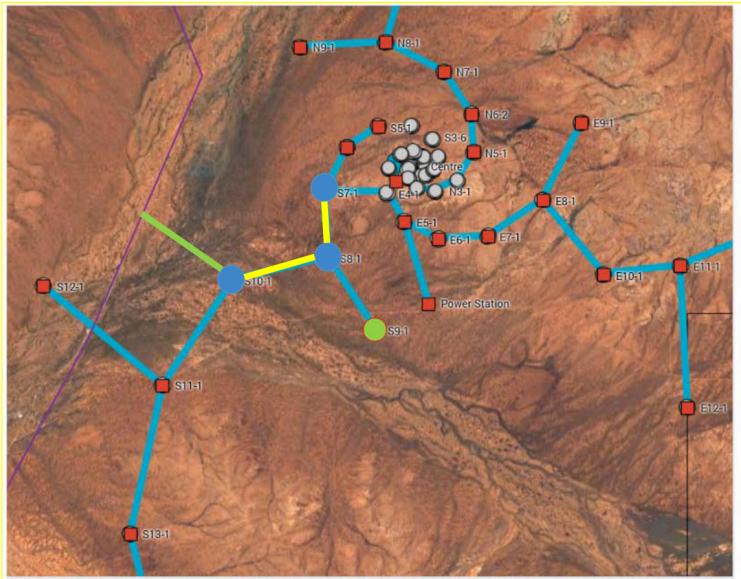




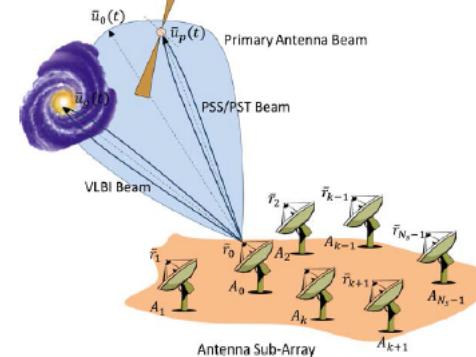
VLBI with SKA: key operational concepts

Multiple VLBI beams produced from a subarray of antennas/stations
typically the core

Independent subarrays:
different purposes, up to 16



Simultaneous/commensal observing modes:
Imaging (continuum, spectral line and fast imaging for slow transients)
Non-Imaging (PSS, PST, transient buffer and VLBI)

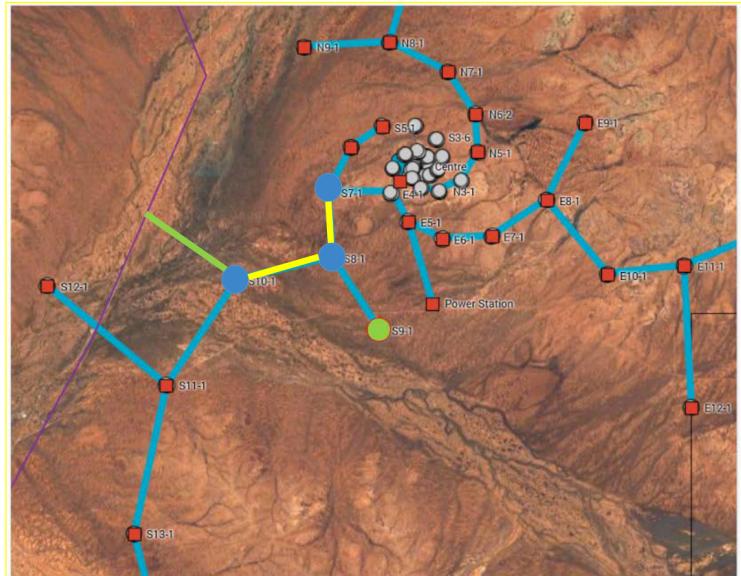




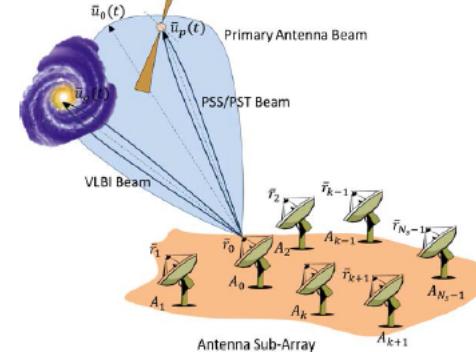
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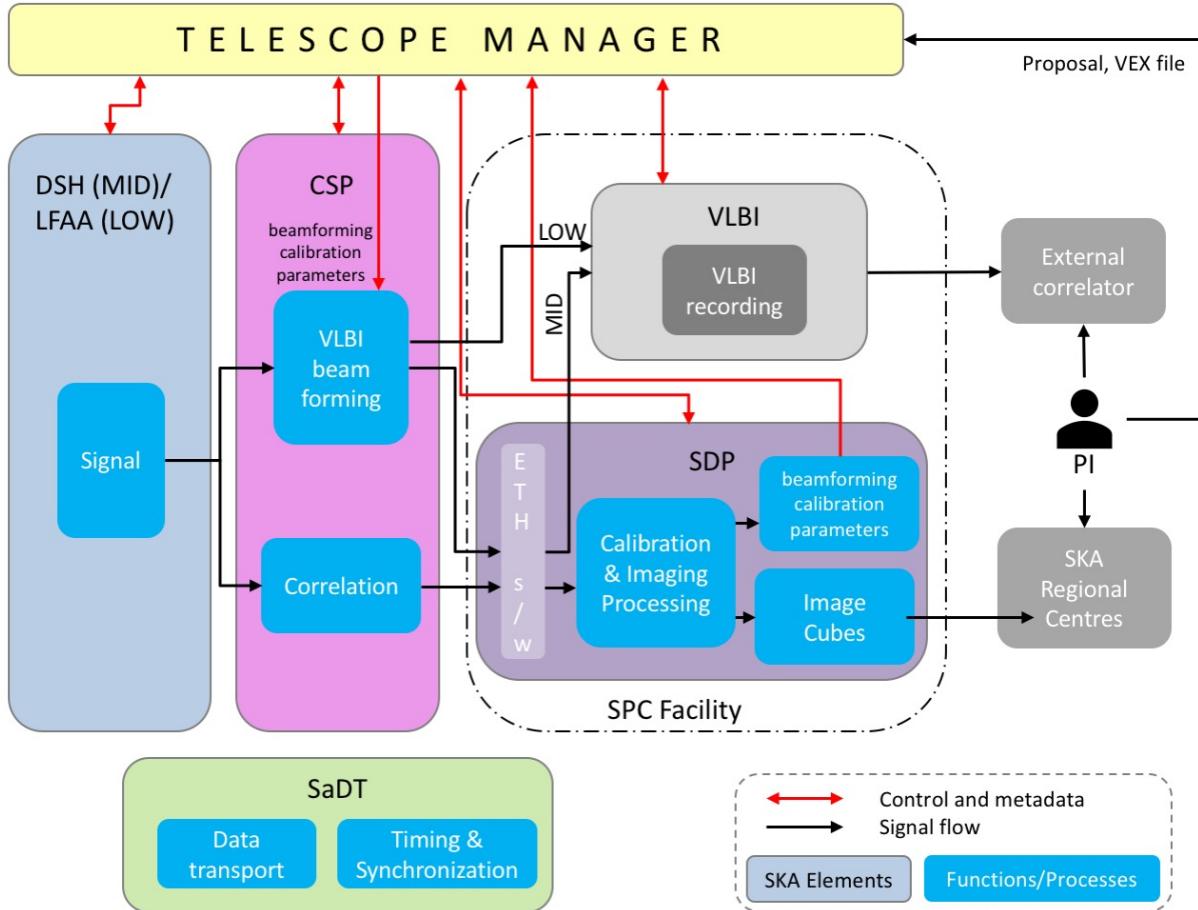


Simultaneous/commensal observing modes:
Imaging (continuum, spectral line and fast imaging for slow transients)
Non-Imaging (PSS, PST, transient buffer and VLBI)



Independent multi-beam capability
within each subarray (in scan boundaries)

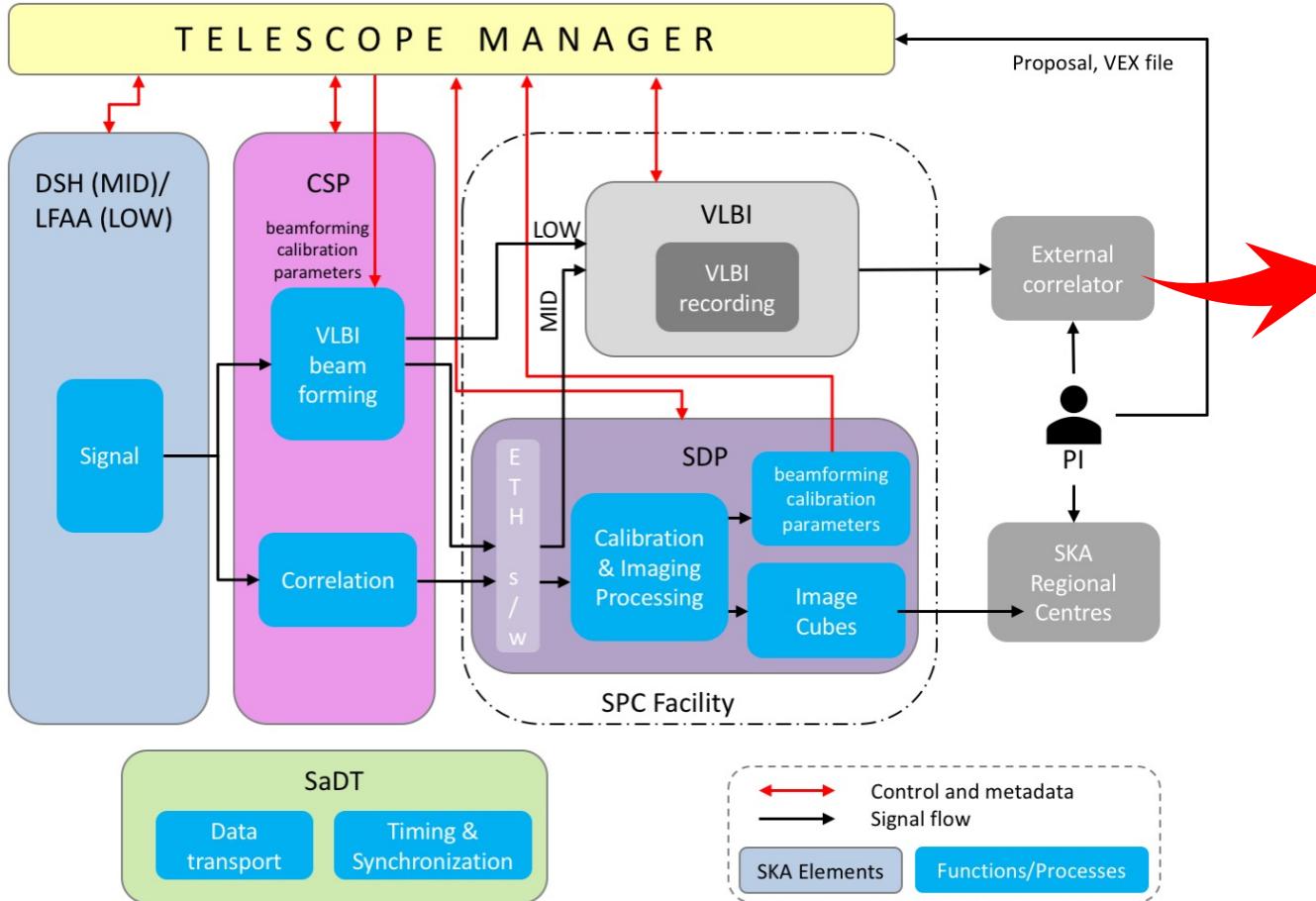
SKA1 VLBI technical implementation



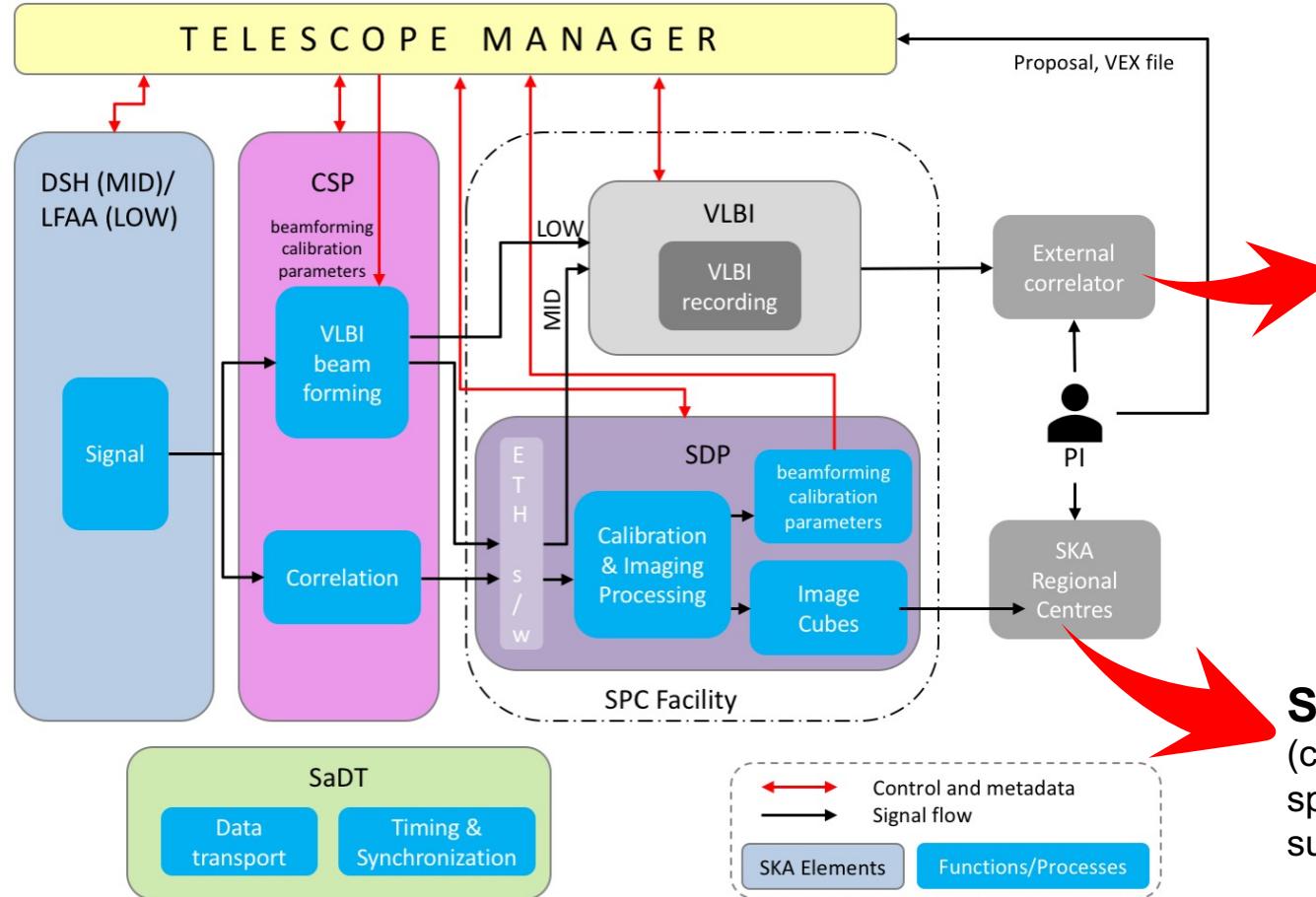
SKA1 VLBI technical implementation



SKA-VLBI
high angular
resolution
images



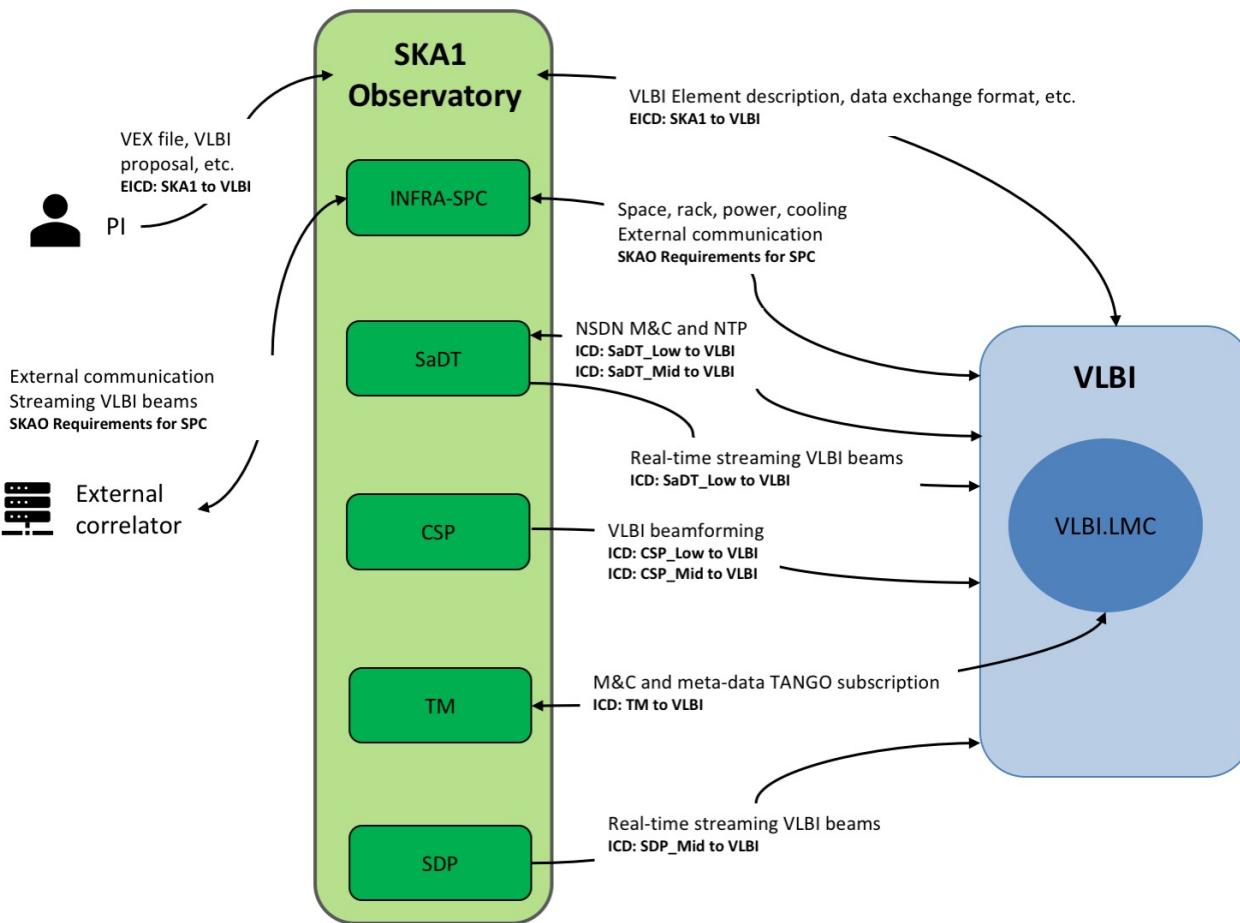
SKA1 VLBI technical implementation



SKA-VLBI
high angular resolution images

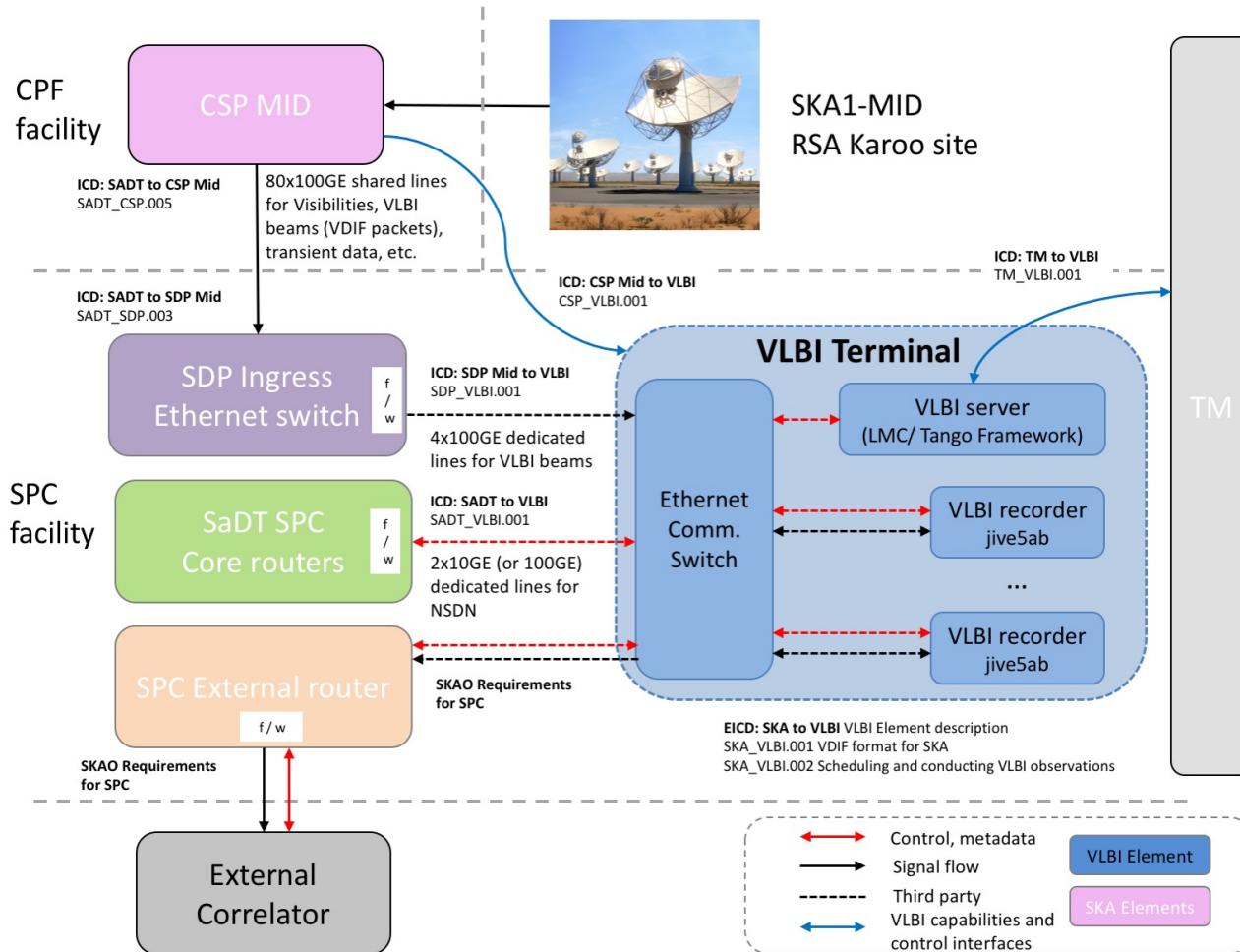
SKA image cubes
(continuum and spectral line, from same subarray)

SKA1 VLBI technical implementation



- ✓ **JUMPING JIVE 1st deliverable:** “Details on VLBI Interfaces to SKA Consortia”
- ✓ **Revision of Level 1 VLBI requirements** (CDR outcomes, CSP assumptions, VLBI SWG inputs).

SKA1 VLBI technical implementation





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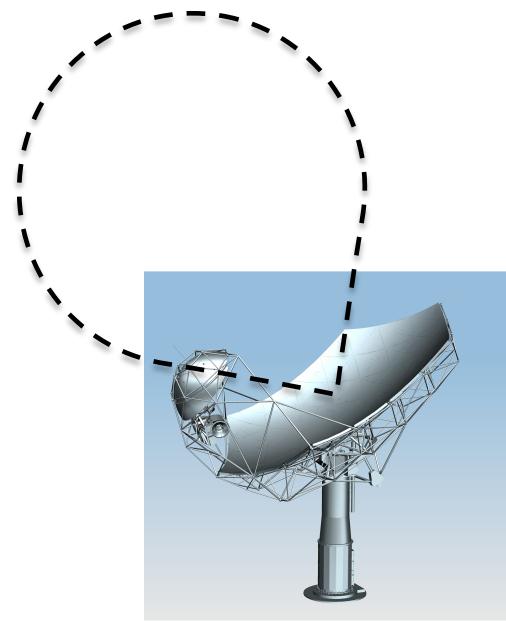
SKA-VLBI with SKA1-MID





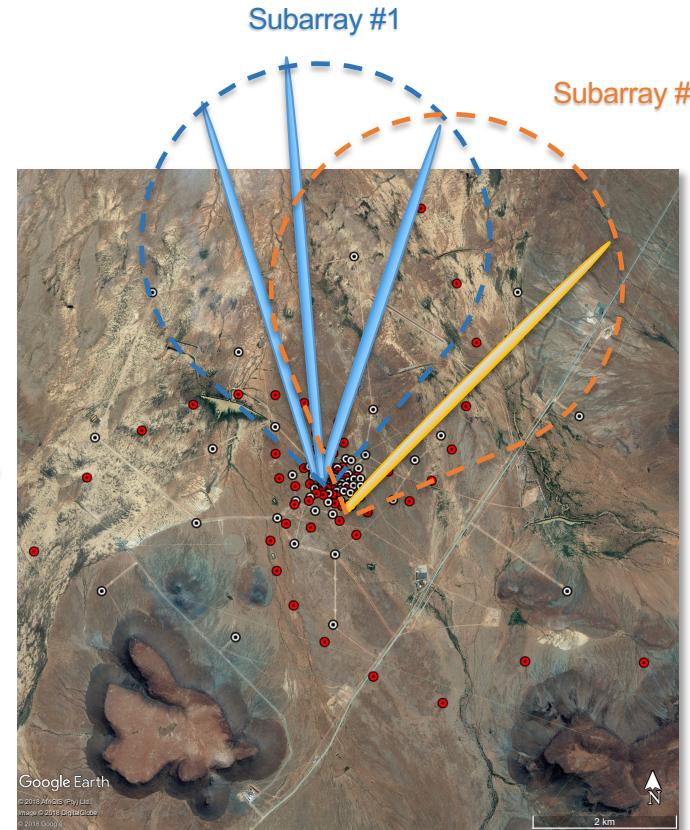
VLBI with SKA1-MID:

All observing modes simultaneously within a subarray
with bandwidth sacrifice



SKA1-MID Antenna

Antenna Beam



SKA1-MID “subarray/array”

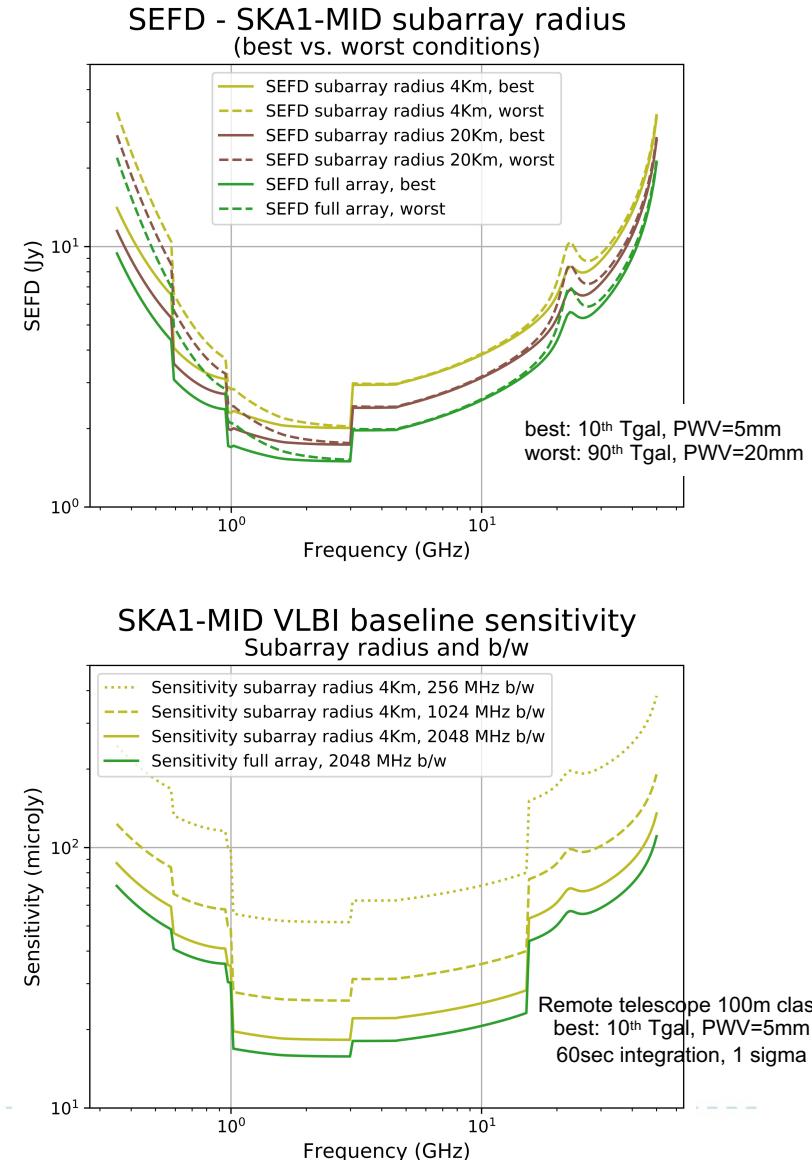
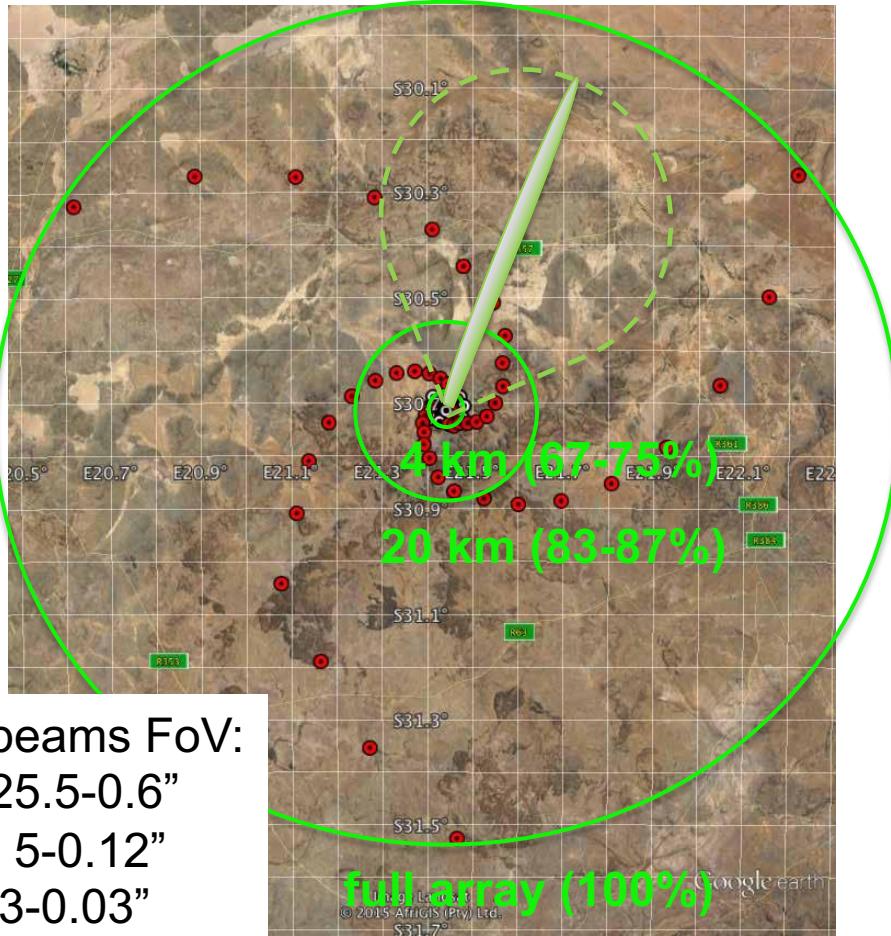
Correlation:

- ✓ **Normal visibilities, zoom**
(100-3 MHz, 6 kHz - 190 Hz)
- ✓ **VLBI coarse visibilities:**
200 kHz

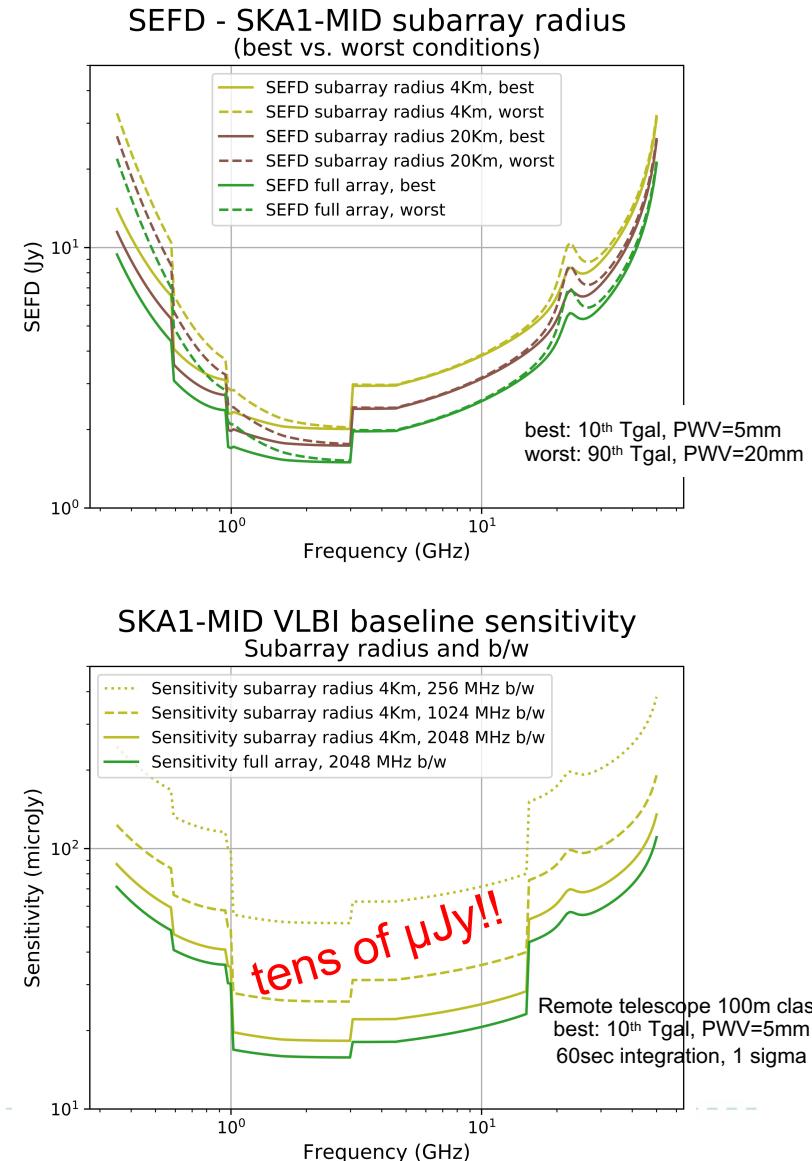
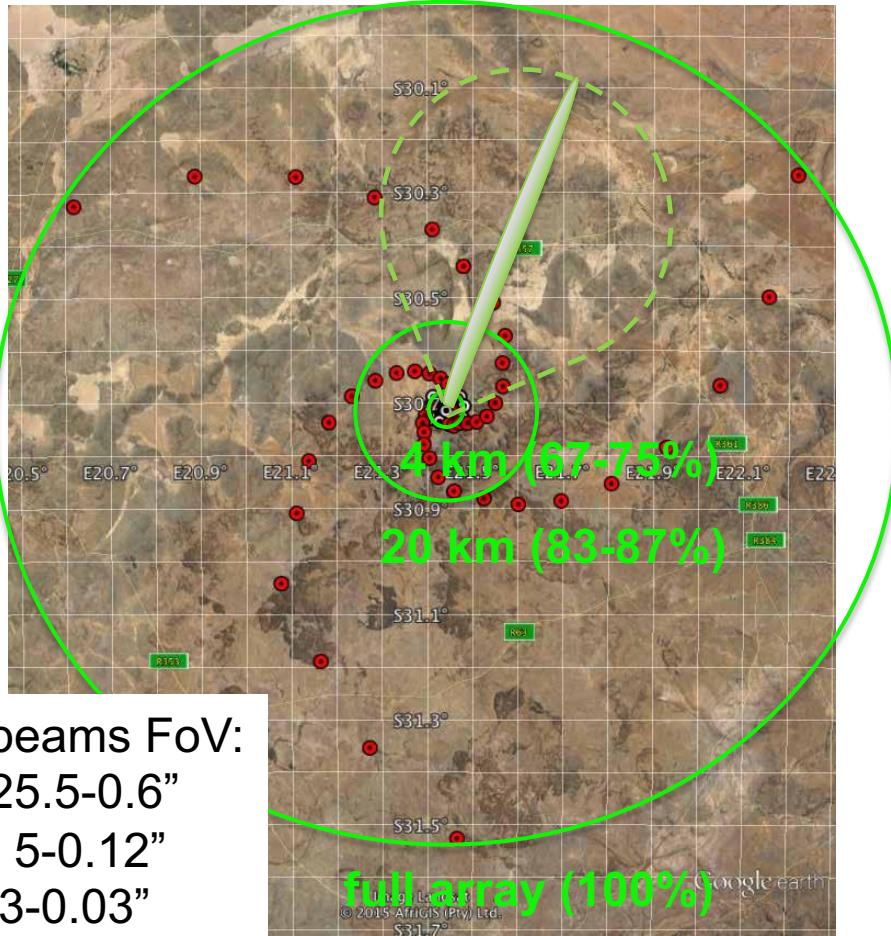
Tied-array beams:

- ✓ **4 VLBI beams** but up to 52 beams max per subarray (200 MHz b/w), from any subarray size
- ✓ **Each VLBI beam:** dual-pol real channels (1-128 & 200 MHz, 2-16 bits, Nyquist)
- ✓ **RFI flagging/excision and polarisation correction**
- ✓ 1500 for **Pulsar Search PSS**
- ✓ 16 for **Pulsar Timing PST**

VLBI with SKA1-MID: configurations



VLBI with SKA1-MID: configurations





Simultaneous Observing with

SKA1-MID: limited by processing resources (26+1 FSP=Frequency Slice Processor)

Band	VLBI + coarse Vis	Imaging	PSS	PST	Zoom
Band 1 (0.35- 1.05GHz)	4beams full (700MHz) (8 FSP)	Full (4 FSP)	1500b 300MHz (8 FSP)	16b full (4 FSP)	2 (2 FSP)
	4b 600MHz (6 FSP)	Full (4 FSP)	1500b 300MHz (8 FSP)	16b full (4 FSP)	4 (4 FSP)
Band 2 (0.95- 1.76GHz)	4beams full (810MHz) (10 FSP)	Full (5 FSP)	1500b 300MHz (8 FSP)	16b 600 MHz (3 FSP)	🚫
	4b 600MHz (6 FSP)	Full (5 FSP)	1500b 300MHz (8 FSP)	16b full (5 FSP)	2 (2 FSP)
Band 5a/b (4.6- 8.5GHz & 8.3- 15.3GHz)	2beams 5GHz (26 FSP)	🚫	🚫	🚫	🚫
	4beams 2.5GHz (26 FSP)	🚫	🚫	🚫	🚫
	4beams 600MHz (6 FSP)	512MHz (3 FSP)	1500b 300MHz (8 FSP)	16b 512 MHz (3 FSP)	6 (6 FSP)





Simultaneous Observing with

SKA1-MID: limited by processing resources (26+1 FSP=Frequency Slice Processor)

Band	VLBI + coarse Vis	Imaging	PSS	PST	Zoom
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Band 5a/b (4.6-8.5GHz & 8.3-15.3GHz)	2beams 5GHz (26 FSP)	🚫	🚫	🚫	🚫
	4beams 2.5GHz (26 FSP)	🚫	🚫	🚫	🚫
	4beams 600MHz (6 FSP)	512MHz (3 FSP)	1500b 300MHz (8 FSP)	16b 512 MHz (3 FSP)	6 (6 FSP)
	🚫	Full (26 FSP)	🚫	🚫	🚫





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SKA-VLBI with SKA1-LOW

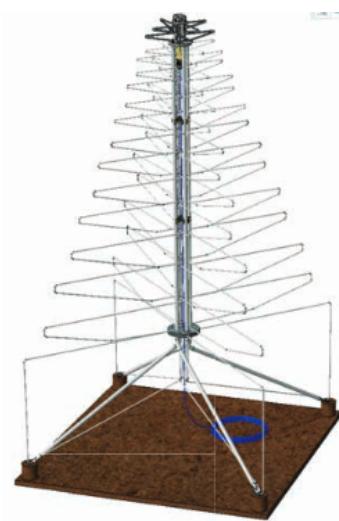


Exploring the Universe with the world's largest radio telescope

VLBI with SKA1-LOW: 50-350 MHz



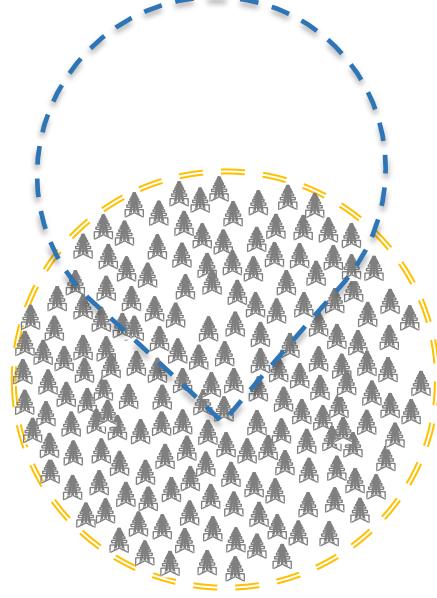
All observing modes simultaneously:
for each subarray, for each station beam



SKA1-LOW
Antenna/Receptor

Antenna Beam

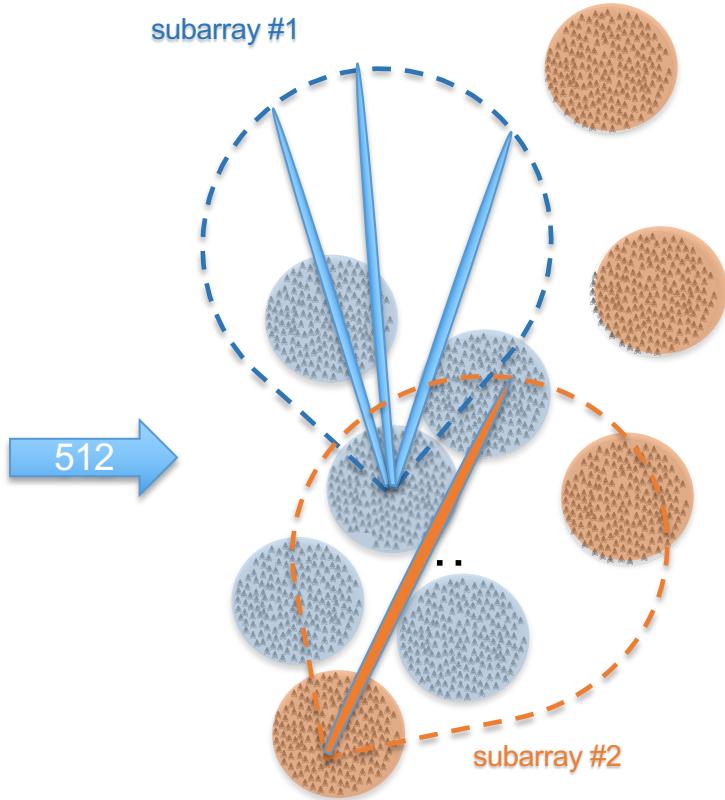
256



SKA1-LOW
“Station”

Station Beam
300 MHz b/w

512



SKA1-LOW “subarray”

Correlation: normal and zoom
(14.4 kHz - 5.4kHz - 226 Hz)

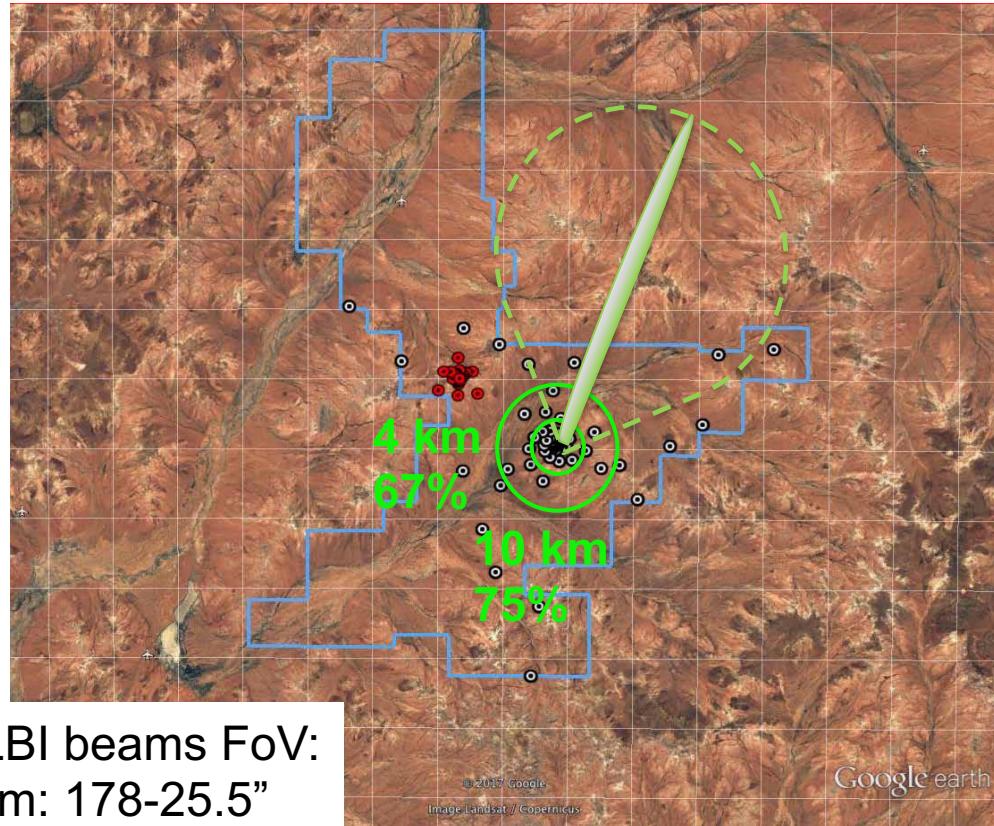
Tied-array beams from 20 Km diameter subarray:

- ✓ **4 VLBI beams (in total)**
- ✓ **500 for Pulsar Search PSS**
- ✓ **16 for Pulsar Timing PST**

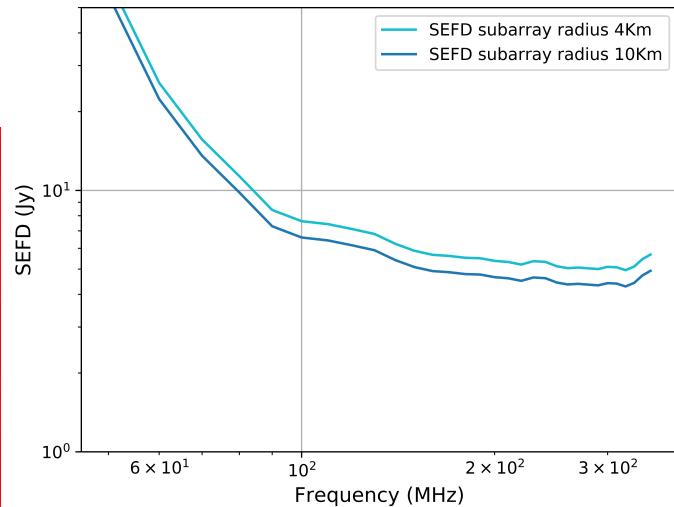




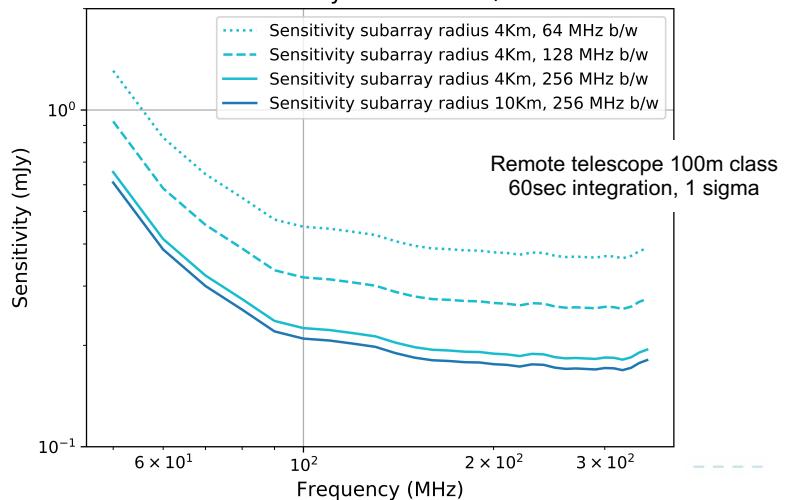
VLBI with SKA1-LOW: configurations



SEFD - SKA1-LOW subarray radius

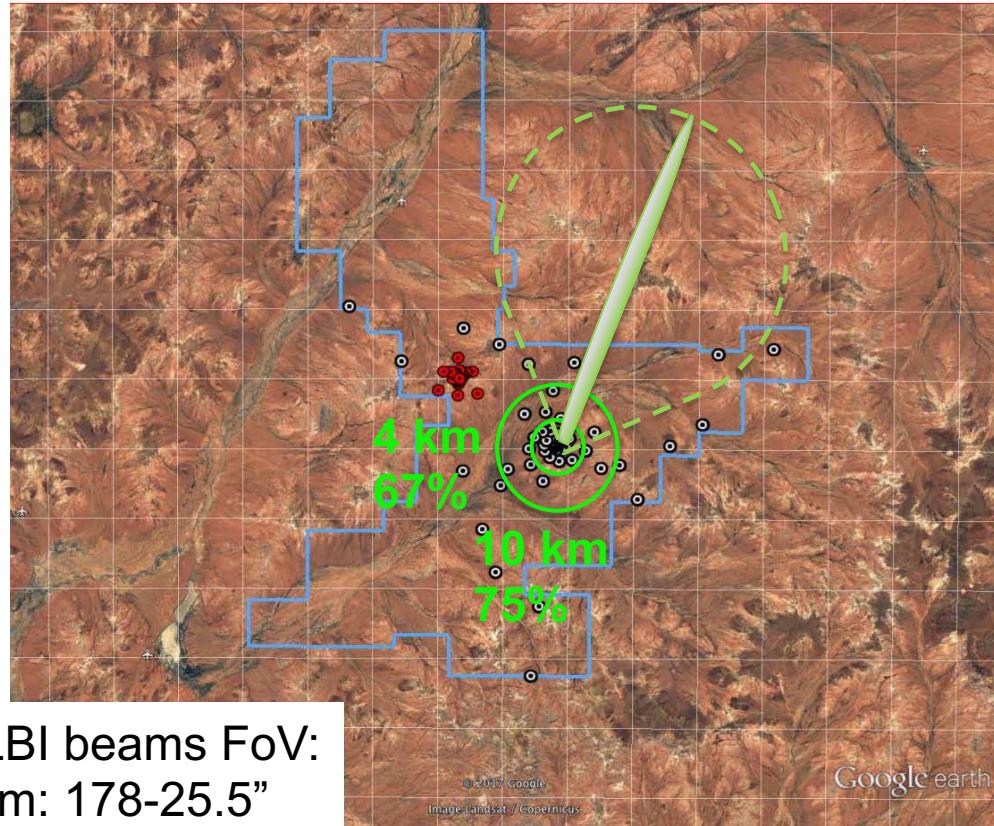


SKA1-LOW VLBI baseline sensitivity
Subarray radius and b/w



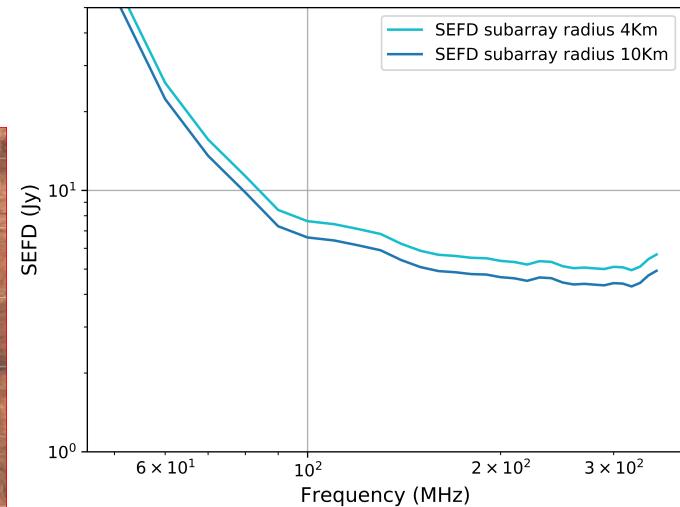


VLBI with SKA1-LOW: configurations

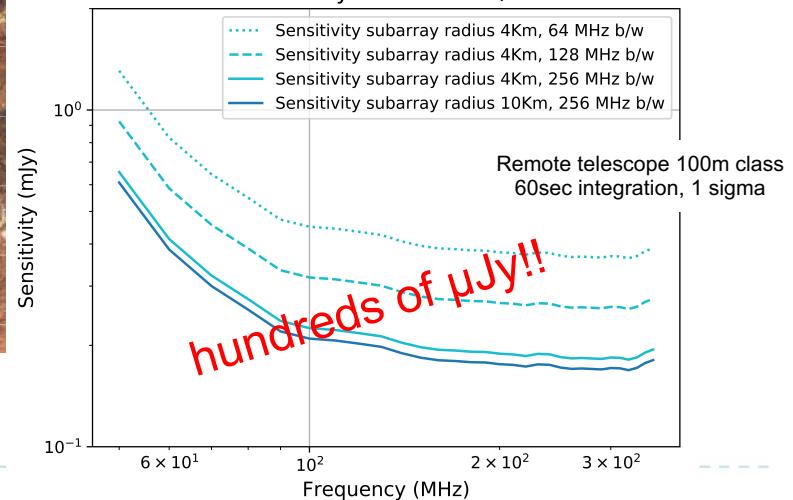


VLBI beams FoV:
4km: 178-25.5"
10km: 71-10"

SEFD - SKA1-LOW subarray radius



SKA1-LOW VLBI baseline sensitivity
Subarray radius and b/w





SKA-VLBI WORKSHOP
THE WORLD'S EYE ON THE SKY

14 – 17 OCTOBER 2019

SKA GLOBAL HQ, UK

SKATESCOPE.ORG / SKA-VLBI-WORKSHOP

INVITED SPEAKERS

DANA SIMARD (U. Toronto, CA): Pulsar scattering
JACK RADCLIFFE (U. of Pretoria/SARAO): Wide-field VLBI
MARCELLO GIROLETTI (INAF, IT): GWEM counterparts VLBI follow-up
JAN FORBRICH (U. Hertfordshire, UK): Stellar continuum, young stellar objects
YOON KYUNG CHOI (MPIfR-Bonn, DE): Measuring astrometry, evolved stars
MANISHA CALEB (U. Manchester, UK): Fast radio bursts
PIKKY ATRI (ICRAR, AU): Black hole X-ray binaries
LEONARD COOPER (Oxford, UK): Low-frequency AGN surveys
JOHN MCKEEAN (ASTRON, RU Groningen, NL): Gravitational lensing, cosmology
JAMES CHIBUEZE (North West U, SA): VLBI in Africa

LOCAL ORGANISING COMMITTEE

SARAH LAMB (SKAO) | JOSEPH DIAMOND (SKAO) | CLAIRE TAYLOR (SKAO) | ROBERT BERWICK (JBOAU, Manchester) | ANTONIO CHRYSOSTOMOU (SKAO) | CRISTINA GARCIA-MIRO (SKAO)

SCIENTIFIC ORGANISING COMMITTEE

LAURA BRITLER, MPIfR-Bonn, Germany | KAZU RYGL, INAF-Roma, Italy | MICHAEL RUPEN, HAAPO, Canada
MARCELLO GIROLETTI, INAF, Italy | CLAUDIO PHILLIPS, School of Physics, University of the Witwatersrand, South Africa
MARI MEZZA, CIE-CSIC, Spain | PREETI KHARB, NCRA-TIFR, India | HIROSHI IMAI, U. Kagoshima, Japan
ROGER DEANE, U. Pretoria, South Africa | JOHN CONWAY, OSGChalmers U, Sweden
PACO COLOMER, JIVE, Netherlands | ANNA BONALDI, SKAO, United Kingdom | TAD AN, SNAO, China
ANTONIO CHRYSOSTOMOU, SKAO, United Kingdom (co-Chair) | ZOEY PRASER, JIVE, Netherlands (co-Chair)

SKA
SQUARE KILOMETRE ARRAY

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RadioNet

- ✓ **65 scientist from 18 countries**
- ✓ **Workshops sessions:** AGN, transients, pulsars and FRBs, high precision stellar astrometry and prospects for SKA-VLBI including African telescopes
- ✓ **SKA-VLBI data challenges**
- ✓ Four working groups to discuss **SKA-VLBI Key Science Projects**

“Women in action” from the astrometry working group





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SKA-VLBI Conclusions



SKA-VLBI Conclusions

- ✓ **VLBI is an observing mode** of the SKA Observatory
- ✓ CSP design compatible with **VLBI standards & networks**
- ✓ SKA-VLBI System level solution ready, it can be applied to **SKA precursors**
- ✓ The SKA-VLBI community supports enthusiastically
- ✓ Let's start cutting-edge VLBI science with SKA precursors:
MeerKAT-VLBI, (MK+)-VLBI, ASKAP-VLBI...

SKA-VLBI capacity and technique



SQUARE KILOMETRE ARRAY

Exploring the Universe with the world's largest radio telescope

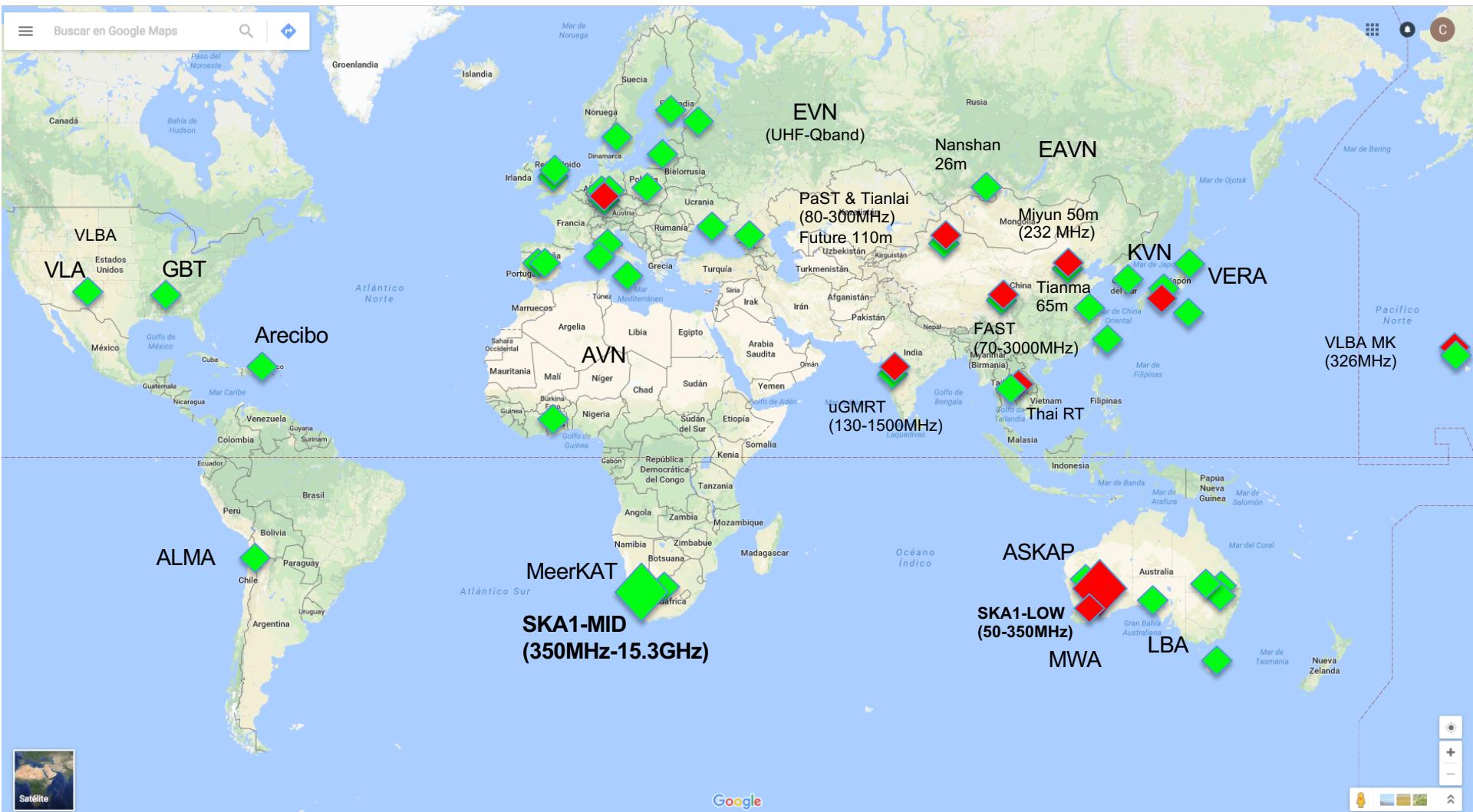
Cristina Garcia Miro / Tao An

JUMPING JIVE Project / SKA VLBI SW

miro@jive.eu JIVE / SHAO



VLBI with SKA: Global VLBI network





SQUARE KILOMETRE ARRAY

Exploring the Universe with the world's largest radio telescope



SKA1-MID

SKA1-LOW

uGMRT
(130MHz-1500MHz)

Table 2: Common visible time in hours for the SKA sites with other array elements

South African SKA									
Dec	W. Aus.	JP	CN	NZ	CL	USA-HI	USA-NM	USA-PR	Europe
+45°	0.0	2.2	5.4	0.0	0.0	0.0	1.5	3.6	5.4
+30°	0.0	2.7	7.6	0.0	0.4	0.0	1.8	4.4	8.0
+15°	0.0	2.7	7.3	0.0	2.3	0.0	1.8	4.8	9.4
0°	0.8	2.5	6.8	0.6	3.7	0.0	1.7	5.0	9.1
-15°	2.0	2.1	6.3	1.9	4.9	0.0	1.4	5.1	7.8
-30°	3.0	1.2	5.4	3.1	6.2	0.0	0.0	5.2	3.0
-45°	4.0	0.0	0.0	5.2	7.9	0.0	0.0	5.2	0.0
-60°	5.5	0.0	0.0	11.8	11.6	0.0	0.0	0.0	0.0
-75°	12.0	0.0	0.0	24.0	24.0	0.0	0.0	0.0	0.0
-90°	24.0	0.0	0.0	24.0	24.0	0.0	0.0	0.0	0.0

Western Australian SKA									
Dec	ZA	JP	CN	NZ	CL	USA-HI	USA-NM	USA-PR	Europe
+45°	0.0	5.3	5.3	0.0	0.0	4.0	1.3	0.0	5.3
+30°	2.0	7.9	7.9	3.2	0.0	4.8	1.7	0.0	4.5
+15°	3.5	8.6	9.2	5.2	0.0	5.1	1.7	0.0	3.8
0°	4.5	8.5	6.7	6.6	0.0	5.3	1.6	0.0	3.2
-15°	5.5	7.8	8.3	7.8	0.4	5.4	1.3	0.0	2.3
-30°	9.0	5.1	6.2	9.1	1.9	5.5	0.8	0.0	0.4
-45°	7.7	0.0	0.0	10.8	5.2	5.4	0.0	0.0	0.0
-60°	11.5	0.0	0.0	14.1	11.7	5.1	0.0	0.0	0.0
-75°	24.0	0.0	0.0	24.0	24.0	0.0	0.0	0.0	0.0
-90°	24.0	0.0	0.0	24.0	24.0	0.0	0.0	0.0	0.0

Key: W. Aus.: Western Australia; ZA: South Africa; JP: Japan, CN: Eastern China; NZ: New Zealand; CL: Chile (latitude of former TIGO site); USA-HI: Hawaii; USA-NM: New Mexico; USA-PR: Puerto Rico; Europe: Id.

uGMRT
5.0
7.0
7.5
8.0
8.5
6.0
6.0
0.0
0.0
0.0



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