

Post-reionisation 21cm Cosmology Sky and Instrument Simulations

Devin Crichton and ETHZ Radio Cosmology Group[†]
SKACH Consortium Meeting @ ETHZ
24.05.22

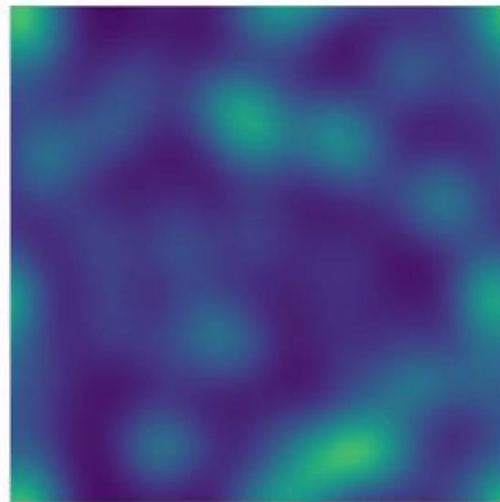
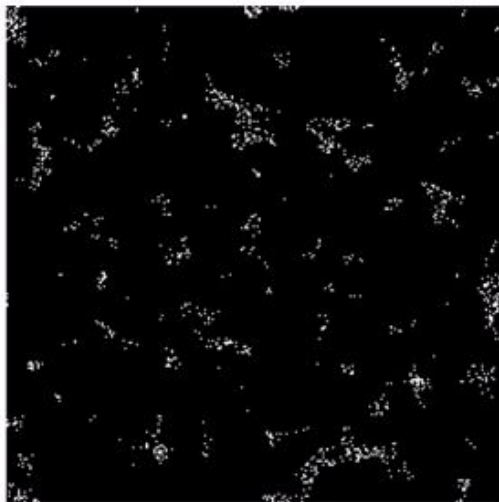
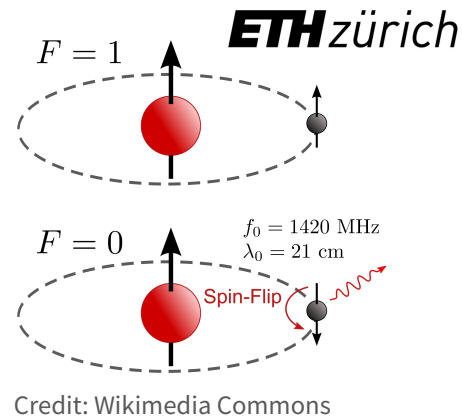


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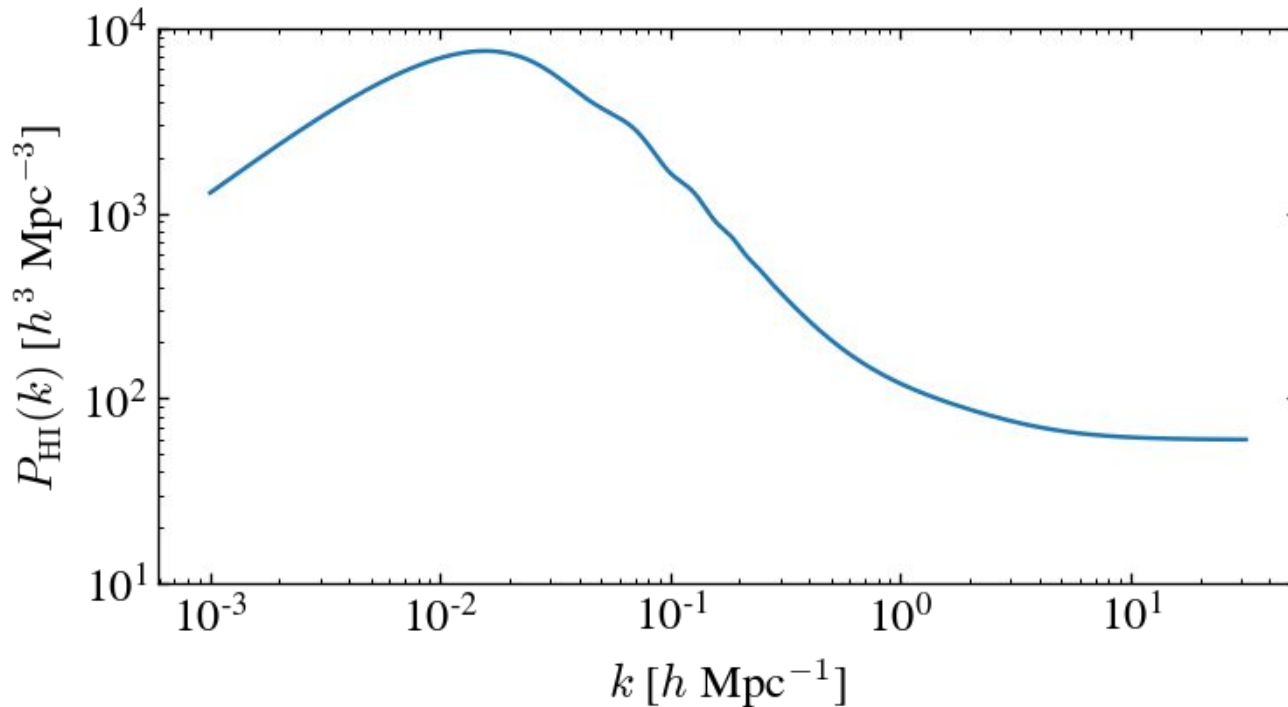
[†]Alexandre Refregier, Marta Spinelli, Sara A. Safari, Pascal Hitz, Pascale Berner

HI Intensity Mapping

- Hyperfine Hydrogen transition line at 1420.4 MHz
- Efficiently and tomographically map cosmological volumes
 - Generally low angular resolution but redshift information cheap
 - Probe epoch of reionisation at low frequencies and large scale structure at high frequencies.
- Post-reionisation IM
 - $\nu > 200\text{-}300\text{MHz}$
 - Biased tracer of large scale structure
 - Cosmological constraints from HI power spectrum
 - Large volumes achievable

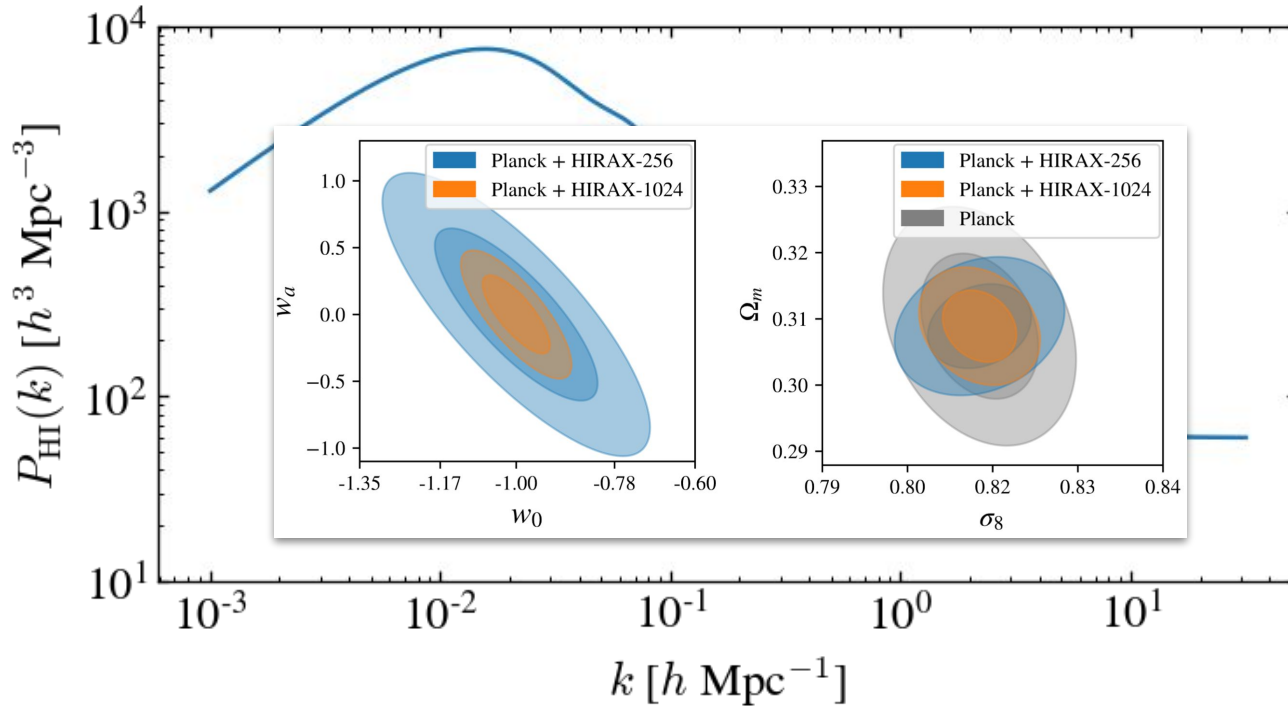


HI Power Spectrum



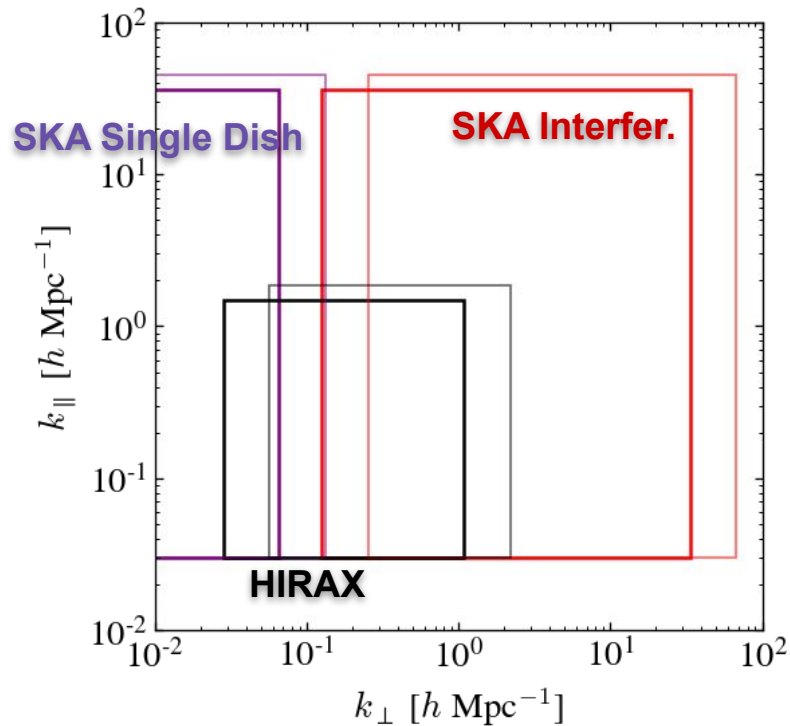
- Primordial non-gaussianity
- Modified gravity theories
- Growth of structure
- Geometric Constraints
- Expansion rate
- Dark energy
- Non-linear dynamics
- HI content of galaxies

HI Power Spectrum



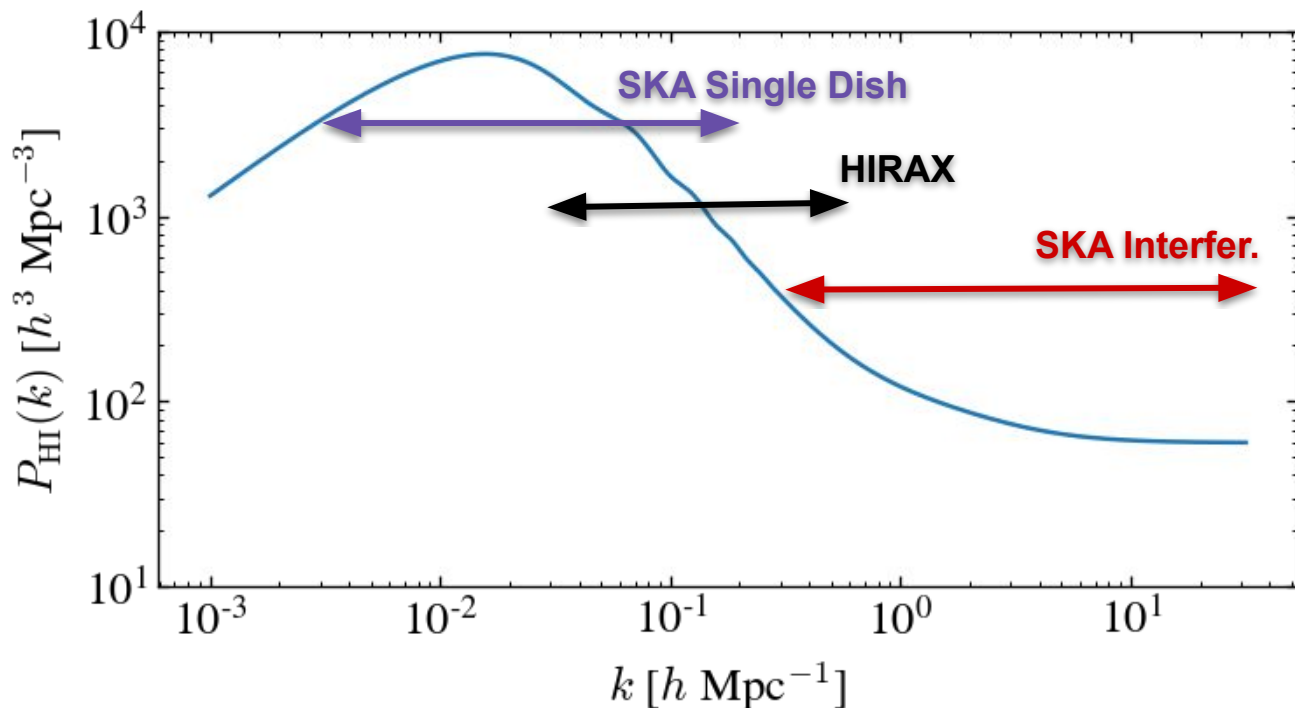
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Instrument Diversity



- Power spectrum sensitivity:
 - k_{\perp} : Angular scales (baselines/beam)
 - k_{\parallel} : $k_{\text{FG}} \rightarrow$ Channel width / k_{NL}

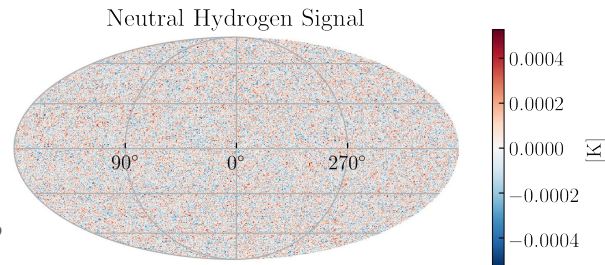
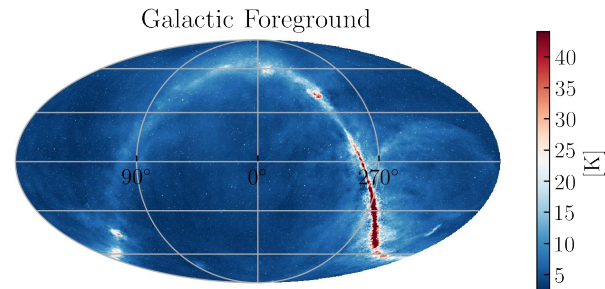
HI Power Spectrum



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Systematics / Chromaticity and Foregrounds

- Foregrounds are the primary challenge for 21cm cosmology
 - Galactic signal brighter by many orders of magnitude
- Signal and Foregrounds have different, *on-sky* properties
 - Galactic emission is:
 - Polarised
 - Strongly correlated over wide frequency bands
 - Structured on the sky in ~known way
 - In principle, there are not many mixed *on-sky* degrees of freedom
- Mode-mixing inherent in measurement is a major issue
 - Instrument has chromatic response fundamentally as well as arising from systematics
 - With perfect knowledge of the instrument, this can be accounted for, however the large contrast in signal strengths can make small reconstruction residuals a big problem

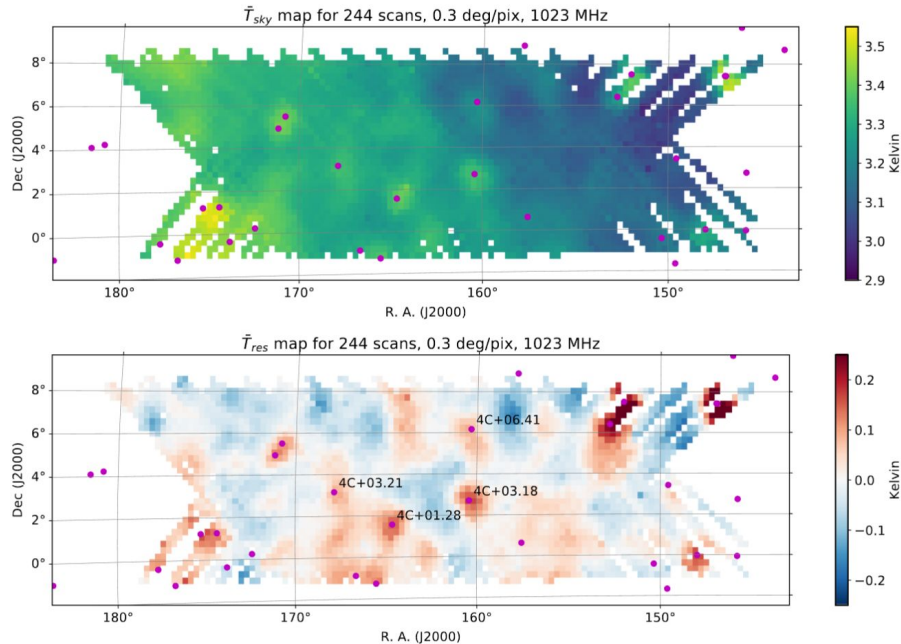


Single Dish Observations with MeerKAT

Cosmology with MeerKAT is underway

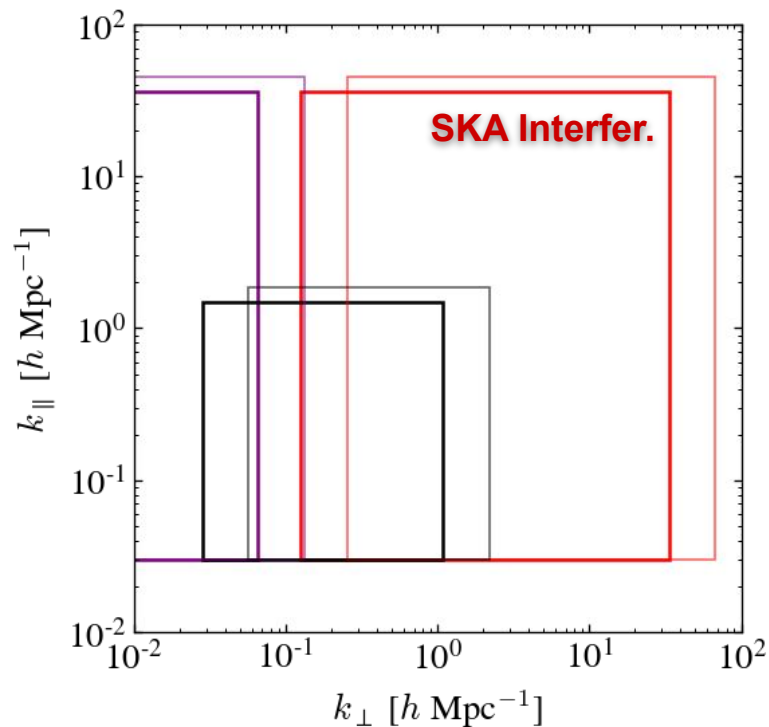
MeerKLASS pilot survey:

- 11 hours, $\sim 200 \text{ deg}^2$
- Demonstrated analysis and foreground mitigation techniques work on real data
- Synchrotron model constraints
- **Cosmological signal detected in cross-correlation with WiggleZ**



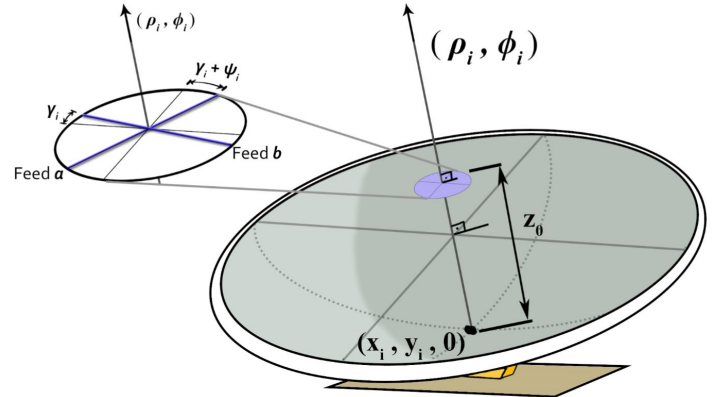
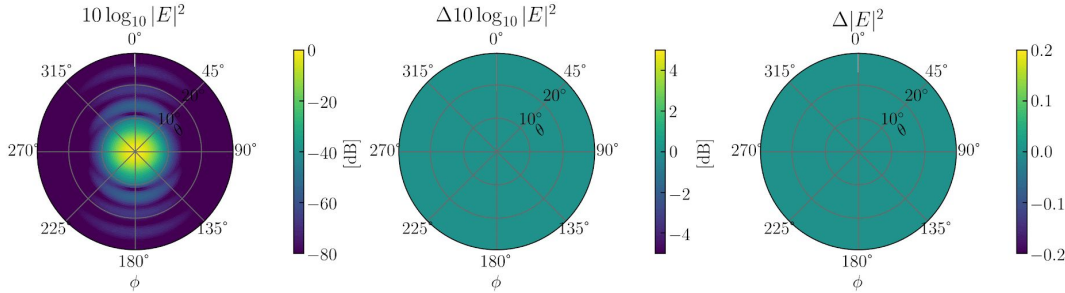
SKA Interferometric Intensity Mapping

- Work package at ETHZ for Digital Twin project (Sara A. Safari @ ETHZ)
 - Develop flexible, scalable power spectrum extraction pipeline
 - Robust to spectral gaps
 - Include primary beam models and foreground mitigation
- Mosaicing / survey strategy optimisation
 - Commensural opportunities



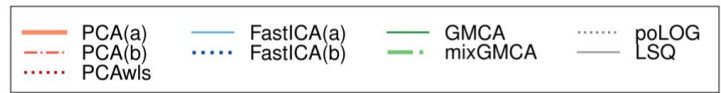
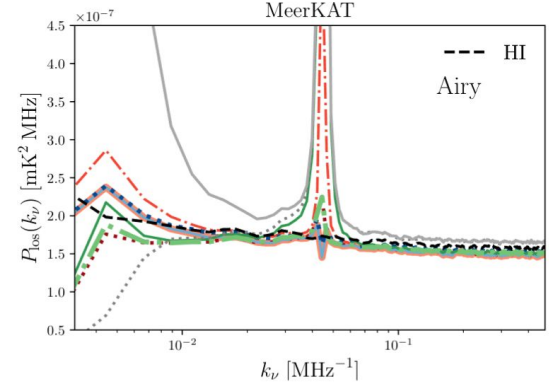
Instrumental Systematics

Focal Sweep x @ 600 MHz; $dx = 0.0$ cm



Credit: Dalian Sunder

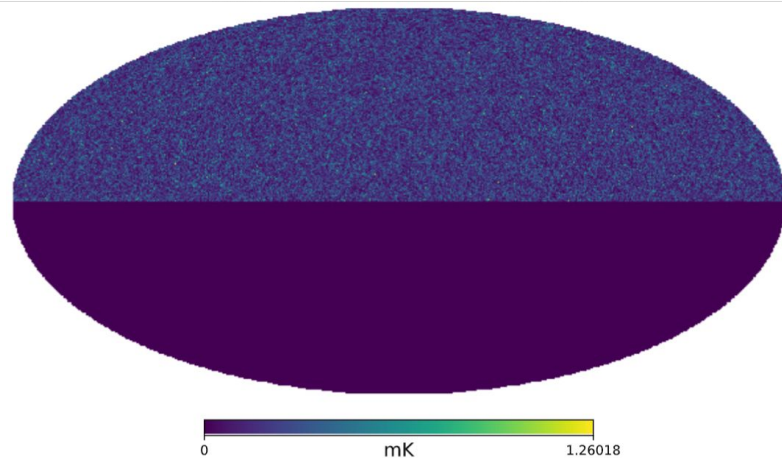
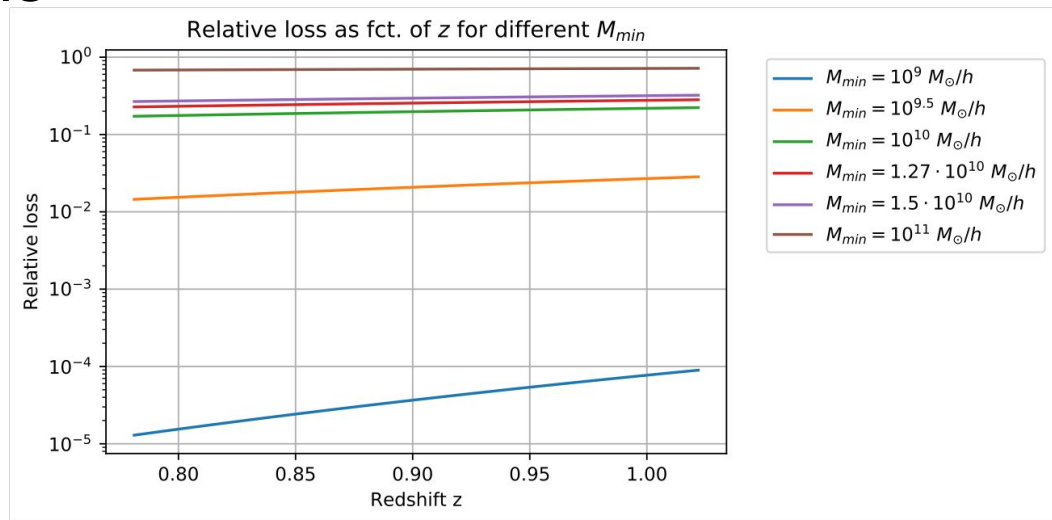
- Instrument modeling is critical
 - Chromatic response needs to be characterised
 - Signal loss needs to be minimal and understood
 - Incorporate instrument simulation techniques
- Beam measurement informative
 - External constraints on chromaticity
 - Holography, drone etc.



Credit: Marta Spinelli et al. 2021 | arXiv: 2107.10814

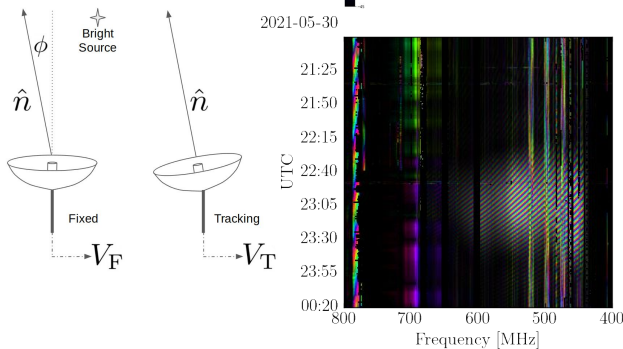
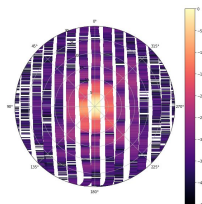
Pinocchio Sky Simulations

- Developing 21cm sky
- PINOCCHIO LPT haloes + HI mass function → Simulated sky maps
 - Use empirical halo model from Padmanabhan et al. 2017
- Fast and robust alternative to n-body simulations suitable for forward modeling and x-corr.
- Aim significant scale up simulations and provide interface in digital twin
 - Lower mass resolution
 - Range of cosmologies and mass function parameters



Other ETH work

- HIRAX Correlator and Science Data Processor
- Drone beam mapping
- Holographic Beam mapping



Christian Monstein, Thierry Viant, Tony Walters



AMD EPYC 7452 CPU	AMD EPYC 7452 CPU	1 TB RAM
NVIDIA A40 GPU		
2x40Gbe NIC		
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2x40Gbe NIC		
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Thierry Viant, Andre Renard, Keith Vanderlinde and others



Thanks!